#### Relativistic Magnetized Jets without Current Sheets

K.N. Gourgouliatos (Purdue)

Collaborators: Ch. Fendt (Heidelberg), J. Braithwaite (Bonn), E. Clausen-Brown (Purdue) & M. Lyutikov (Purdue)

#### Outline

Motivation

Mathematical Setup
Ctatic Calution

Static Solution

Relativistic Solutions

Simulation

Stability

Observational Signature

 We describe the structure of jets containing both plasma and magnetic field

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- The external timescales are slow enough, compared to the internal ones, so that the system can reach an equilibrium state
- If the reconnection takes place fast enough we expect that any current sheets will be dissipated

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  - Endpoints of dissipative evolution
  - Simulation trial solutions initial conditions

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- We assume cylindrical symmetry
- Some hot plasma confines the cylinder
- What is the equilibrium solution for this structure?

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The boundary condition between regions (I) and (II) is that the magnetic field should drop to zero and there are no current sheets



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 $\boldsymbol{B} = \nabla P_p(R) \times \nabla \phi + I_p(R) \nabla \phi$ 

 $\boldsymbol{B} = \nabla P_t(R) \times \nabla z + I_t(R) \nabla z$ 

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For the Force-Free case, for a linear relation between *P* and *I*, there is the Lundquist (1951) solution For the Force-Free case, for a linear relation between *P* and *I*, there is the Lundquist (1951) solution


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In the G-S case there are two classes of solutions, depending on the choice of the pressure: proportional to the poloidal or toroidal flux

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0.8

0.6

0.4

0.2

0 -

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We can construct two basic families of solutions, where both the toroidal and the poloidal field are zero at the boundary of the cylinder.

To ensure the absence of current sheets we need a minimum pressure





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$$B_t = c_t \alpha_t J_1(\alpha_t R)\hat{\phi} + (c_t \alpha_t J_0(t))$$

$$D_{000} = 0$$

$$D_{000} =$$

nt

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Motion along the z axis induces electric field

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 $H^2 = B_\phi^2 - E_R^2$ 

Using the same means as in the static case we can solve the relativistic

$$E_R = \frac{v_z}{(1 - v_z^2)^{1/2}} \left( c_p \alpha_p J_1(\alpha_p R) - \frac{F_p R}{\alpha_p} \right)$$
$$B_\phi = \frac{1}{(1 - v_z^2)^{1/2}} \left( c_p \alpha_p J_1(\alpha_p R) - \frac{F_p R}{\alpha_p} \right)$$
$$B_z = c_p \alpha_p J_0(\alpha_p R) - \frac{2F_p}{\alpha_p^2} ,$$
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$$V_F = \frac{E \times B}{B^2} = \frac{(-v \times B) \times B}{B^2} = v - \frac{v \cdot B}{B^2}B$$

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The  $V_F$  velocity depends on the details of the magnetic field structure



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### Polarization

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- They answer the question of the pressure lying between the pure hydro models and the magnetic models
- In the observational side they do not have significant differences from the Force-Free structures used

# Thanks