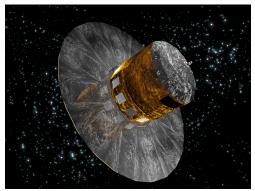
# Towards an Automated Processing of Gaia Eclipsing Binaries

#### **Christos Siopis**

Institut d'Astronomie et d'Astrophysique Université Libre de Bruxelles Belgium

10th Hel.A.S. Conference, Ioannina, Greece





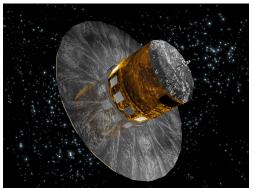
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 Goal: To provide a 6D map of the Galaxy

#### Gaia is a scanning mission

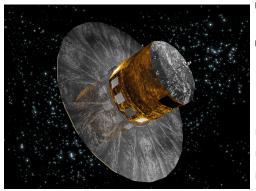
- no pointing
- no change in schedule
- 30-200 transits per object (~80 transits on average)
- $\triangleright$  ~ 10<sup>9</sup> stars to V = 20 22
- $\blacktriangleright \sim 10^6 10^7$  galaxies
- $\blacktriangleright$  ~ 500,000 quasars
- ~ 3 × 10<sup>5</sup> solar system objects
  ~ 10<sup>4</sup> exoplanets

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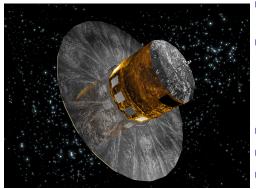


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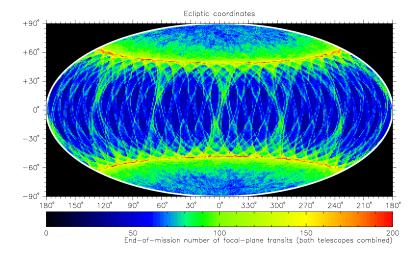


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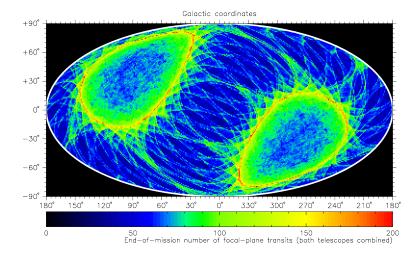
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#### Gaia sky coverage – Ecliptic coordinates



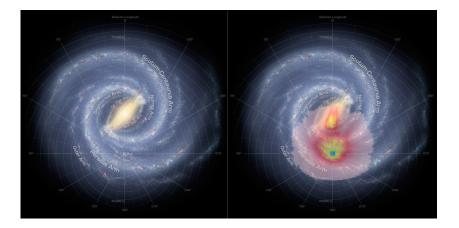
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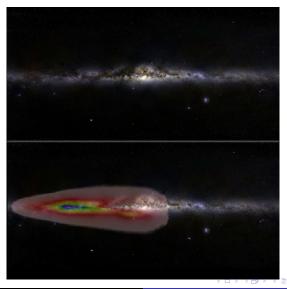
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### Gaia Catalog Coverage: View from Galactic Pole



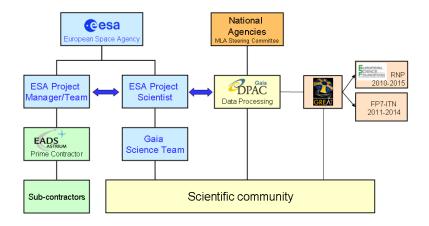
The Gaia Mission The Gaia Organization Diagram

### Gaia Catalog Coverage: View from Galactic Equator

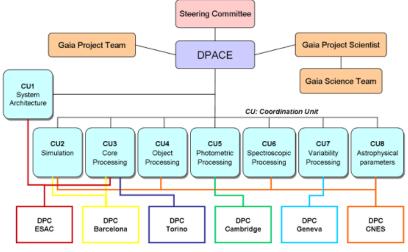


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# The Gaia Organization Diagram



# DPAC: Data Processing & Analysis Consortium



DPC: Data Processing Centre

- GREAT ITN (Initial Training Network) programme
- ▶ GREAT ESF (European Science Foundation) programme
  - A pan-European research infrastructure to facilitate the fullest exploitation of the Gaia Catalog
  - Funding for workshops, conferences, training events, exchange visits, ...
  - Some key topic areas:
    - Origin, structure, and evolution of the Milky Way.
    - Galactic Dynamics & Archaeology
    - Stellar Astrophysics
    - Binary Stars
    - Extrasolar planets
    - The Solar System

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## Doing Eclipsing Binaries with Gaia

- Responsibility of CU4 Development Unit (DU) 436
  - Christos Siopis (DU436 leader)
  - Brandon Tingley (now at IAC)
  - Gilles Sadowski (physicist, computer scientist)
  - Associate members

#### ▶ Why bother with eclipsing binaries?

- Intrinsically interesting!
- Provide only general way to determine stellar masses
- Distance indicators

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#### Unique aspects of Gaia EB processing:

- Need for automated processing
- ▶ DPAC enforces deadlines, programming guidelines, ...
  - Software must be implemented in Java!
  - Performance requirements
- ▶ 5 (6?)-year baseline
  - Constrains EB periods
  - Slightly simplifies modeling

#### ▶ Important points of light curve not necessarily sampled

- A consequence of the Gaia scanning law
- Photometry: 30-200 transits per object (~80 mean)
- Spectroscopy: 20-120 transits per object (~40 mean)
- Non-uniform (albeit quisi-regular) sampling

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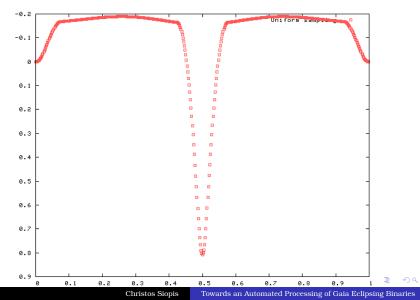
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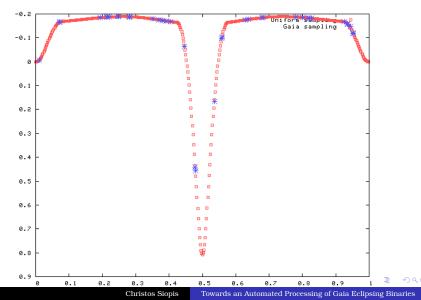
Particularities Estimation of Physical Parameters Remaining Work

## EB and Gaia Scanning Law



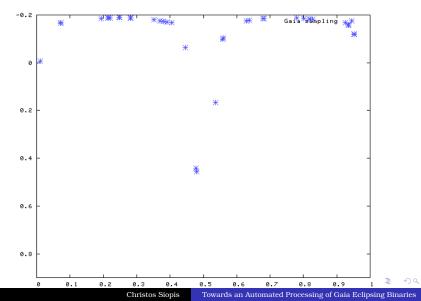
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- ► EB Model Generator: Given a set of epochs {*t<sub>i</sub>*} and physical parameters **p**, generate EB physical model *M*(*t<sub>i</sub>*; **p**)
  - ► Full Roche-lobe modeling

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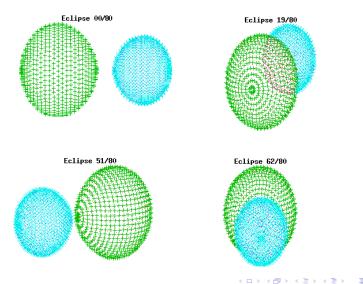
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Problem: Global minimum is hard and time-consuming to find!

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#### Following a two-step procedure:

- 1. Use database of precalculated light curves to come close to global minimum
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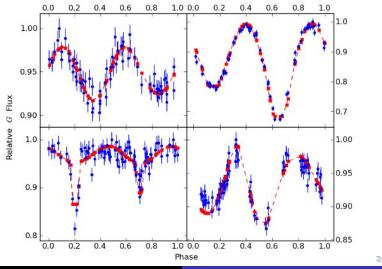
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Christos Siopis

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- ► Fine-tuning of the fitting procedure
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