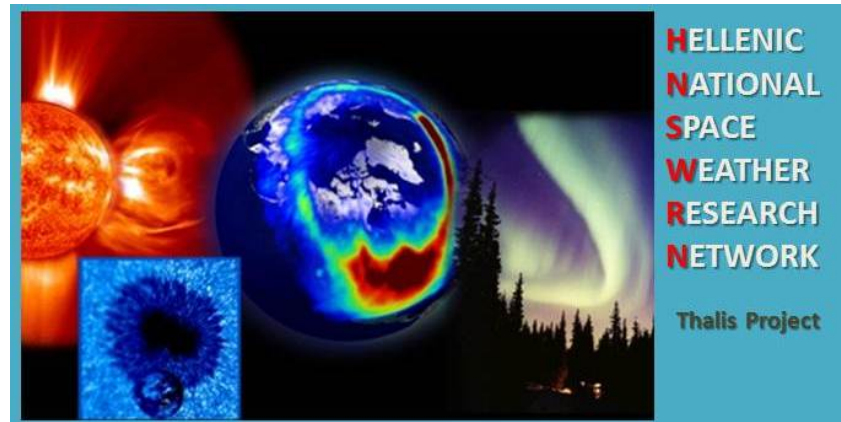
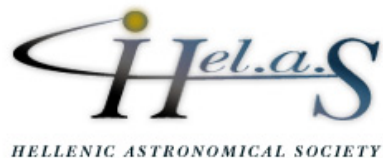


The applications of Complexity Theory and Tsallis Non-extensive Statistics at Space Plasma Dynamics

Thessaloniki, Greece
Thursday 30-6-2015

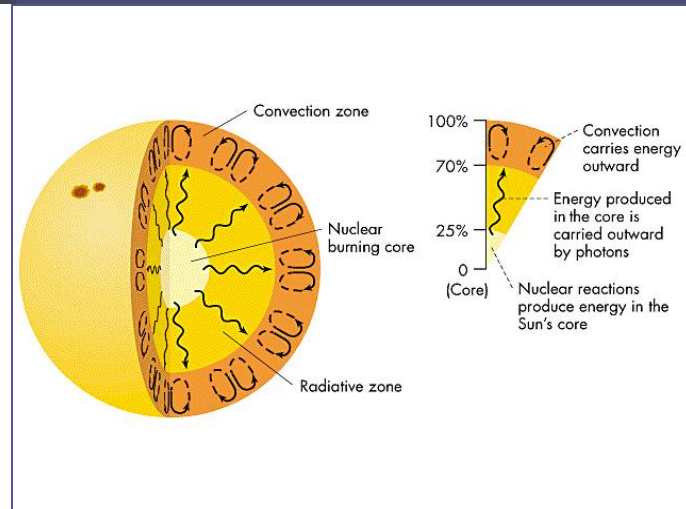
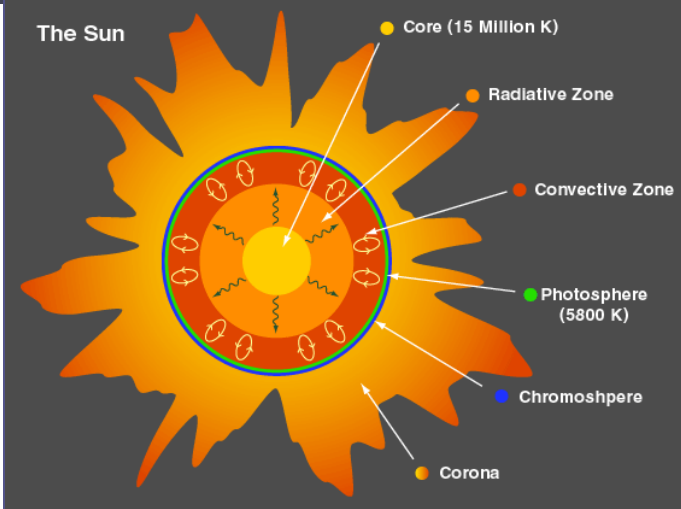
George P. Pavlos

Democritus University of Thrace
Department of Electrical and Computer Engineering
Xanthi, Greece

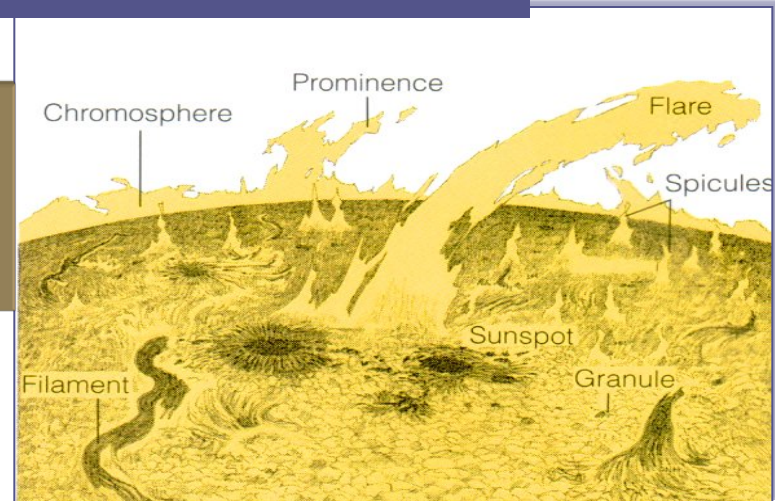
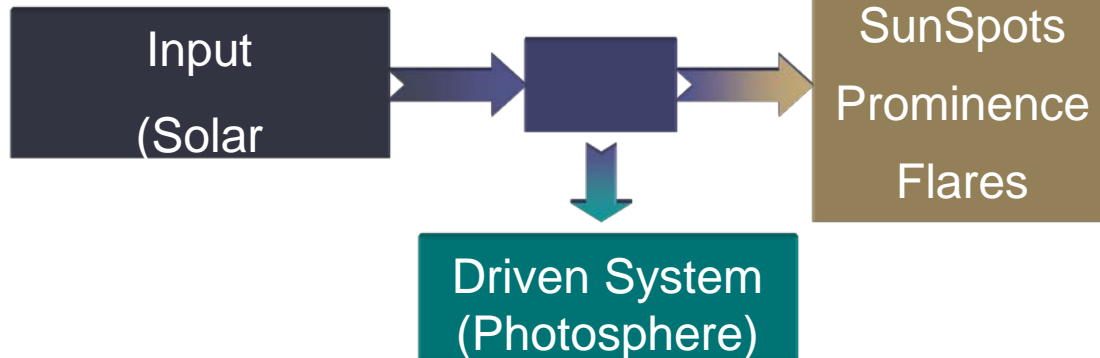


DRIVEN SYSTEMS (LOADING - UNLOADING)

Solar Dynamics

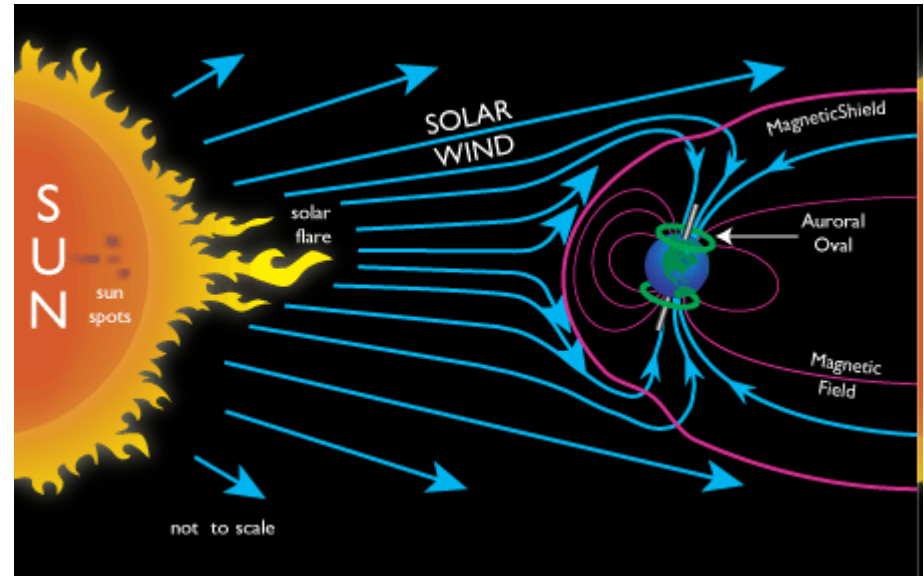
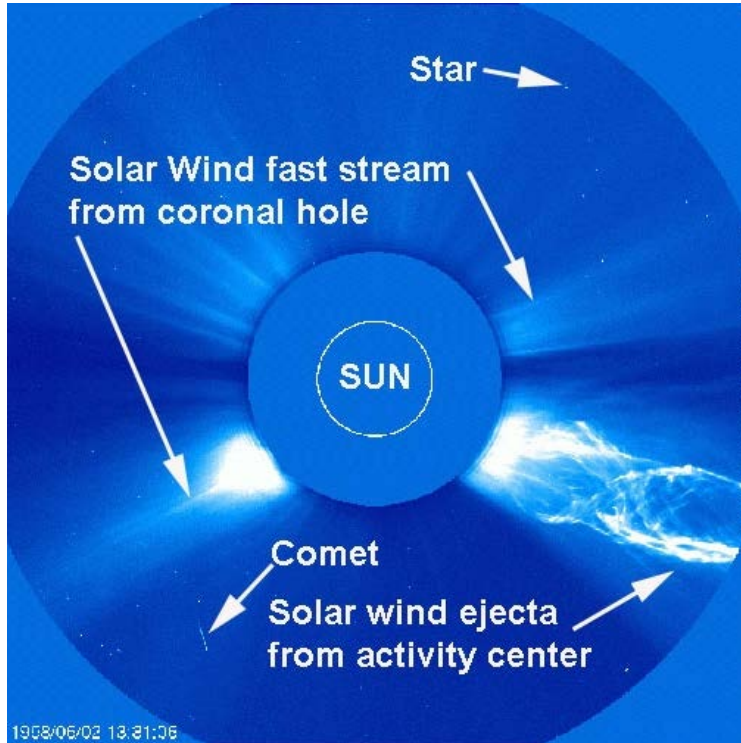


Description of Sun's Interior



Description of Sun's Photosphere

Solar Wind



Complexity Theory



Entropy Principle

Nonequilibrium self-organization

$$S_q = k \ln_q W = \max$$

$$F_q = U_q - TS_q = -\frac{1}{\beta} \ln_q Z_q = \min$$

F_q (Free energy) = minimum \Rightarrow Nonequilibrium stationary states (NESS)

Tsallis Non-extensive Statistical Mechanics

$$S_q = -k \sum_i p_i^q \ln_q p_i = k \frac{1 - \sum_i p_i^q}{q-1}$$

$$P_{opt} = \frac{[1 - (1-q)\beta E]^{1/1-q}}{Z_q} = \frac{e_q^{-\beta E}}{Z_q}$$

$$Z_q = \sum e_q^{-\beta E}$$

$q=1 \Rightarrow$ Boltzman – Gibbs Theory

THEORETICAL CONCEPTS

Near Thermodynamic Equilibrium

Linear Dynamics

Euclidean Geometry-Topology
Smooth Functions – Smooth differential Equations
Normal derivatives-integrals

MHD –Vlasov-Blotzman theory
Normal diffusion-Brownian motion
Gaussian statistics-dynamics
Normal Langevin-FP equations
Extensive statistics – BG entropy
Infinite dimensional noise
White-colored noise
Normal Central Limit Theorem (CLT)

Normal Liouville theory
Locality in space and time
Separation of time-spatial scales
Microscopic-macroscopic locality
Equilibrium RG

Far from Thermodynamic Equilibrium

Non-Linear Dynamics

Fractal Geometry -singular functions
Fractional differential-integral equations

Fractional MHD theory
Anomalous diffusion – motion
Strange Kinetics
Non-Gaussian statistics-dynamics
Fractional Langevin-FP equations
Non-Extensive statistics – Extremization of Tsallis entropy
q-extended CLT
Intermittent turbulence

Fractional Liouville theory
(multi)Fractal Topology
Power laws – multiscale processes
Memory – long range correlations
Nonlocality in space and time
Non-Equilibrium RG

Renormalization Group Theory (RGT) (Chang, 1992)

Non-Equilibrium – Non-linear Space Plasma Physics

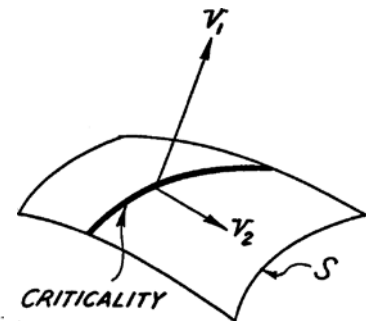
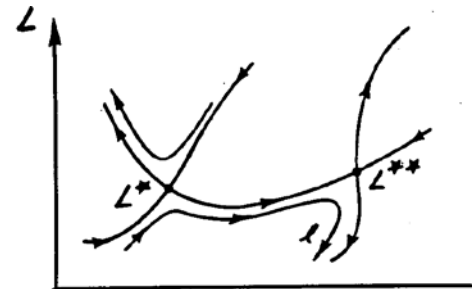
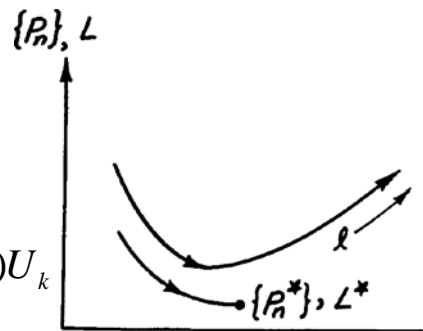
$$\frac{\partial \phi_i}{\partial t} = f_i(\phi, \mathbf{x}, t) + n_i(\mathbf{x}, t) \quad i = 1, 2, \dots \quad \text{Generalized Langevin Stochastic Equation}$$

$$P(\phi(\mathbf{x}, t)) = \int D(\mathbf{x}) \exp\left\{-i \cdot \int L(\phi, \phi, \mathbf{x}) d\mathbf{x}\right\} dt$$

Random Field Distribution Function
Plasma System Stochastic Lagrangian

Scale Invariance - RGT - Fixed Points (SOC, Chaos, Intermittency etc.)

$$\begin{aligned} \partial L / \partial l &= RL \\ \partial L' / \partial l &= R_L L' \quad \dot{\eta} \\ dP'_m / dl &= \Sigma(R_L)_{mn} P'_n \\ L'(l) &= \Sigma V_k(l) U_k = \Sigma V_{k0} \exp(\lambda_k l) U_k \end{aligned}$$



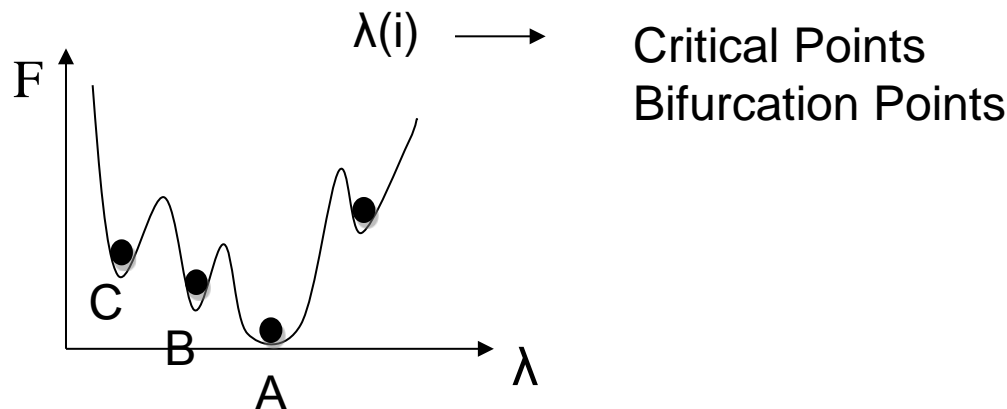
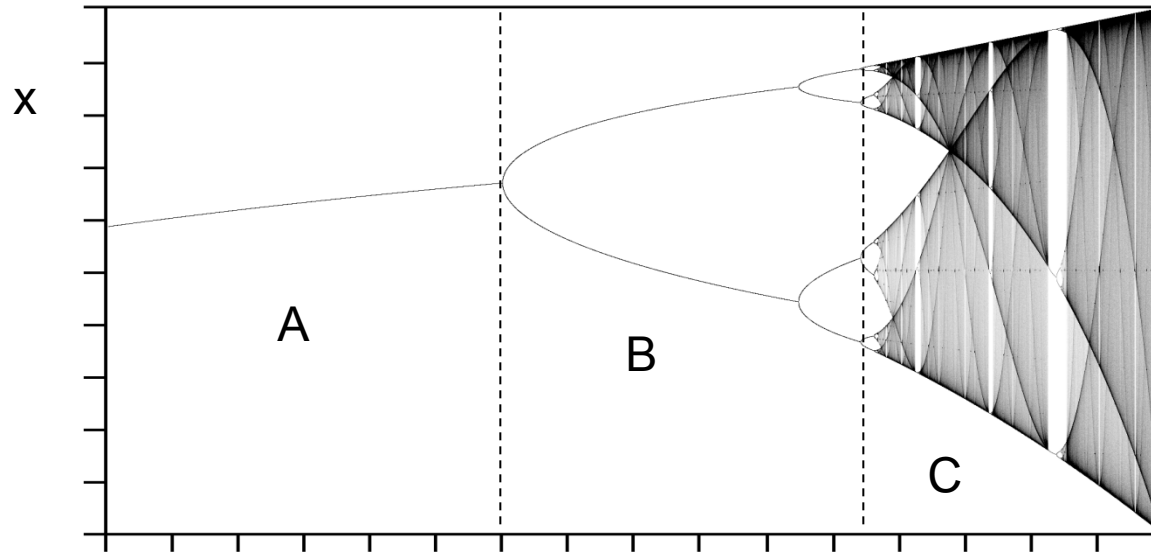
Non-equilibrium Non-extensive Random Field Theory (Partition Function Theory)

$$G_N^q(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_N) = \frac{1}{Z} \frac{\delta^N Z_q(J(\vec{x}))}{\delta J(\vec{x}_1) \cdot \delta J(\vec{x}_2) \cdot \dots \cdot \delta J(\vec{x}_N)} \quad Z_q = \lim_{J \rightarrow 0} Z(J(\vec{x})) \quad Z(J(\vec{x})) = \int D[\Psi] \cdot e^{-\int F_q(J) d^D x}$$

Dynamical Bifurcation

$$dx/dt = f(x, \lambda)$$

Non – Linear Dynamics



Critical Points
Bifurcation Points

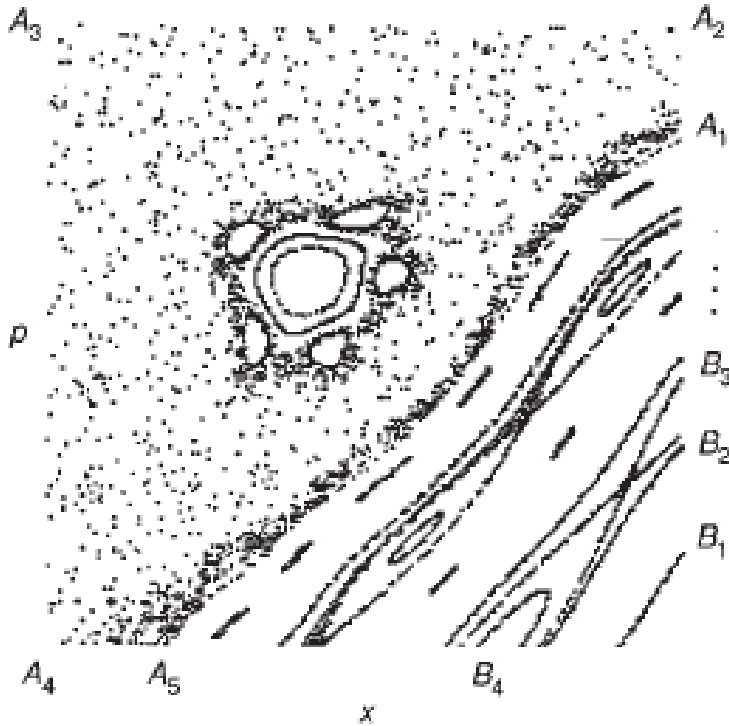
(A)
Gaussian Equilibrium
Critical States
Equilibrium Phase
Transition
Power Laws

(B)
Self-Organization Structure
Dissipative Structures
Long Range Correlations

(C)
Spatiotemporal Chaos
Strange Dynamics (Attractors)
Levy Processes
Scale Invariance
Intermittent Turbulence
Tsallis Entropy
Fractal Topology
Anomalous Diffusion

Strange kinetics

Michael F. Shlesinger, George M. Zaslavsky & Joseph Klafter NATURE · VOL 363 · 6 MAY 1993



Poincaré Section of Phase Space Intermittency in Phase Space

- Stochastic Sea
- Fractal Set (Island, Cantori)
- Trapping – Stickiness – Levy flights
- Lyapunov Exponents ≥ 0
- Strange Topology – Strange Dynamics
- Scale Invariance (RGT)

Levy Distribution – Anomalous Diffusion

$$\langle R^2(t) \rangle \sim t^\gamma \quad p_n(k) = \exp(-\text{constant} \times n|k|^\alpha)$$

$$\langle |R| \rangle \sim t^\mu \quad (t \leftarrow \infty) \quad p_n(x) \sim \text{constant} \times n/x^{1+\alpha}$$

$$\mathcal{B} = \lim_{\Delta t \rightarrow 0, \Delta x \rightarrow 0} |\Delta x|^{2\alpha} / |\Delta t|^\beta = \text{constant}$$

Basic Space Plasma Theory

BBGKY Hierarchy \longrightarrow Boltzmann - Vlasov, MHD Theory

Scale Invariance - Singularities

$$\vec{\tau}' = \lambda \vec{\tau} \quad t' = \lambda^{1-\alpha/3} t = \lambda^{1-h_t} \quad \vec{u}' = \lambda^{\alpha/3} \vec{u} = \lambda^h u \quad \vec{b}' = \lambda^{\alpha/3} \vec{b}'$$

Fractional Extension

Fractional Real - Space Derivative

$$\frac{\partial^\beta}{\partial x_i^\beta} f(t, \vec{r}) = \frac{1}{\Gamma(1-\beta)} \frac{\partial}{\partial x_i} \int_{-\infty}^x \frac{dx_i}{(x_i - x_i')^\beta} f(t, \vec{r})$$

$$\nabla_{\vec{r}}^\beta \equiv \frac{\partial^\beta}{\partial \vec{r}^\beta} \equiv \sum_i \hat{e}^i \frac{\partial^\beta}{\partial x_i^\beta}$$

$$\nabla_{\vec{r}}^{2\beta} \equiv \frac{\partial^{2\beta}}{\partial \vec{r}^{2\beta}} \equiv \sum_{i,k} \delta^{i,k} \frac{\partial^{2\beta}}{\partial x_i^\beta \partial x_k^\beta}$$

Riesz – Weyl operator

Levy flights – Levy walks

Power Law distribution

$$\varphi(\ell) = \frac{1}{\ell^{1+2\beta}}$$

Fractional Time Derivative

$$\frac{\partial}{\partial t^a} f(t, \vec{r}) = \frac{1}{\Gamma(m-a)} \frac{\partial^m}{\partial t^m} \int_0^t \frac{d\theta}{(t-\theta)^{1+\alpha-m}} f(\theta, \vec{r})$$

Riemann – Liouville operator

Fractal – time random walks (FTRW)

Fractal active time (Cantor set)

$1 \leq a \leq r$ persistent (super-diffusive) process

$0 \leq a \leq 1$ anti-persistent (sub-diffusive) process

Waiting time Power Law distribution

$$\phi(r) \sim \frac{1}{r^{1+a}}$$

FRACTIONAL KINETIC EQUATION (FRACTIONAL FOKKER-PLANCK-KOLMOGOROV EQUATION)

Zaslavsky G.M., Chaos, fractional kinetics, and anomalous transport, *Physics Reports* 371, 461-580, 2002.

$$\frac{\partial^\beta P(x, t)}{\partial t^\beta} = \frac{\partial^\alpha}{\partial (-x)^\alpha} (\mathcal{A}(x)P(x, y)) + \frac{\partial^{\alpha_1}}{\partial (-x)^{\alpha_1}} (\mathcal{B}(x)P(x, y))$$

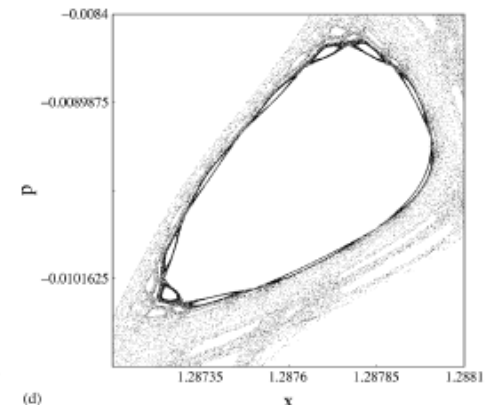
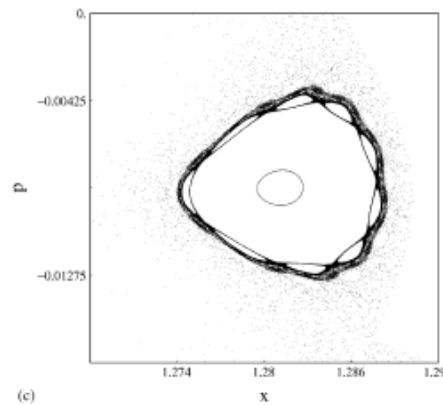
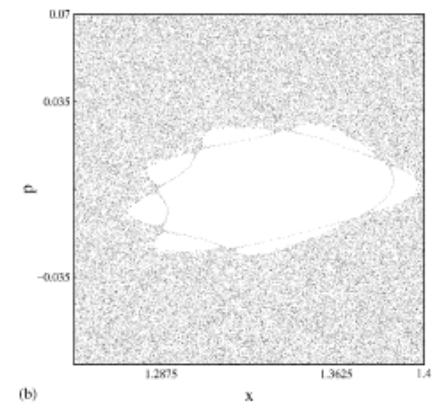
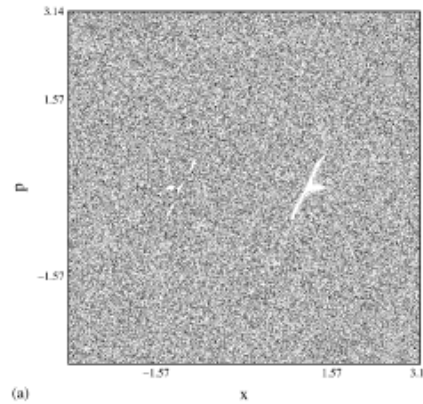
$$\mathcal{B}(x) = \frac{1}{\Gamma(2 + \alpha)} \lim_{\Delta t \rightarrow 0} \frac{\langle |\Delta x|^{\alpha+1} \rangle}{(\Delta t)^\beta},$$

$$\mathcal{A}(x) = \frac{1}{\Gamma(1 + \alpha)} \lim_{\Delta t \rightarrow 0} \frac{\langle |\Delta x|^\alpha \rangle}{(\Delta t)^\beta}$$

$$\langle R^2(t) \rangle \sim t^\gamma$$

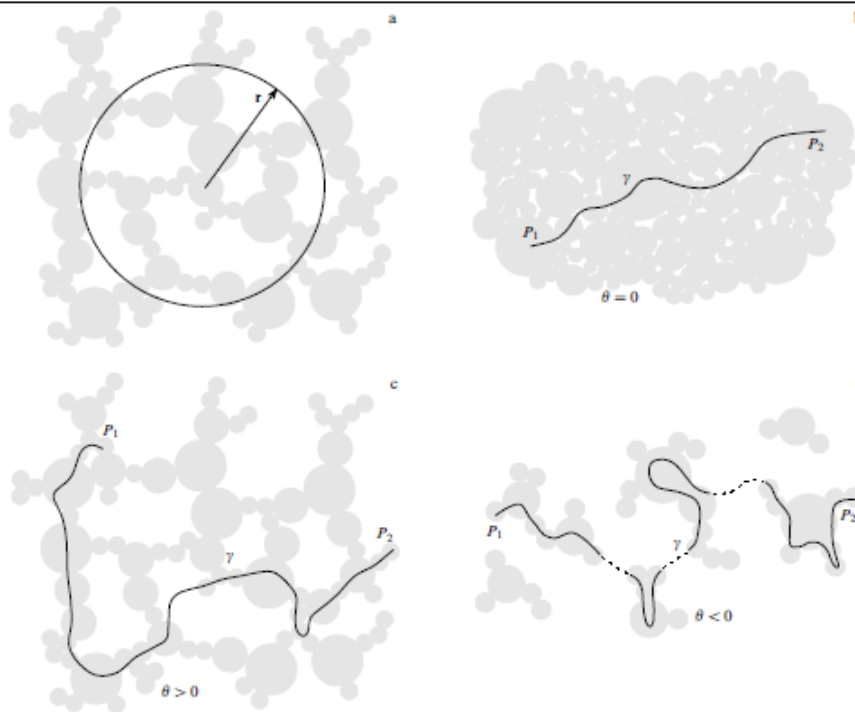
$$p_n(k) = \exp(-\text{constant} \times n |k|^\alpha)$$

$$p_n(x) \sim \text{constant} \times n / x^{1+\alpha}$$



Fractal topology and strange kinetics: from percolation theory to problems in cosmic electrodynamics

L M Zelenyĭ, A V Milovanov *Physics–Uspekhi* **47** (8) 749–788 (2004)



Mean – square displacement
Normal – Anomalous Diffusion

$$\langle x^2(t) \rangle \cong Dt^\mu = Dt^{2H}, 0 \leq H \leq 1$$

$$\mu = \frac{\beta}{\alpha} = \frac{ds}{df} = \frac{2}{2+\theta}, 1 < \mu \leq 2$$

(Non Gaussian dynamics)

$$\mu \neq 1, \left\{ \begin{array}{l} \theta > 0, \mu < 1 \leftrightarrow \text{subdiffusion} \\ \theta < 0, \mu > 1 \leftrightarrow \text{superdiffusion} \end{array} \right\}$$

$\mu = 2H$, d_s =spectral fractal dimension

θ =connectivity index

d_f = Hausdorff fractal dimension of space

H =Hurst exponent

$d_w=1/H$ =fractal dimension of trajectories

$\mu = 1, H = 1/2, d_w = 2 \leftrightarrow$ Normal diffusion (Gaussian dynamics)

μ ={competition between FTRW – Levy process}

Fractional Solar Plasma

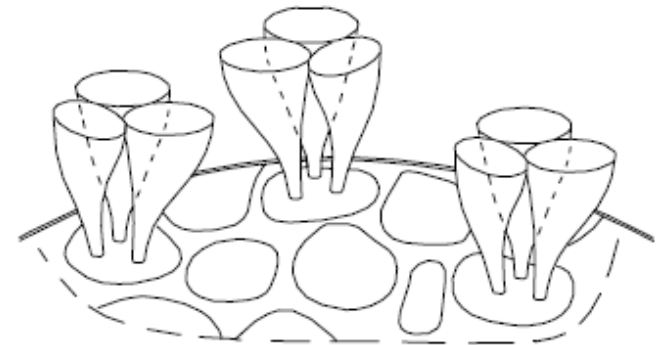
L M Zelenyĭ, A V Milovanov (1999,2004)

$$\frac{\partial^\alpha \psi}{\partial t^\alpha} = \nabla_{\mathbf{r}}^{2\beta} (\mathcal{B}\psi).$$

$$\frac{\partial^\alpha}{\partial t^\alpha} \psi(t, \mathbf{r}) = \frac{1}{\Gamma(m-\alpha)} \frac{\partial^m}{\partial t^m} \int_0^t \frac{d\vartheta}{(t-\vartheta)^{1+\alpha-m}} \psi(\vartheta, \mathbf{r}),$$

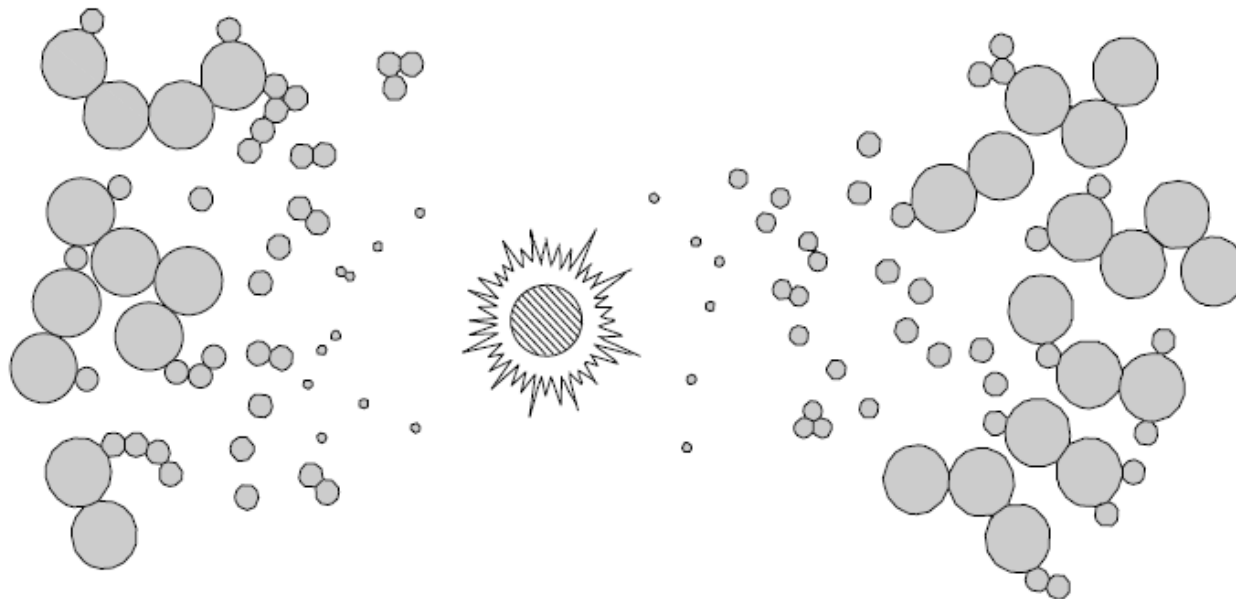
$$\frac{\partial^\beta}{\partial x_i^\beta} \psi(t, \mathbf{r}) = \frac{1}{\Gamma(1-\beta)} \frac{\partial}{\partial x_i} \int_{-\infty}^{x_i} \frac{dx'_i}{(x_i - x'_i)^\beta} \psi(t, \mathbf{r}').$$

Solar Photosphere



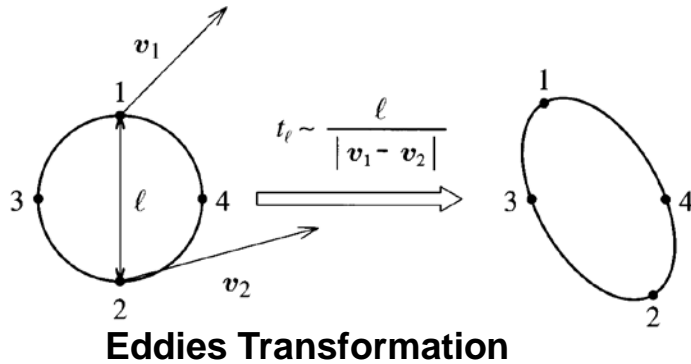
Magnetic Flux - Fractal Distribution

Solar Wind - Fracton Theory - Intermittent Turbulence



Turbulence - Intermittent Turbulence

Uriel Frisch - Turbulence (The Legacy of A.N. Kolmogorov)



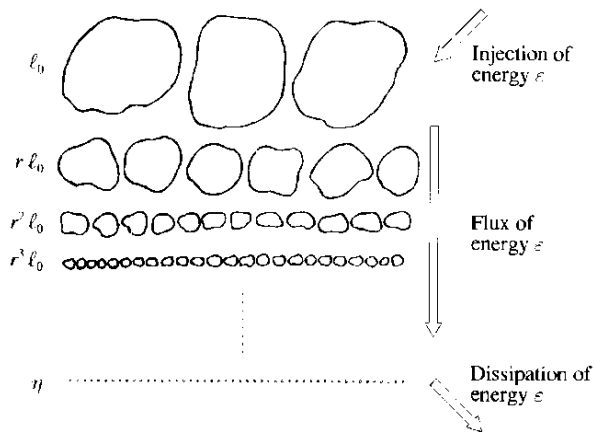
$$S_p(\ell) \equiv \langle (\delta v_{\parallel}(\ell))^p \rangle.$$

$$S_p(x) \equiv \langle [(\mathbf{v}(\mathbf{r} + x\ell^0) - \mathbf{v}(\mathbf{r})) \cdot \ell^0]^p \rangle$$

$$S_p(\ell) \propto \ell^{p/3}.$$

$$S_p(\ell) = C_p \varepsilon^{p/3} \ell^{p/3},$$

'mother - eddy' \rightarrow 'daughters'



**Space filling Cascade over Energy
dissipation (K41)**

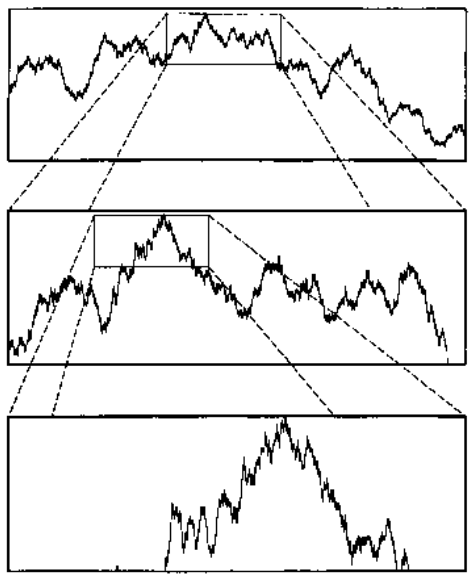
$$\Pi'_\ell \sim \frac{v_\ell^3}{\ell} \sim \varepsilon.$$

$$v_\ell \sim \varepsilon^{1/3} \ell^{1/3},$$

$$v_\ell \sim \sqrt{\langle \delta v_{\parallel}^2(\ell) \rangle},$$

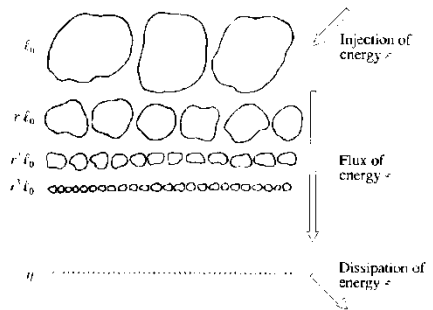
$$t_\ell \sim \frac{\ell}{v_\ell}.$$

Uriel Frisch – Turbulence (The Legacy of A.N. Kolmogorov)



Brownian self similarity

'mother – eddy' → 'daughters'



β-model cascade (eddies without space filling)

(β-model)

$$\beta \quad (0 < \beta < 1).$$

$$l = r^n l_0$$

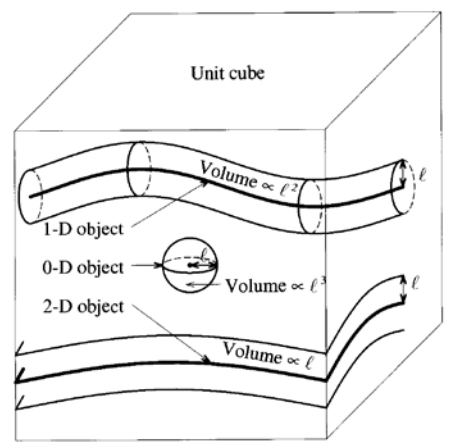


Fig. 8.10. The probability that a sphere of radius l encounters an object of dimension D behaves as l^{3-D} as $l \rightarrow 0$.

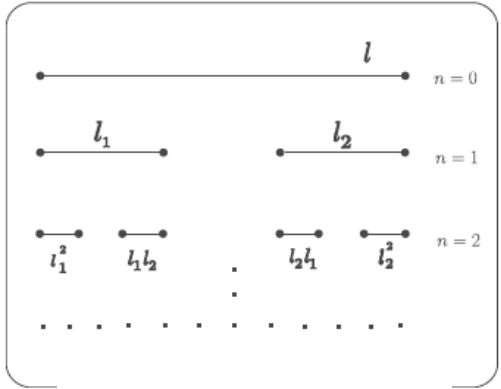
(probability density)

$$p_l = \beta^n = \beta^{\frac{\ln(l/l_0)}{\ln r}} = \left(\frac{l}{l_0}\right)^{3-D}$$

$$p_l \propto l^{3-D}, \quad l \rightarrow 0.$$

$$p_l \propto l^{d-D}, \quad l \rightarrow 0.$$

'p-model



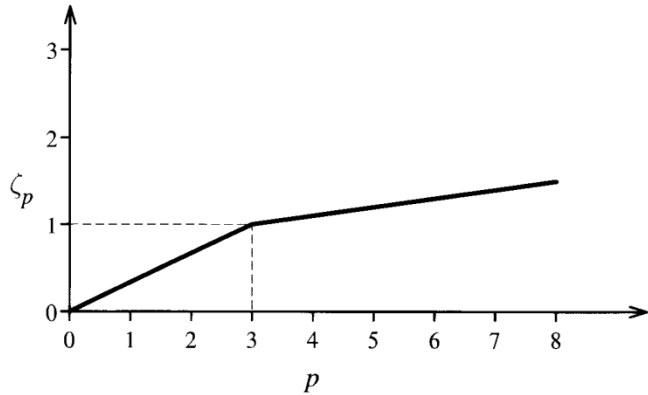
Two-scale Cantor set

$$\frac{p^q}{l_1^{\tau(q)}} + \frac{(1-p)^q}{l_2^{\tau(q)}} = 1$$

$$f(\alpha) = q\alpha(q) - \tau(q)$$

Multifractal intermittency velocity structure

Uriel Frisch – Turbulence (The Legacy of A.N. Kolmogorov)



bifractal model

$$v_\ell \sim v_0 \left(\frac{\ell}{\ell_0} \right)^{\frac{1}{3} - \frac{3-D}{3}},$$

singular exponent $h = \frac{1}{3} - \frac{3-D}{3}$

$$S_p(\ell) = \langle \delta v_\ell^p \rangle \sim v_0^p \left(\frac{\ell}{\ell_0} \right)^{\zeta_p}$$

Bifractal β Model

$$\frac{\delta v_\ell(r)}{v_0} \sim \begin{cases} \left(\frac{\ell}{\ell_0} \right)^{h_1}, & r \in \mathcal{S}_1, \dim \mathcal{S}_1 = D_1 \\ \left(\frac{\ell}{\ell_0} \right)^{h_2}, & r \in \mathcal{S}_2, \dim \mathcal{S}_2 = D_2. \end{cases}$$

singularity manifold fractal set-fractal dimension

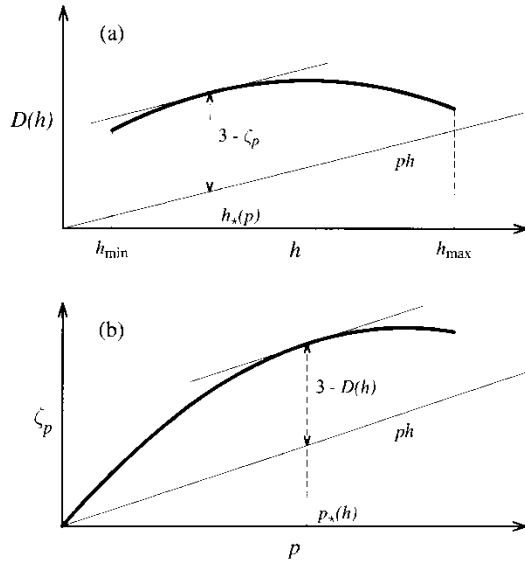
$$\langle \delta v_\ell^p \rangle \propto \ell^{\zeta_p}, \quad \zeta_p = \min(ph_1 + 3 - D_1, ph_2 + 3 - D_2).$$

$$\zeta_p = \begin{cases} p/3 & 0 \leq p \leq 3 \\ p/3 + (3 - D_2)(1 - p/3) & p \geq 3. \end{cases}$$

multiscaling exponents spectrum

Multifractal B model

Uriel Frisch – Turbulence (The Legacy of A.N. Kolmogorov)



Legendre transform

$$\frac{\delta v_\ell(\mathbf{r})}{v_0} \sim \left(\frac{\ell}{\ell_0}\right)^h, \quad \mathbf{r} \in \mathcal{S}_h.$$

$$\frac{S_p(\ell)}{v_0^p} \equiv \frac{\langle \delta v_\ell^p \rangle}{v_0^p} \sim \int_I d\mu(h) \left(\frac{\ell}{\ell_0}\right)^{ph+3-D(h)}.$$

Legendre transformation

$$\lim_{\ell \rightarrow 0} \frac{\ln S_p(\ell)}{\ln \ell} = \zeta_p,$$

$$\zeta_p = \inf_h [ph + 3 - D(h)].$$

$$\frac{S_p(\ell)}{v_0^p} \sim \left(\frac{\ell}{\ell_0}\right)^{\zeta_p}, \quad \ell \rightarrow 0.$$

$$\zeta_p = ph_*(p) + 3 - D(h_*(p)).$$

$$D'(h_*(p)) = p,$$

$$D(h) = \inf_p (ph + 3 - \zeta_p).$$

$$\frac{d\zeta_p}{dp} = h_*(p) + [p - D'(h_*(p))] \frac{dh_*(p)}{dp} = h_*(p)$$

H_{pmf} Under the same assumptions as in H1, there is a universal function $D(h)$ which maps real scaling exponents h to scaling dimensions $D \leq 3$ (including negative values and the value $-\infty$), such that for any h , the probability of velocity increments satisfies

$$\lim_{\ell \rightarrow 0} \frac{\ln \bar{P}_\ell^{\text{inc}}(\pm \ell^h)}{\ln \ell} = 3 - D(h). \quad (8.50)$$

Multifractal Intermittent Energy Dissipation

$$\varepsilon_\ell(\mathbf{r}) = \frac{1}{(4/3)\pi\ell^3} \int_{|\mathbf{r}'-\mathbf{r}|<\ell} d^3r' \frac{1}{2} v \sum_{ij} [\partial_j v_i(\mathbf{r}') + \partial_i v_j(\mathbf{r}')]^2$$

Definition. The dissipation is said to be multifractal if there is a function $F(\alpha)$ which maps real scaling exponents α to scaling dimensions $F \leq 3$ (including negative values and the value $-\infty$), such that for any α

$$\lim_{\ell \rightarrow 0} \frac{\ln \bar{P}_\ell^{\text{diss}}(\ell^{\alpha-1})}{\ln \ell} = 3 - F(\alpha), \quad (8.79)$$

$$\left. \begin{aligned} \frac{\varepsilon_\ell(\mathbf{r})}{v_0^3/\ell_0} &\sim \left(\frac{\ell}{\ell_0}\right)^{\alpha-1} \quad \text{as } \ell \rightarrow 0, \\ \text{for } \mathbf{r} \in \mathcal{D}_\alpha \subset \mathbb{R}^3; \quad \dim \mathcal{D}_\alpha &= F(\alpha). \end{aligned} \right\}$$

$$\langle \varepsilon_\ell^q \rangle \sim \left(\frac{v_0^3}{\ell_0}\right)^q \left(\frac{\ell}{\ell_0}\right)^{\tau_q}, \quad \tau_q = \min_\alpha [q(\alpha-1) + 3 - F(\alpha)]$$

$$\varepsilon_\ell(x) \equiv \frac{1}{2\ell} \int_{|x'-x|<\ell} dx' \frac{1}{2} v \sum_{ij} [\partial_j v_i(x') + \partial_i v_j(x')]^2$$

$$\frac{\varepsilon_\ell(x)}{v_0^3/\ell_0} \sim \left(\frac{\ell}{\ell_0}\right)^{\alpha-1} \quad \text{as } \ell \rightarrow 0 \quad \text{for } x \in \mathcal{D}'_\alpha; \quad \dim \mathcal{D}'_\alpha = f(\alpha),$$

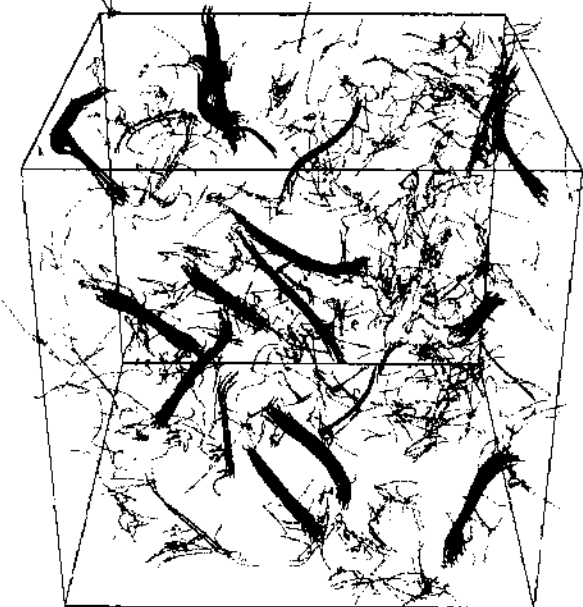
$$f(\alpha) \equiv F(\alpha) - 2.$$

Velocity – Energy Dissipation Multifractal Structures

$$h = \frac{\alpha}{3}, \quad D(h) = F(\alpha) = f(\alpha) + 2, \quad \zeta_p = \frac{p}{3} + \tau_{p/3}.$$

Space time scale invariance

$$\mathbf{r}' = \lambda \mathbf{r}, \quad \mathbf{u}' = \lambda^{\alpha/3} \mathbf{u}, \quad t' = \lambda^{1-\alpha/3} t, \quad (p/\rho)' = \lambda^{2\alpha/3} (p/\rho).$$



**Intermittent vortex filaments.
Three dimensional turbulence simulation**

Multifractal Theory

Theiler J., Vol. 7, No. 6/June 1990/J. Opt. Soc. Am. A, 1055

generalized dimensions $D_q = \frac{1}{q-1} \lim_{r \rightarrow 0} \frac{\log \sum_i P_i^q}{\log r}$.

most-dense points $D_\infty = \lim_{r \rightarrow 0} \frac{\log(\max_i P_i)}{\log r}$,

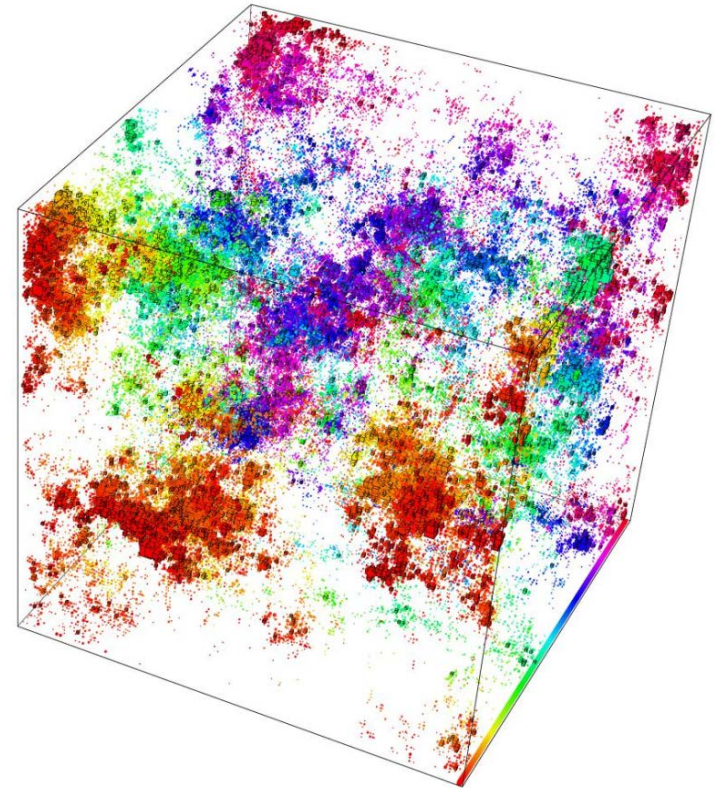
least-dense points $D_{-\infty} = \lim_{r \rightarrow 0} \frac{\log(\min_i P_i)}{\log r}$.

Information entropy $S(r) = - \sum_i P_i \log_2 P_i$

$$D_I = \lim_{r \rightarrow 0} \frac{-S(r)}{\log_2 r}$$

Information dimension
$$= \lim_{r \rightarrow 0} \frac{\sum_i P_i \log_2 P_i}{\log_2 r}$$
.

Rényi entropy $S_q(r) = \frac{1}{q-1} \log \sum_i P_i^q$,



$$D_q = \lim_{r \rightarrow 0} \frac{-S_q(r)}{\log r} = \frac{1}{q-1} \lim_{r \rightarrow 0} \frac{\log \sum_i P_i^q}{\log r},$$

Multifractal Theory

Theiler J., Vol. 7, No. 6/June 1990/J. Opt. Soc. Am. A, 1055

$$P_i = r^\alpha$$

$$n(\alpha, r) \sim r^{-f(\alpha)} \Delta\alpha.$$

$$\sum_i P_i^q = \int n(\alpha, r) r^{q\alpha} d\alpha$$

$$\theta = \min_\alpha \{q\alpha - f(\alpha)\}$$

$$\sim \int r^{-f(\alpha)} r^{q\alpha} d\alpha \sim r^\theta,$$

$$\sum_i P_i^q \sim r^{(q-1)D_q},$$

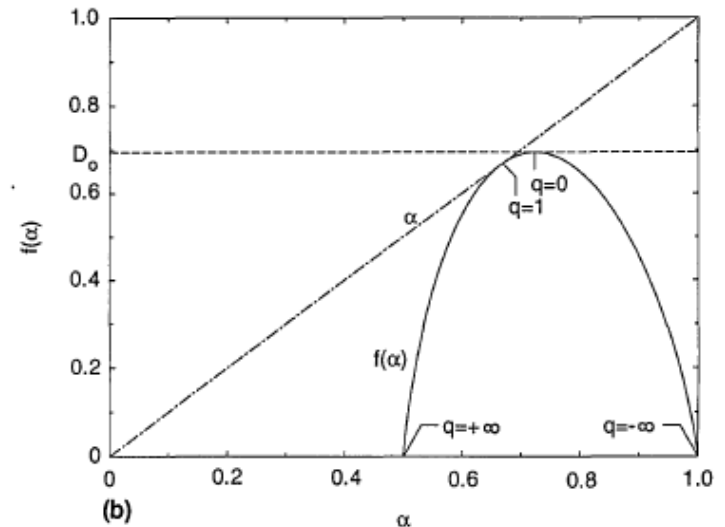
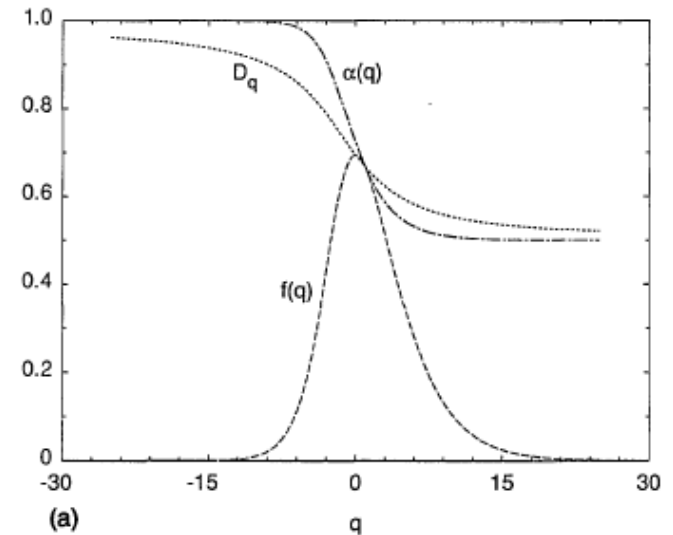
$$\alpha = \frac{\partial \tau}{\partial q} \quad f = \alpha q - \tau$$

and

$$q = \frac{\partial f}{\partial \alpha} \quad \tau = \alpha q - f.$$

$$f(\alpha) = \min_q \{q\alpha - \tau(q)\},$$

$$\tau(q) = \min_\alpha \{q\alpha - f(\alpha)\}.$$



Tsallis Extension of Statistics

Nonextensive Statistical Mechanics

Microscopic Level



Macroscopic Level

Quantum Complexity
Quantum Phase
Transition
(Q.P.T.)

Equilibrium Phase Transition
(E.P.T.)

Non – Equilibrium
Phase Transition
(N.E.P.T.)

Tsallis Theory

Nonextensive Statistical Mechanics

$$S_q = k \ln_q W \quad (S_1 = S_{BG}). \quad S_q = k \langle \ln_q(1/p_i) \rangle. \quad S_q = k \frac{1 - \sum_{i=1}^W P_i^q}{q - 1}.$$

$$y = [1 + (1 - q)x]^{1/(1-q)} \equiv e_q^x \quad (e_1^x = e^x). \quad \frac{dy}{dx} = y^q \quad y = \frac{x^{1-q} - 1}{1 - q} \equiv \ln_q x \quad (x > 0; \ln_1 x = \ln x),$$

$$\frac{S_q(A + B)}{k} = \frac{S_q(A)}{k} + \frac{S_q(B)}{k} + (1 - q) \frac{S_q(A)}{k} \frac{S_q(B)}{k}. \quad S_q[A + B] = S_q[A] + S_q[B|A] + (1 - q)S_q[A]S_q[B|A],$$

Generalized Fokker-Planck Equations

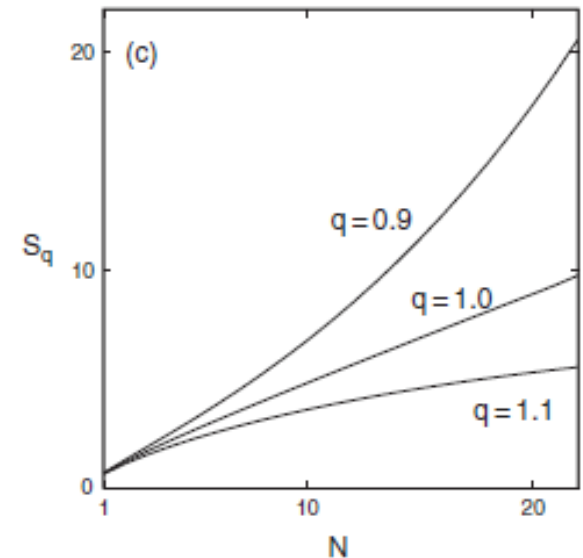
$$\frac{\partial^\beta p(x, t)}{\partial |t|^\beta} = D_{\beta, \gamma, q} \frac{\partial^\gamma [p(x, t)]^{2-q}}{\partial |x|^\gamma} \quad (0 < \beta \leq 1; 0 < \gamma \leq 2).$$

$$\frac{\partial p(x, t)}{\partial t} = D_{\gamma, q} \frac{\partial^\gamma [p(x, t)]^{2-q}}{\partial |x|^\gamma} \quad (0 < \gamma \leq 2; q < 3).$$

$$L_\gamma(x) \propto \frac{1}{|x|^{1+\gamma}} \quad (|x| \rightarrow \infty; 0 < \gamma < 2), \quad \text{Levy Distribution}$$

$$p_q(x) \propto \frac{1}{|x|^{2/(q-1)}} \quad (|x| \rightarrow \infty; 1 < q < 3). \quad \text{q-Gaussian Distribution}$$

$$\gamma = \begin{cases} 2 & \text{if } q \leq 5/3, \\ \frac{3-q}{q-1} & \text{if } 5/3 < q < 3, \end{cases}$$



q-extension of thermodynamics

$$Z_q = \sum_{conf} e_q^{-\beta q(E_i - V_q)}$$

$$\beta_q = \beta / \sum_{conf} p_i^q \quad \beta = 1 / KT$$

$$\langle E \rangle_q \equiv \sum_{conf} p_i^q E_i / \sum_{conf} p_i^q = U_q$$

$$F_q \equiv U_q - TS_q = -\frac{1}{\beta} \ln qZ_q$$

$$U_q = \frac{\partial}{\partial \beta} \ln qZ_q, \quad \frac{1}{T} = \frac{\partial S_q}{\partial U_q}$$

$$C_q \equiv T \frac{\partial \delta_q}{\partial T} = \frac{\partial U_q}{\partial T} = -T \frac{\partial^2 F_q}{\partial T^2}$$

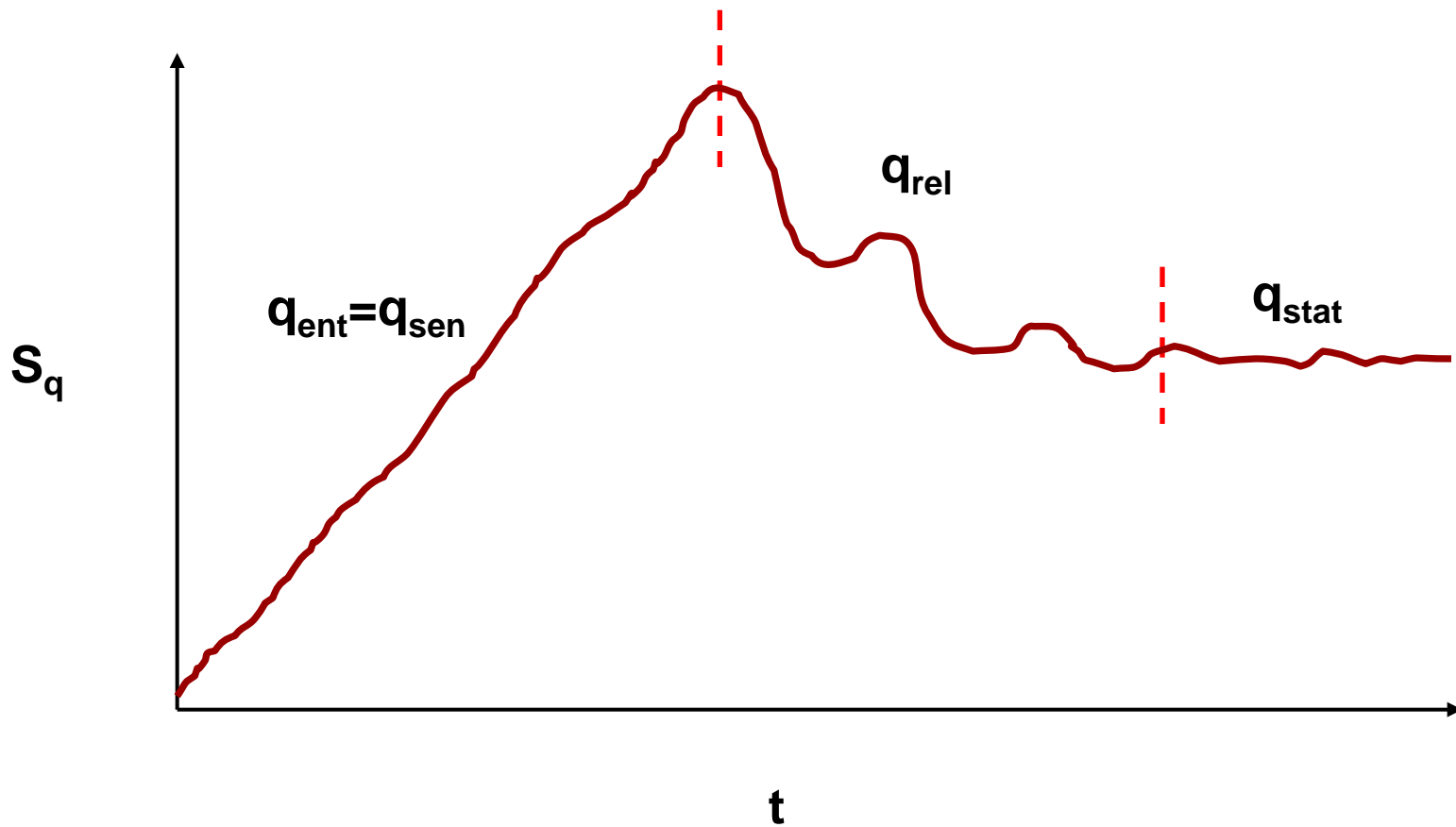
q-extension of central limit theorem

q-independent random variables

$$F_q[X + Y](\xi) = F_q[X](\xi) \otimes_{\frac{1+q}{3-q}} F_q[Y](\xi), \quad F_q[f](\xi) \equiv \int dx e_q^{i\xi x} \otimes_q f(x).$$

where $h(x, y)$ is the joint distribution. Therefore, *q-independence* means *independence* for $q = 1$ (i.e., $h(x, y) = f_X(x) f_Y(y)$), and it means *strong correlation* (of a certain class) for $q \neq 1$ (i.e., $h(x, y) \neq f_X(x) f_Y(y)$).

q-triplet Tsallis One-Dimensional Systems (Timeseries)



The q-triplet of Tsallis

$$\mathbf{q - CLT} \longrightarrow (q_{k-1}, q_k, q_{k+1}) = q_{stat}, q_{rel}, q_{sen}$$

Gaussian-BG equilibrium ($q_{stat}=q_{sen}=q_{rel}=1$) \longrightarrow Nonequilibrium ($q_{stat}, q_{sen}, q_{rel}$)

$$\frac{dy}{dx} = y^q, (y(0) = 1, q \in \mathfrak{R}) \longrightarrow (q_{stat}, q_{sen}, q_{rel})$$

Equilibrium PDF \longrightarrow Metaequilibrium PDF

$$\mathbf{(q_{stat})} \quad \frac{d(p_i Z_{stat})}{dE_i} = -\beta q_{stat} (p_i Z_{stat})^{q_{stat}} \longrightarrow p(x) \propto [1 - (1-q)\beta_{q_{stat}} x^2]^{1/(1-q)}$$

Equilibrium BG entropy production \longrightarrow Metaequilibrium q-entropy production

$$\mathbf{(q_{sen})} \quad K_q \equiv \lim_{t \rightarrow \infty} \lim_{W \rightarrow \infty} \lim_{N \rightarrow \infty} \frac{\langle S_q \rangle(t)}{t} \longrightarrow \frac{1}{1 - q_{sen}} = \frac{1}{\alpha_{min}} - \frac{1}{\alpha_{max}} \quad f(a_{min}) = f(a_{max}) = 0$$

$$\frac{d\xi}{dt} = \lambda_{q_{sen}} \xi^{q_{sen}}, \longrightarrow \xi = e^{\lambda_{q_{sen}} t}, \quad \xi \equiv \lim_{\Delta x(0) \rightarrow 0} \frac{\Delta x(t)}{\Delta x(0)}$$

Equilibrium relaxation process \longrightarrow Metaequilibrium nonextensive relaxation process

$$\mathbf{(q_{rel})} \quad \Omega(t) \equiv [O(t) - O(\infty)] / [O(0) - O(\infty)] \frac{d\Omega}{dt} = -\frac{1}{T_{q_{rel}}} \Omega^{q_{rel}} \longrightarrow \Omega(t) = e^{-t/T_{q_{rel}}}$$

Intermittent Turbulence - Tsallis Theory

Fractal dissipation of Energy

$$\varepsilon_n \sim \varepsilon_0 \left(l_n / l_0 \right)^{\alpha-1}$$

$$\sum_n \varepsilon_n^q l_n^d \sim l_n^{(q-1)D_n} = l_n^{\tau(q)}$$

$$dn(\alpha) \sim l_n^{-f_d(\alpha)} d\alpha$$

$$\left. \begin{aligned} f_d(a) &= a\bar{q} - (\bar{q} - 1)(D_{\bar{q}} - d + 1) + d - 1 \\ a &= \frac{d}{d\bar{q}} [(\bar{q} - 1)(D_{\bar{q}} - d + 1)] \end{aligned} \right\}$$

$$f(a) = a\bar{q} - \tau(\bar{q}),$$

$$a = \frac{d}{d\bar{q}} [(q-1)D_q] = \frac{d}{d\bar{q}} \tau(\bar{q})$$

$$\bar{q} = \frac{df(a)}{da}$$

Tsallis Entropy Extremization - Singular Spectrum

$$P(a) = Z_q^{-1} \left[1 - (1-q) \frac{(a-a_0)^2}{2X/\ln 2} \right]^{\frac{1}{1-q}}$$

$$f(a) = D_0 + \log_2 \left[1 - (1-q) \frac{(a-a_0)^2}{2X/\ln 2} \right] / (1-q)^{-1}$$

$$\tau(\bar{q}) = \bar{q}a_0 - 1 - \frac{2X\bar{q}^2}{1 + \sqrt{C_{\bar{q}}}} - \frac{1}{1-q} [1 - \log_2(1 + \sqrt{C_{\bar{q}}})]$$

$$a_{\bar{q}} - a_0 = (1 - \sqrt{C_{\bar{q}}}) / [\bar{q}(1-q)\ln 2]$$

$$\frac{1}{1-q} = \frac{1}{a_-} - \frac{1}{a_+}$$

Intermittence Turbulence Spectrum

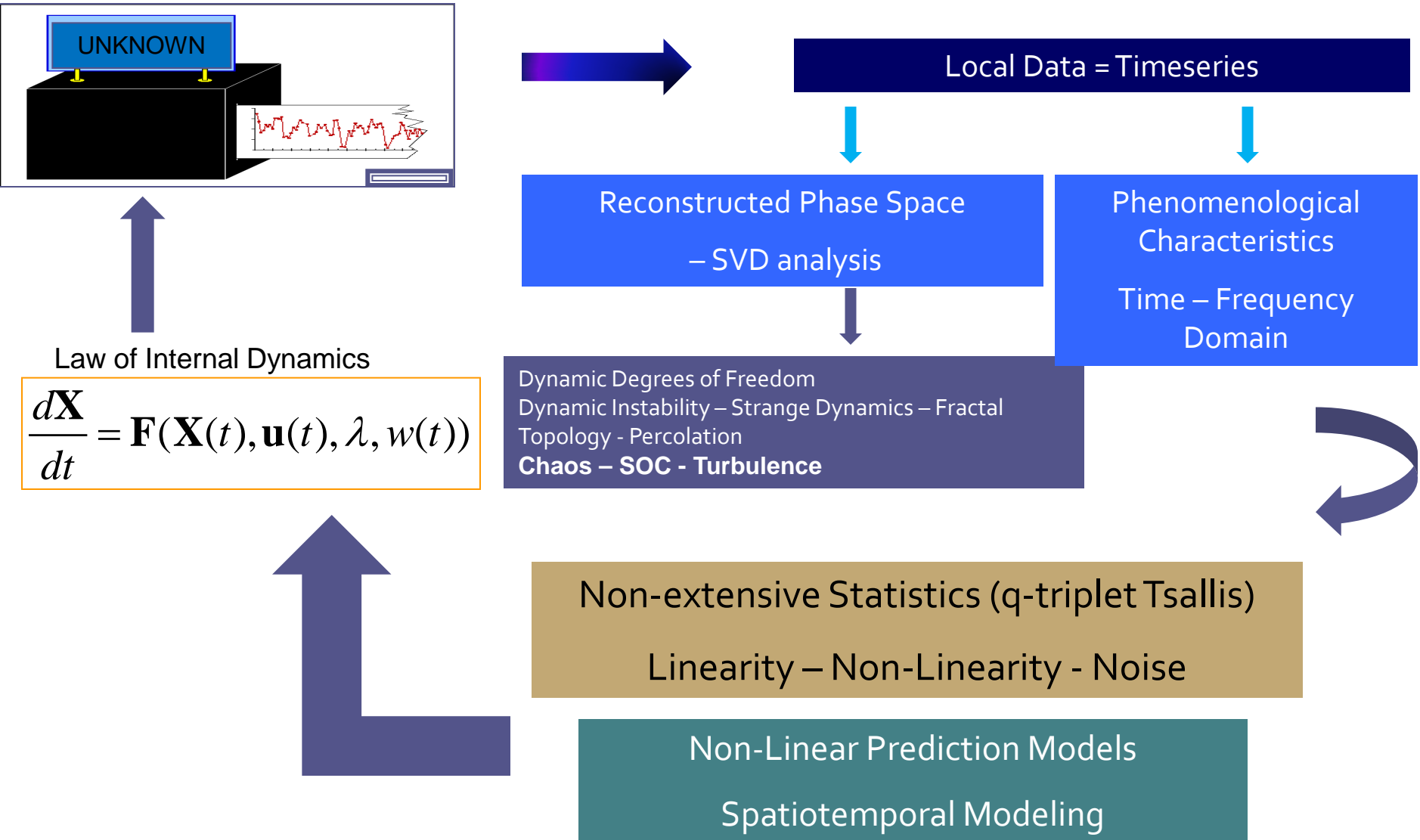
$$Sp(l^n) = \langle |\delta u_n|^p \rangle$$

$$J(p) = 1 + \tau(\bar{q} = \frac{p}{3})$$

$$J(p) = \frac{a_0 p}{3} - \frac{2Xp^2}{q(1 + \sqrt{C_{p/3}})} - \frac{1}{1-q} [1 - \log_2(1 + \sqrt{C_{p/3}})]$$

$$J(p) = \frac{p}{3} + T^{(u)}(p) + T^{(F)}(p)$$

FLOW CHART OF NON LINEAR ANALYSIS



CHAOTIC ALGORITHM

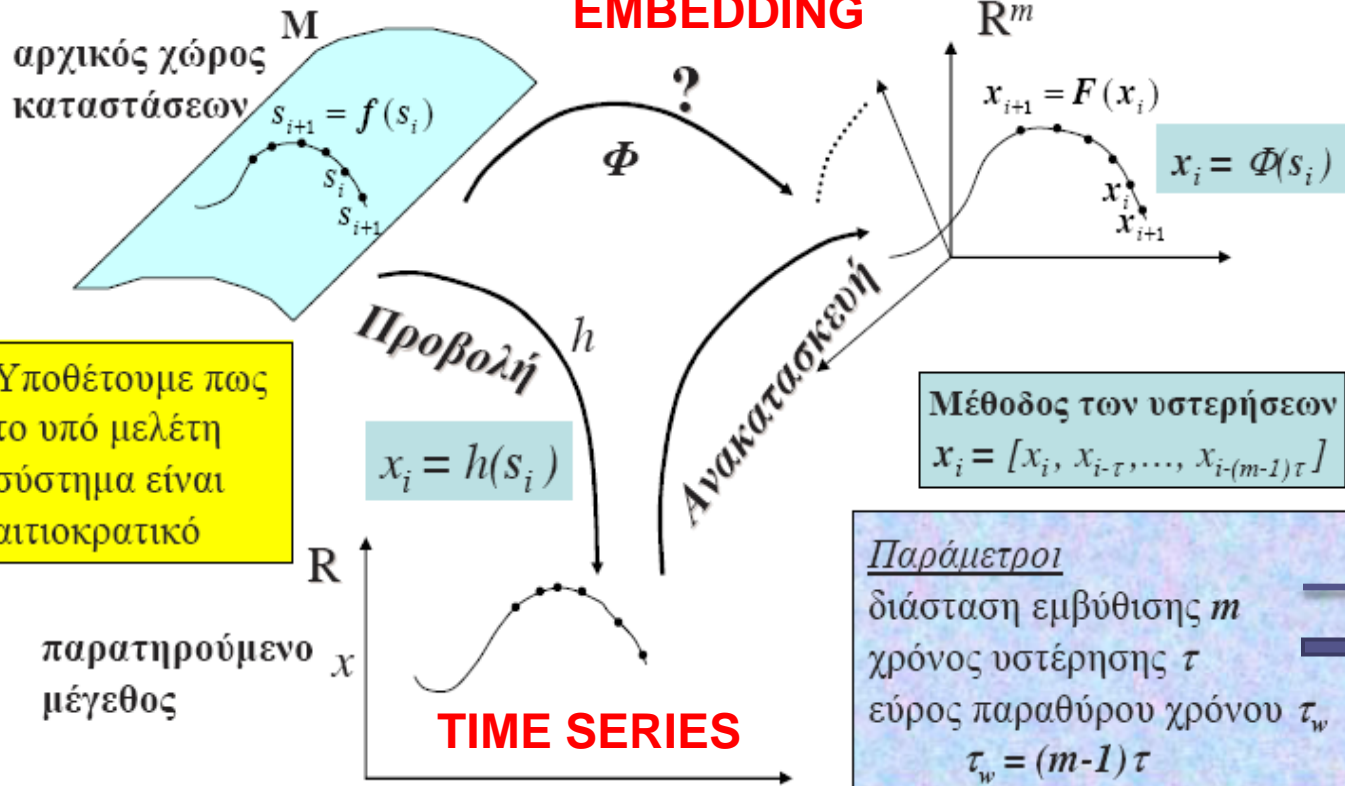
Takens Theorem (1981)

State Space Reconstruction

ORIGINAL STATE SPACE συνθήκη: $m \geq 2D + 1$

RECONSTRUCTED STATE SPACE

EMBEDDING



The Chaotic Algorithm is based in proficient mathematical concepts: **Embedding theory, metrics, fractals, multiple manifolds, probability – information theory, dynamical systems and maps**

- ✓ m small, self-cutting
- ✓ m big, noise
- ✓ Autocorrelation
- ✓ Mutual Information

COMPLEXITY AND TIME SERIES ANALYSIS

I. Analysis in the time-frequency domain

- Autocorrelation and Power Spectrum (noise-chaoticity)
- Mutual Information (Nonlinear Correlations)
- Flatness Coefficient (F) (Gaussian – non-Gaussian process)
- Hurst Exponent (temporal fractality – singularities, persistence – anti persistence, normal – anomalous diffusion)
- structure function (normal – intermittent turbulence, singularity spectrum ($h, D(h)$)
- q-entropy production (S_q)

II. Fractal geometry and fractional dynamics

- Singularity spectrum ($a, f(a)$)
- Generalized Spectrum ($\bar{q}, D(\bar{q})$)
- P –model (multiplicative process)
- Probability distribution function (Tsallis distributions)
- q-extension of CLT and Tsallis q-triplet parameters ($q_{sen}, q_{rel}, q_{stat}$)
- Tsallis modeling of non-equilibrium stationary states (NESS)

III. Low dimensionality and Nonlinearity

- Correlation Dimension of reconstructed phase space and existence of strange attractor.
- Strange Dynamics and sensitivity of initial conditions (Lyapunov exponents spectrum)
- Discrimination of stochasticity and nonlinear randomness - chaoticity
- Discrimination of dynamical components

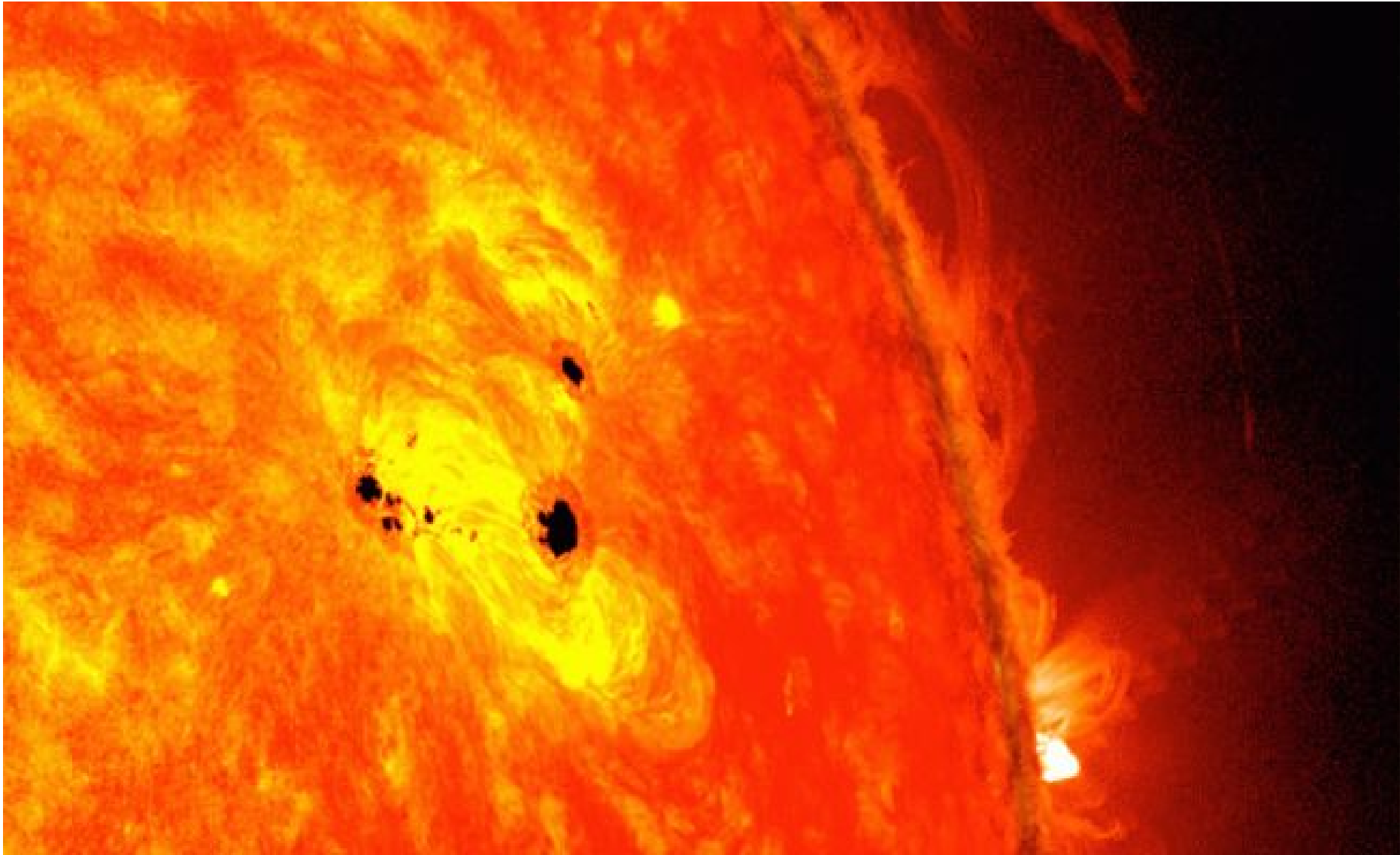
Synopsis

Chaotic Analysis
Dimensional Analysis

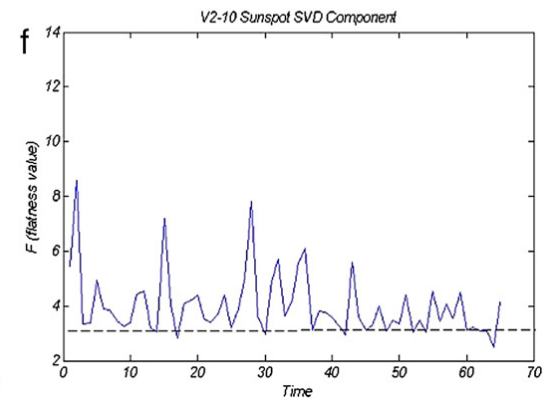
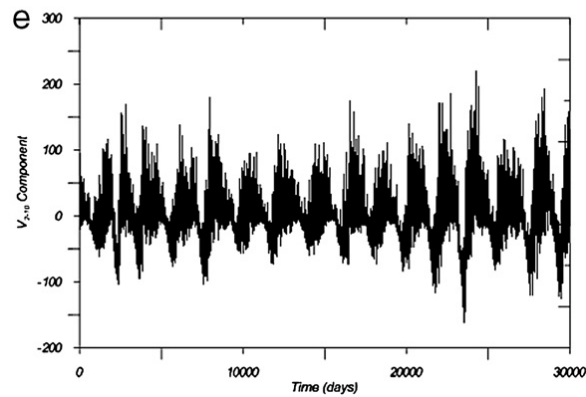
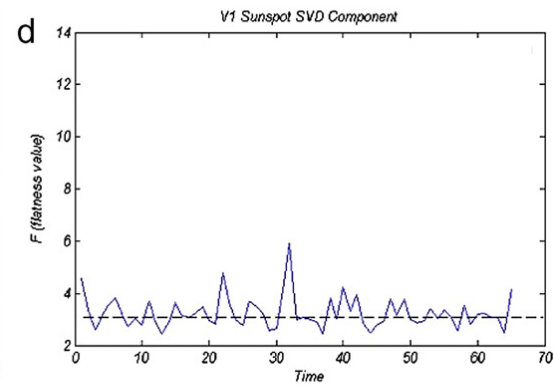
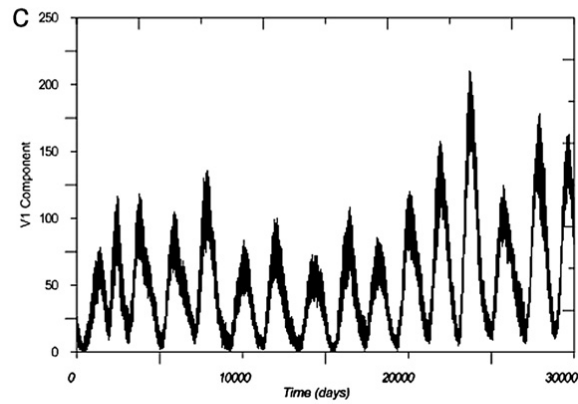
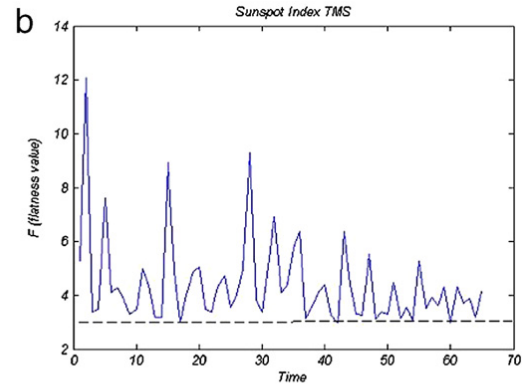
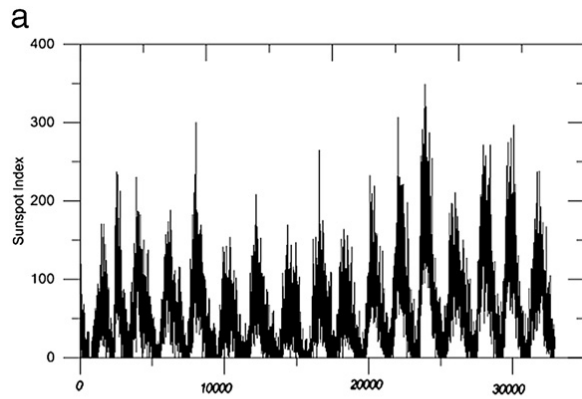
Statistical Analysis

Turbulence Analysis

Sunspot



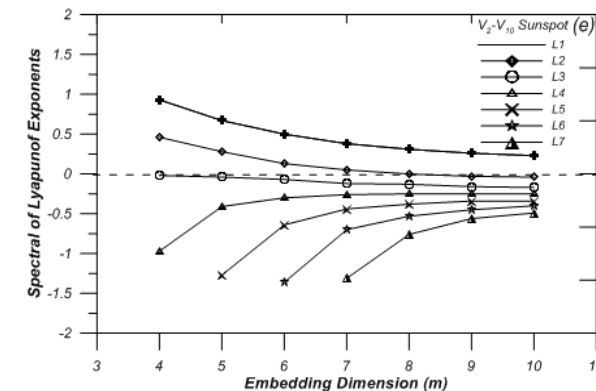
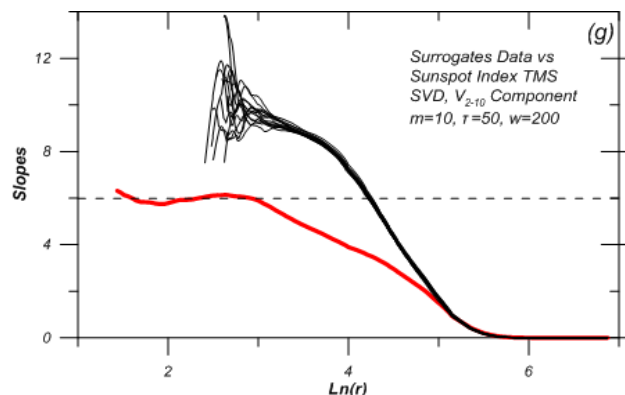
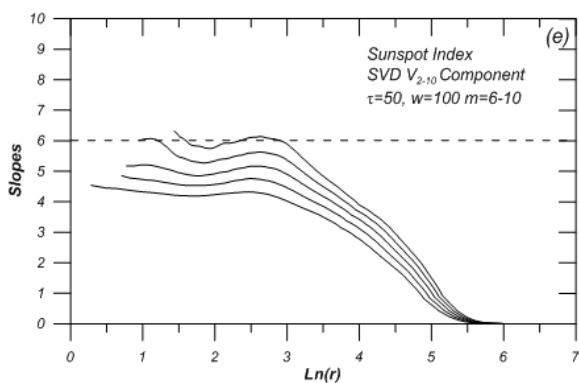
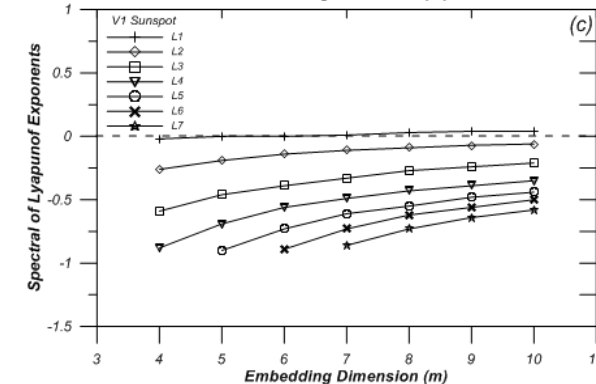
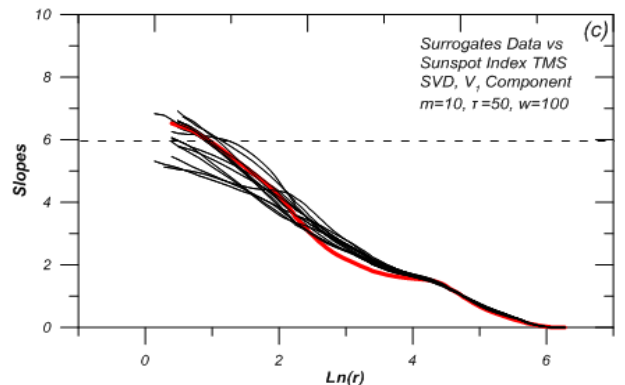
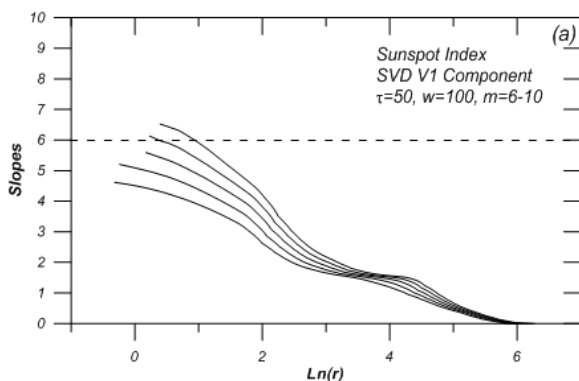
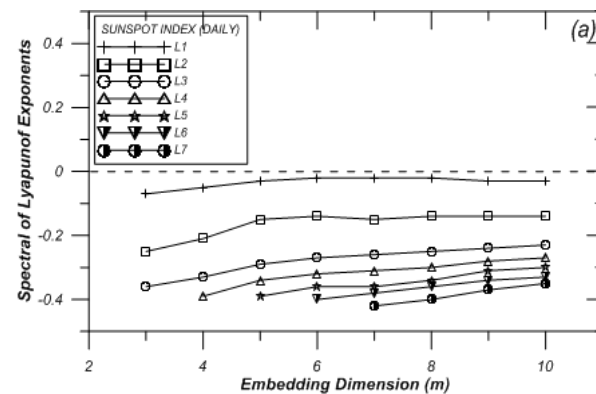
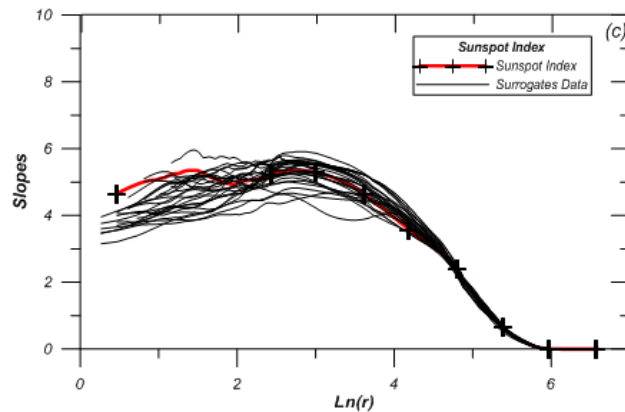
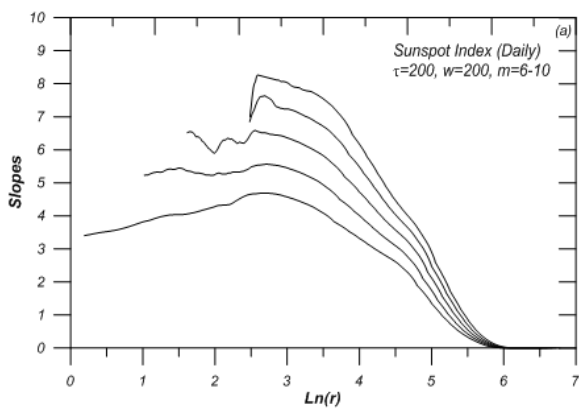
Sunspot Index (Original – SVD Component)



Sunspot Index (Correlation Dimension & Lyapunov Exponent)

Correlation Dimension

Lyapunov Spectrum Exponents

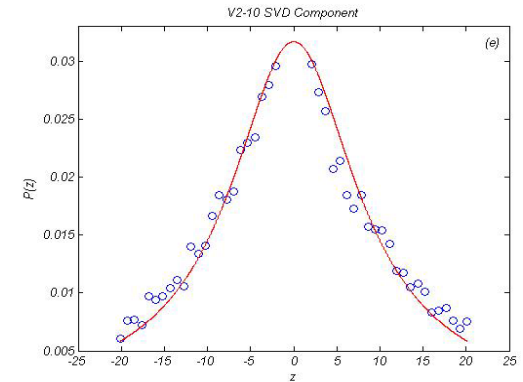
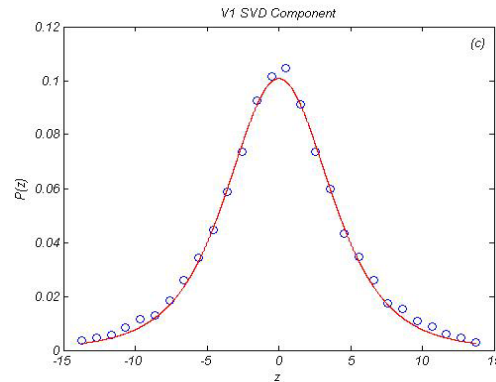
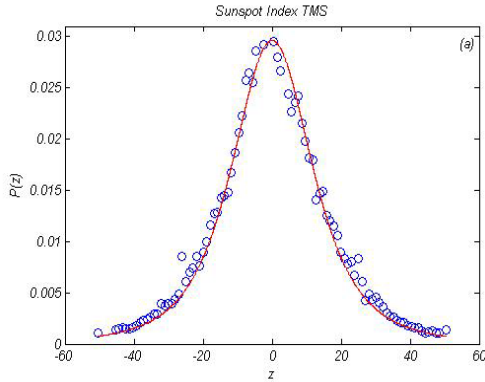


Sunspot Index (Tsallis Statistics)

Sunspot TMS - $q_{stat}=1.53 \pm 0.04$

V1 SVD Component - $q_{stat}=1.40 \pm 0.08$

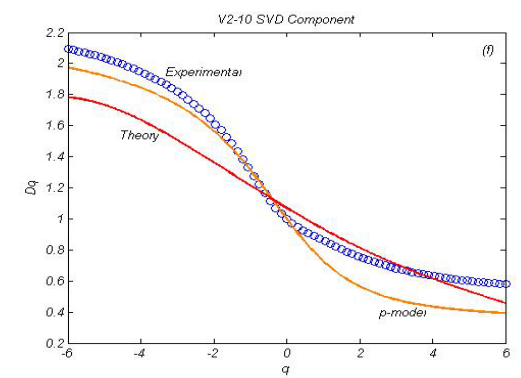
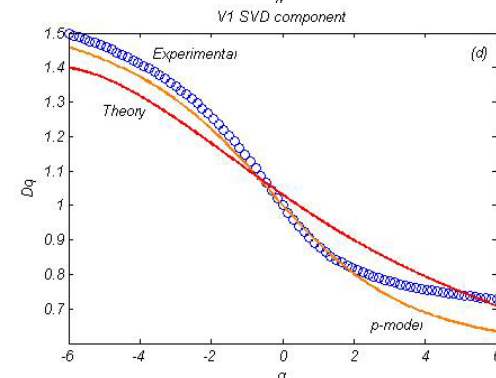
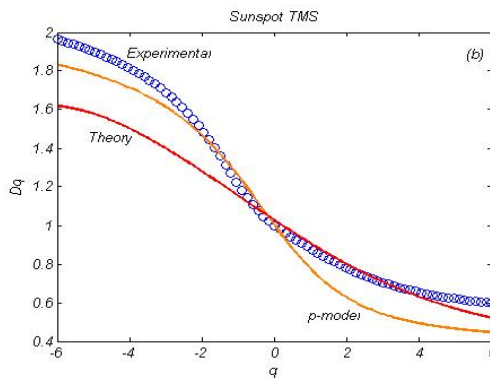
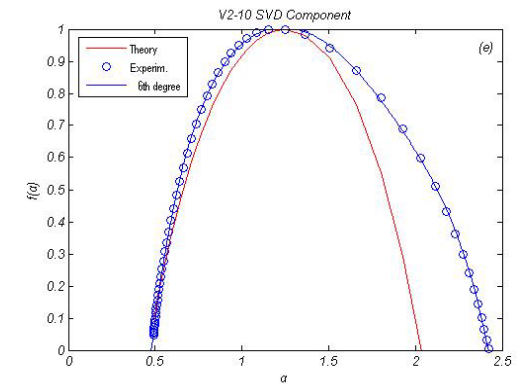
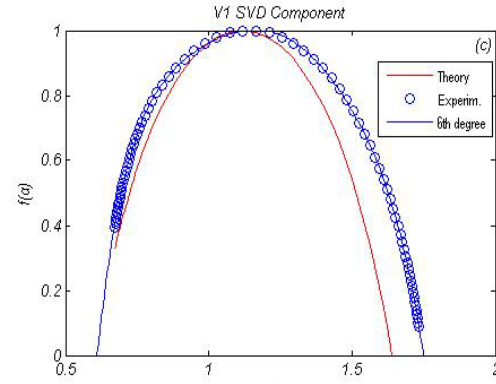
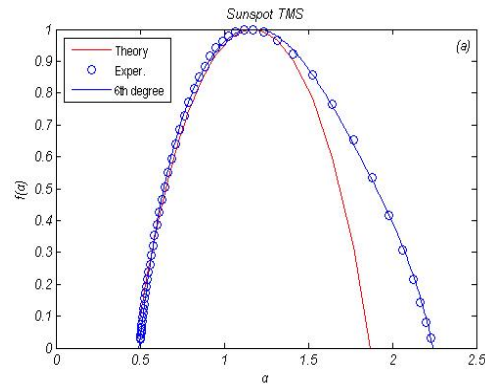
V2-10 SVD Component - $q_{stat}=2.12 \pm 0.20$



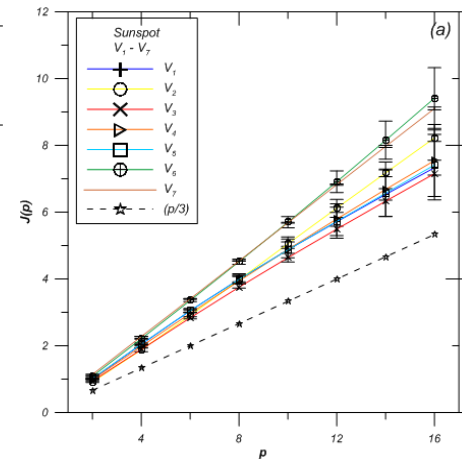
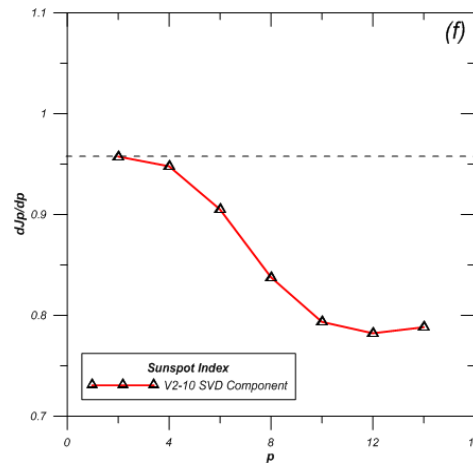
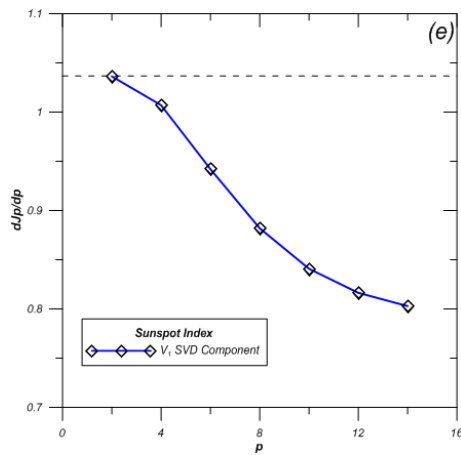
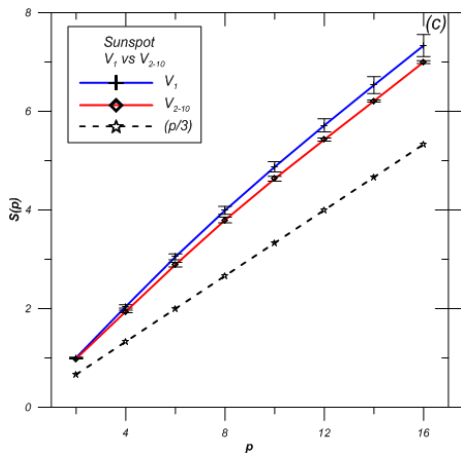
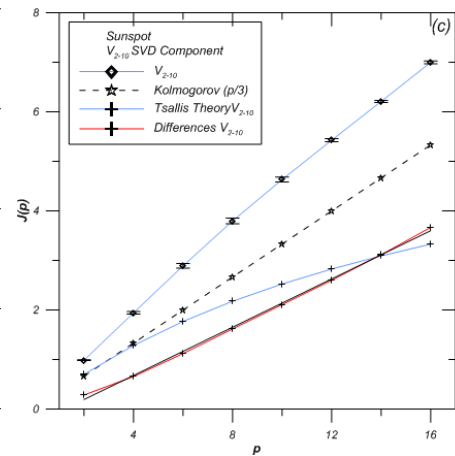
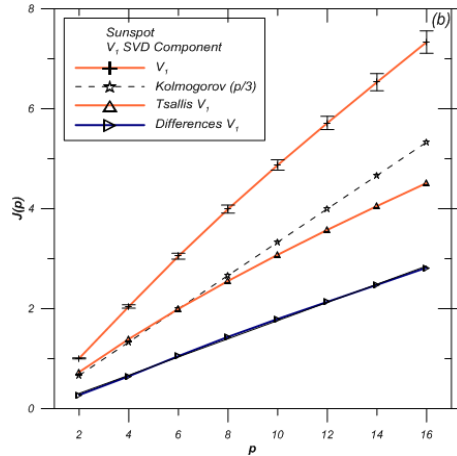
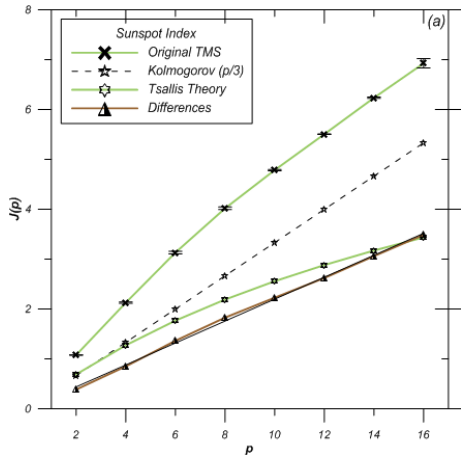
**Sunspot TMS TMS - $q_{sen}=0.368$
 $\Delta\alpha=1.752$**

**V1 SVD Component - $q_{sen}=-0.055$
 $\Delta\alpha=1.133$**

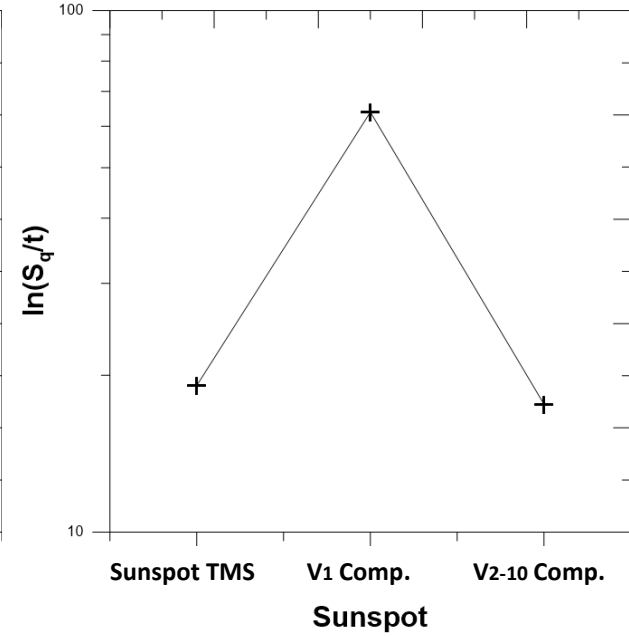
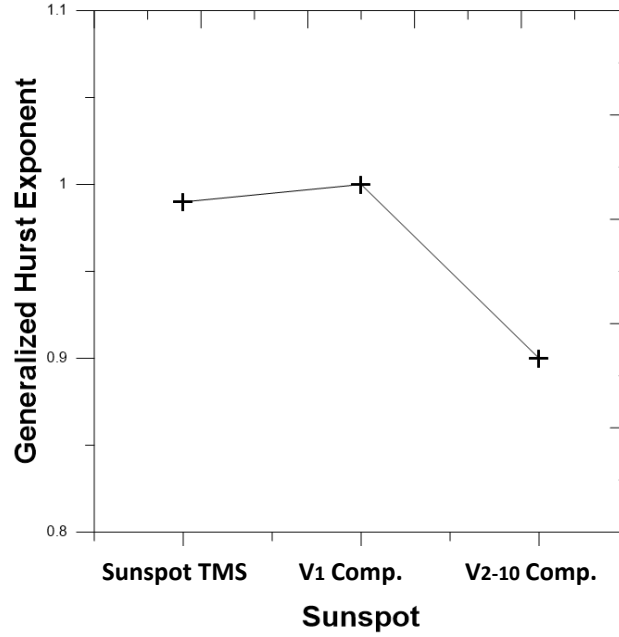
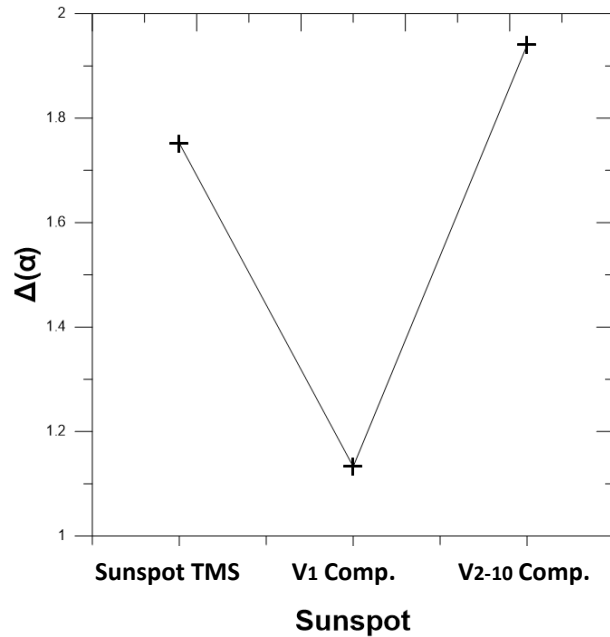
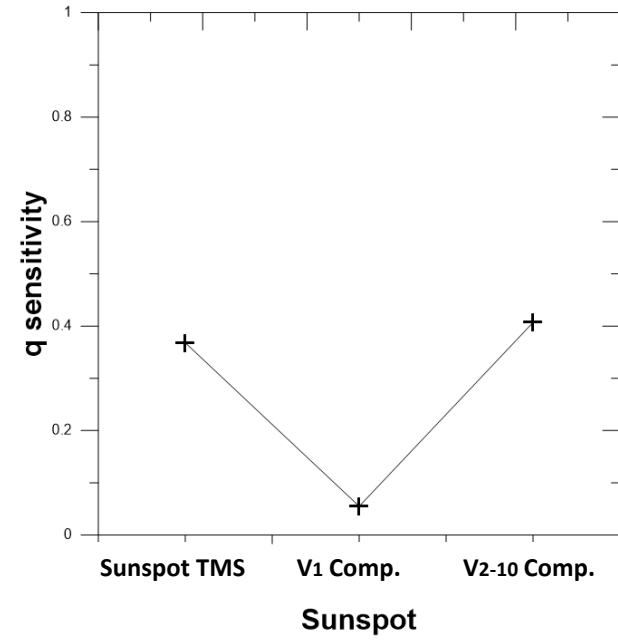
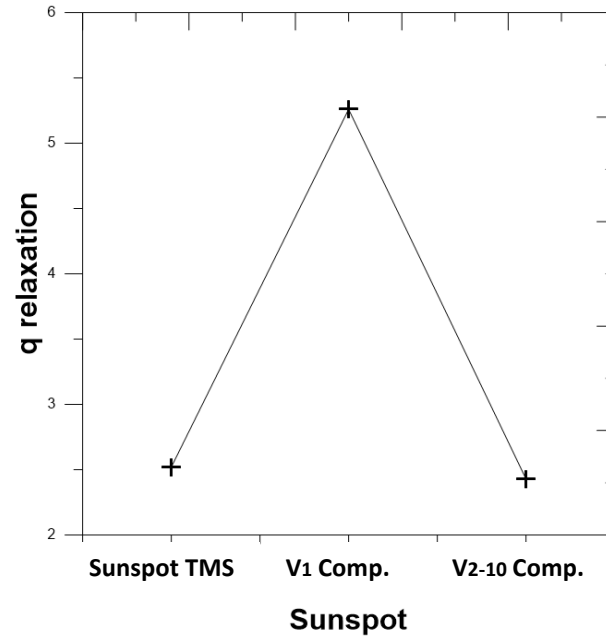
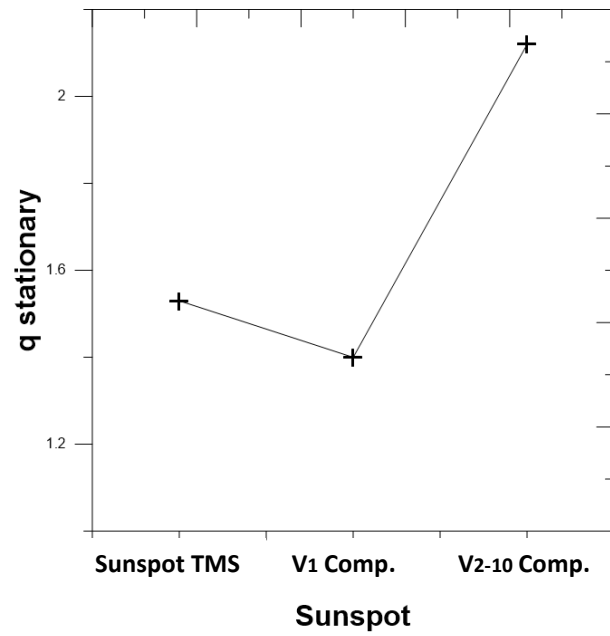
**V2-10 SVD Component - $q_{sen}=0.407$
 $\Delta\alpha=1.940$**



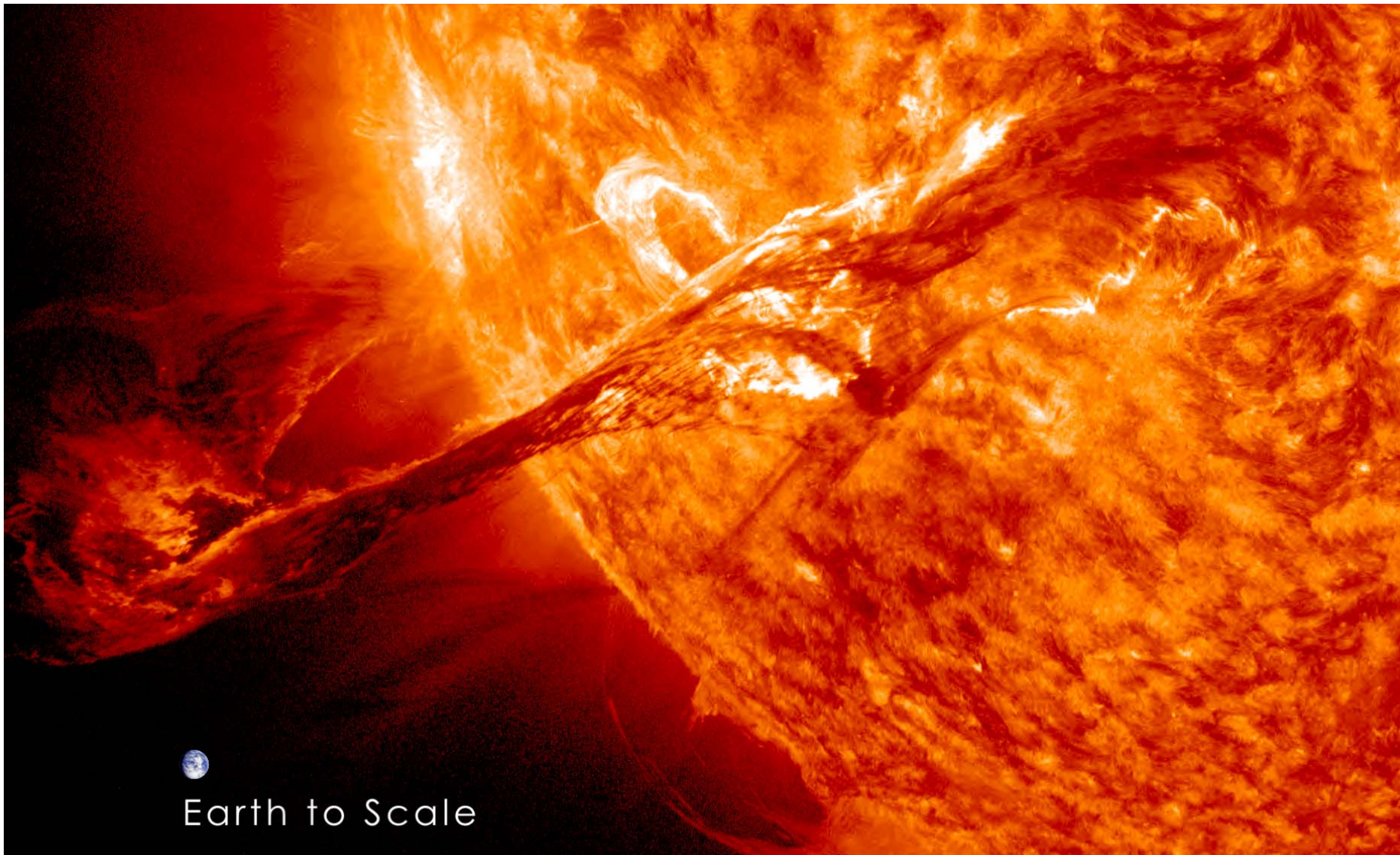
Sunspot Index (Turbulence Analysis)



Sunspot Index (Parameters, Generalized Hurst Exponent & Entropy Production)

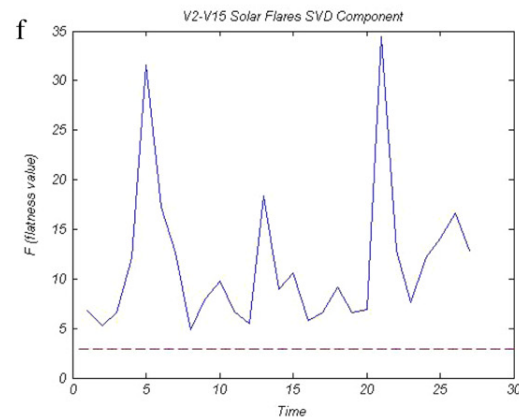
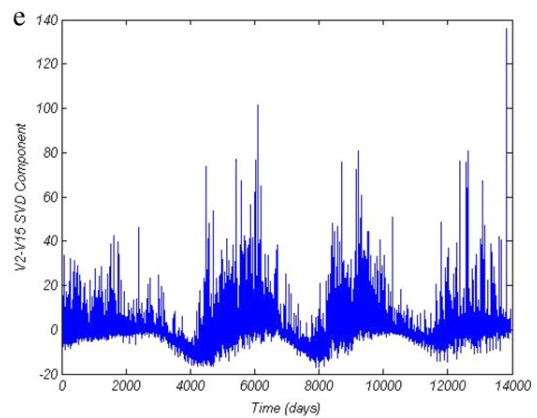
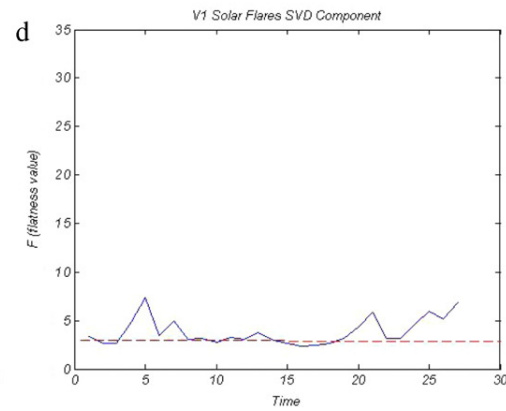
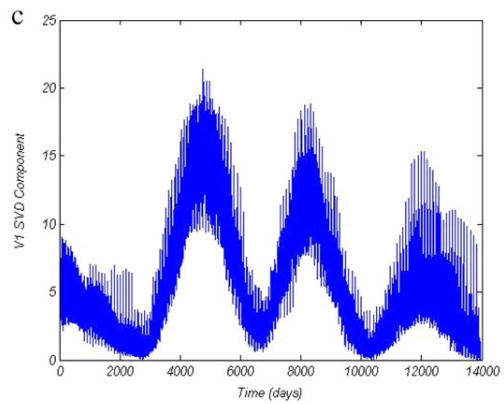
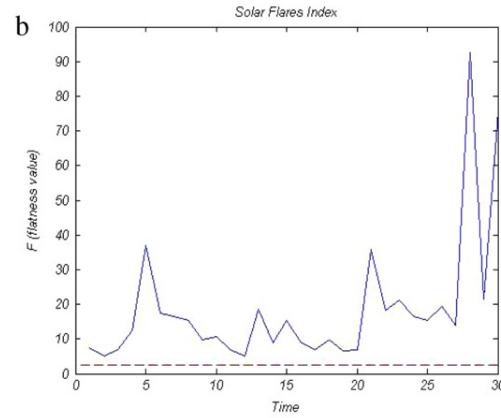
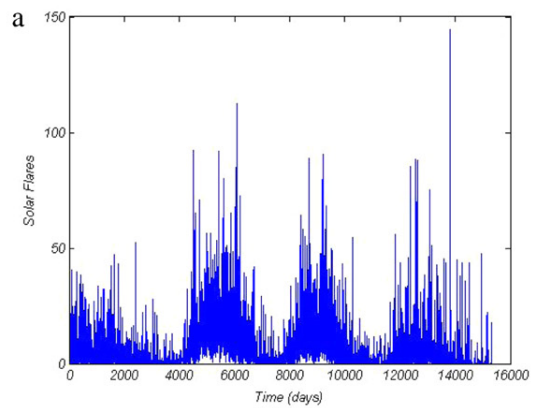


Solar Flares



Earth to Scale

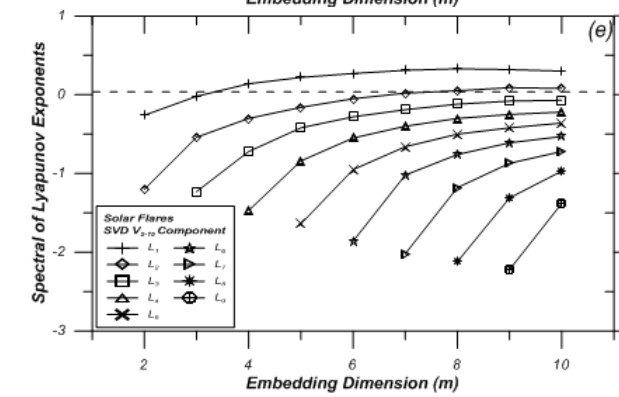
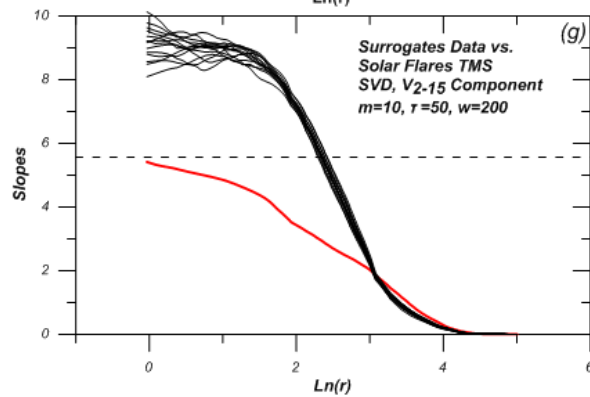
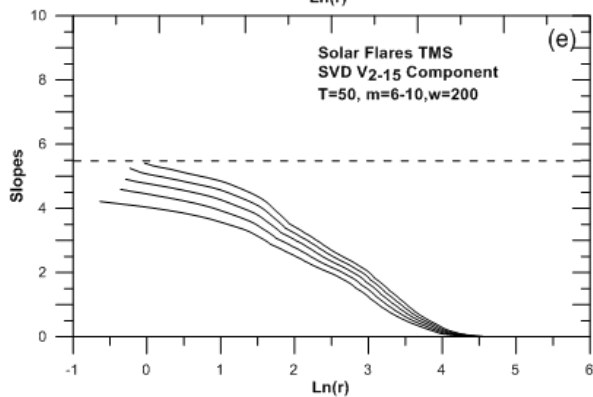
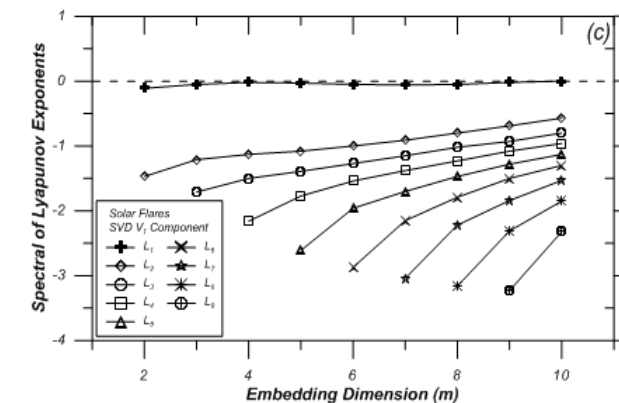
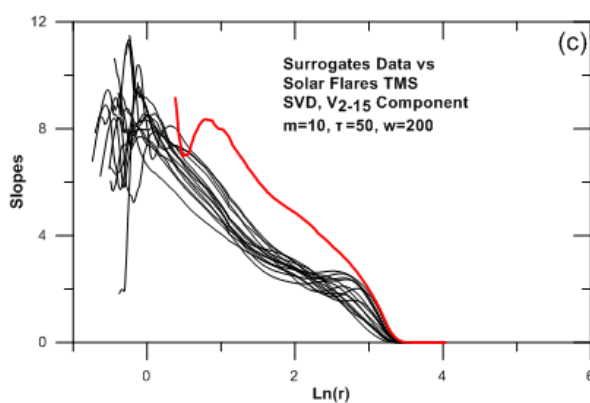
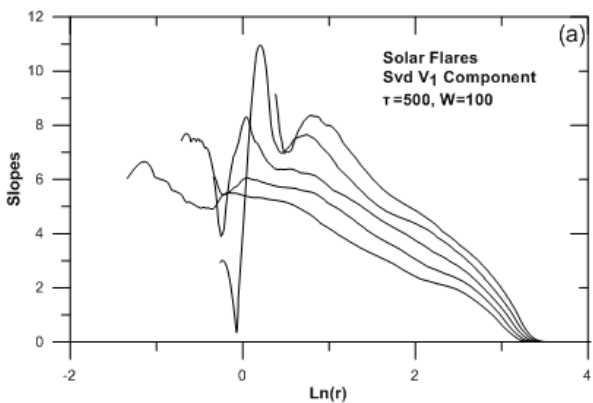
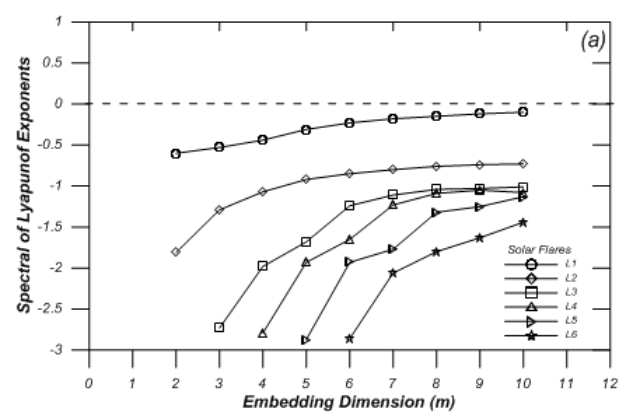
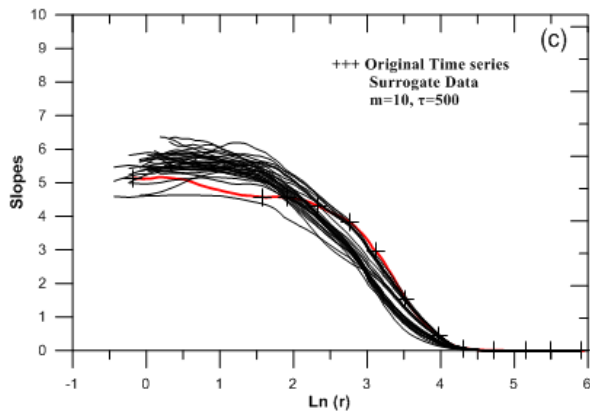
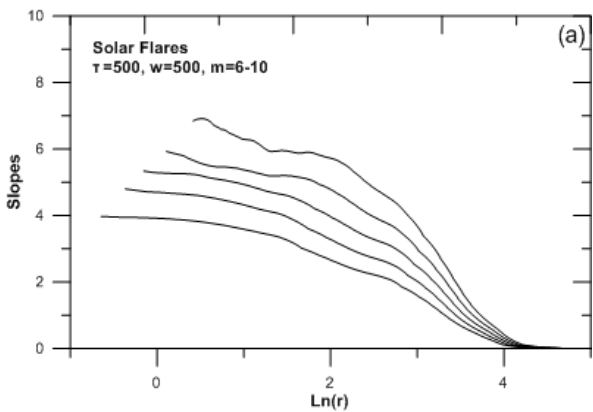
Solar Flares Index (Original - SVD Components)



Solar Flares Index (Correlation Dimension & Lyapunov Exponent)

Correlation Dimension

Lyapunov Spectrum Exponents

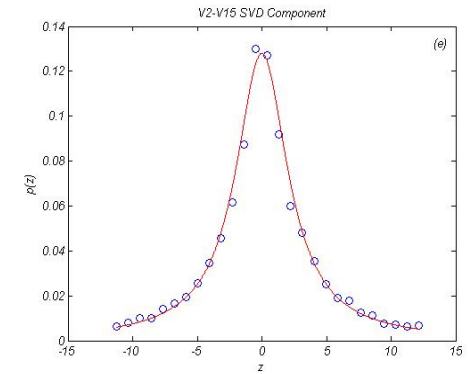
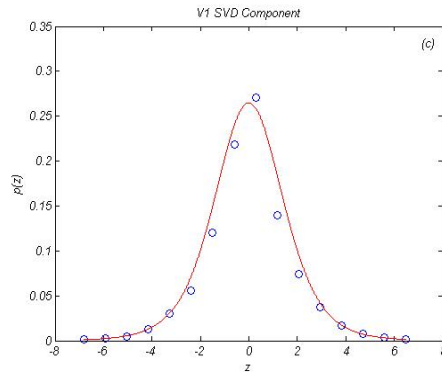
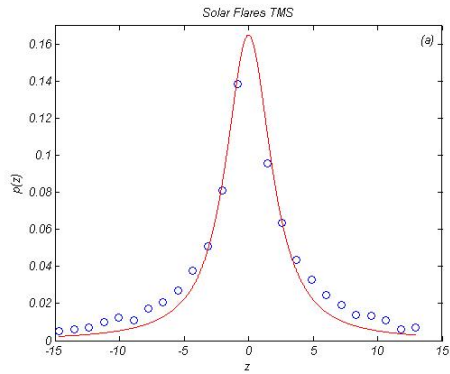


Solar Flares Index (Tsallis Statistics)

Solar Flare TMS - $q_{stat}=1.87 \pm 0.05$

V1 SVD Component - $q_{stat}=1.28 \pm 0.04$

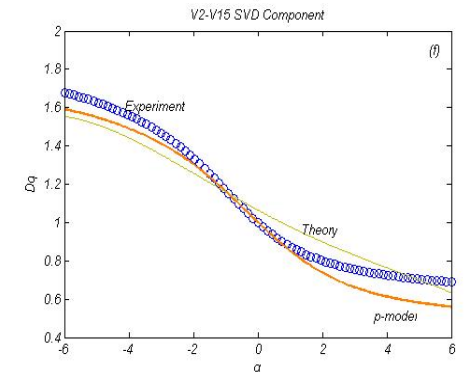
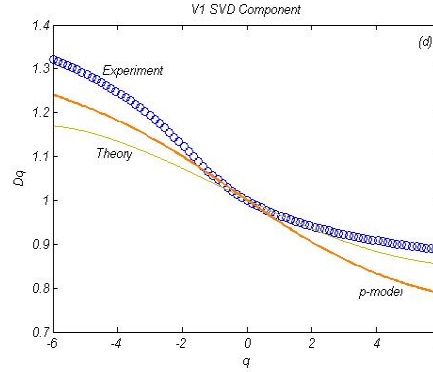
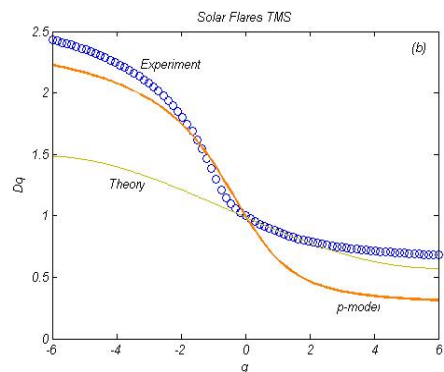
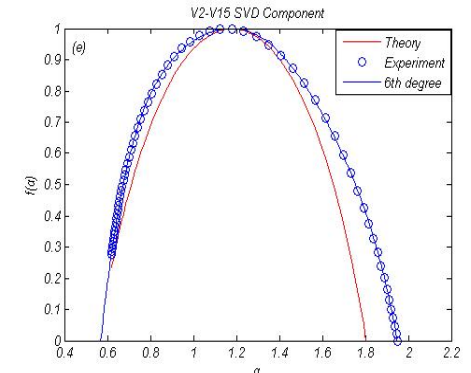
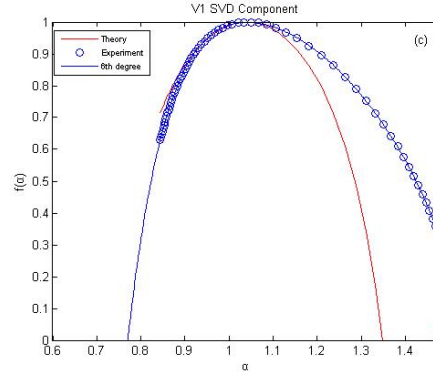
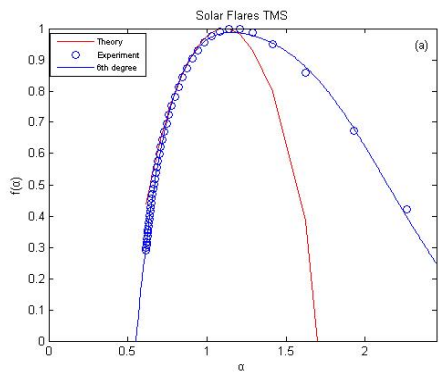
V2-15 SVD Component - $q_{stat}=2.02 \pm 0.15$



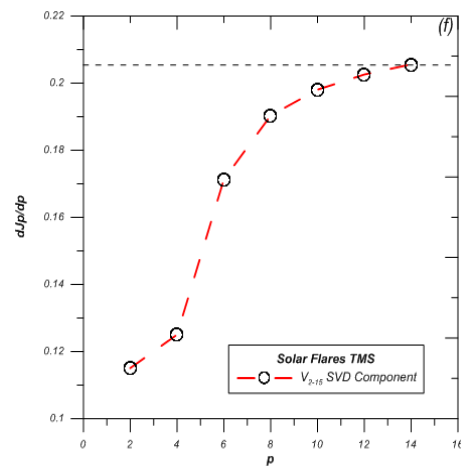
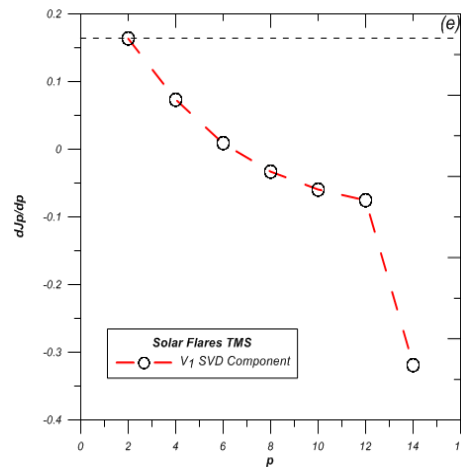
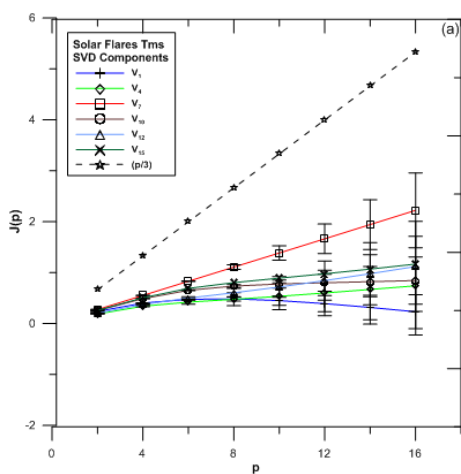
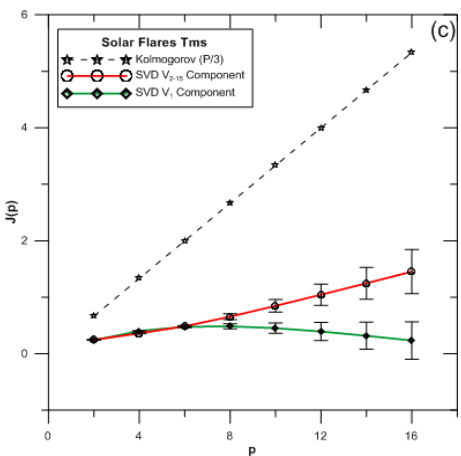
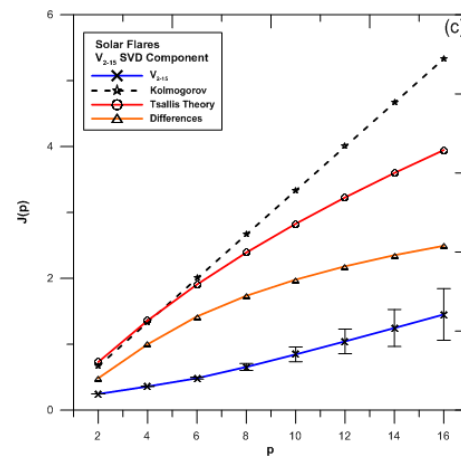
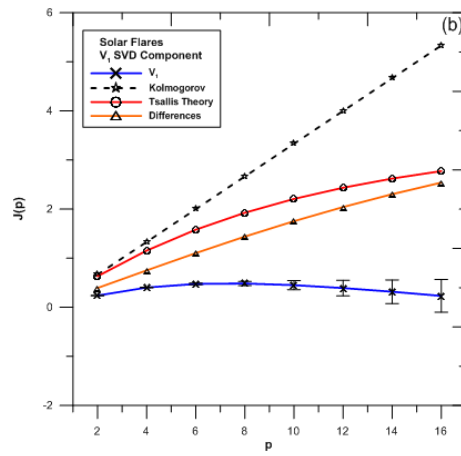
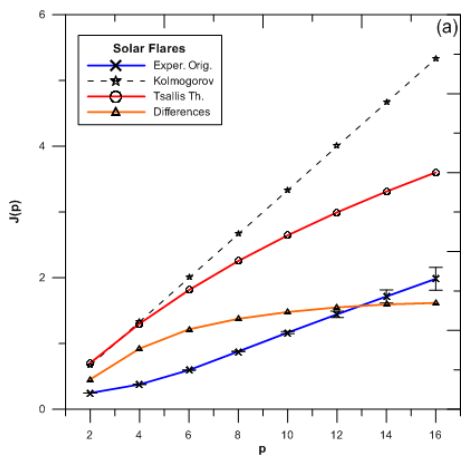
**Solar Flare TMS - $q_{sen}=0.308$
 $\Delta\alpha= 2.15$**

**V1 SVD Component - $q_{sen}=-0.540$
 $\Delta\alpha= 0.77$**

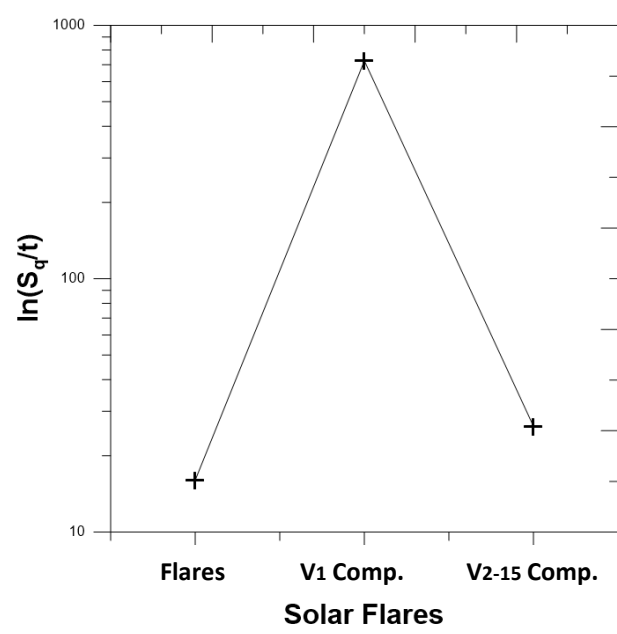
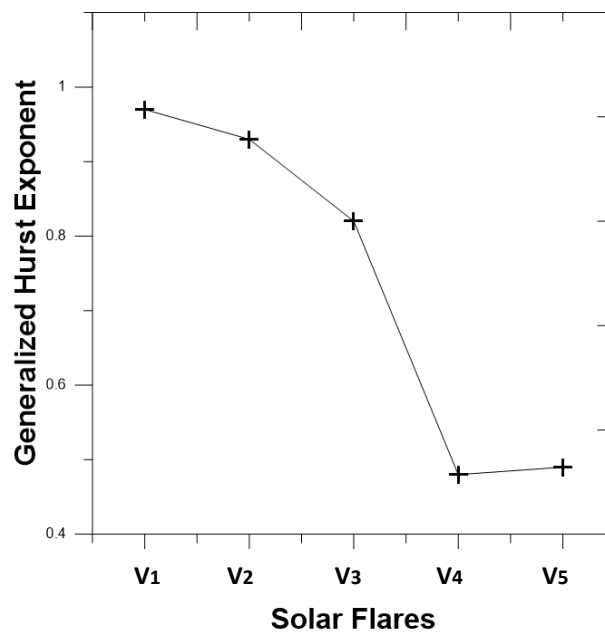
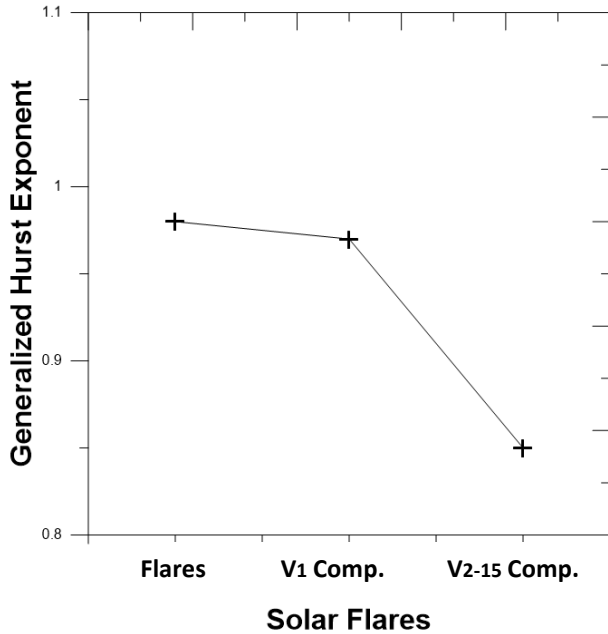
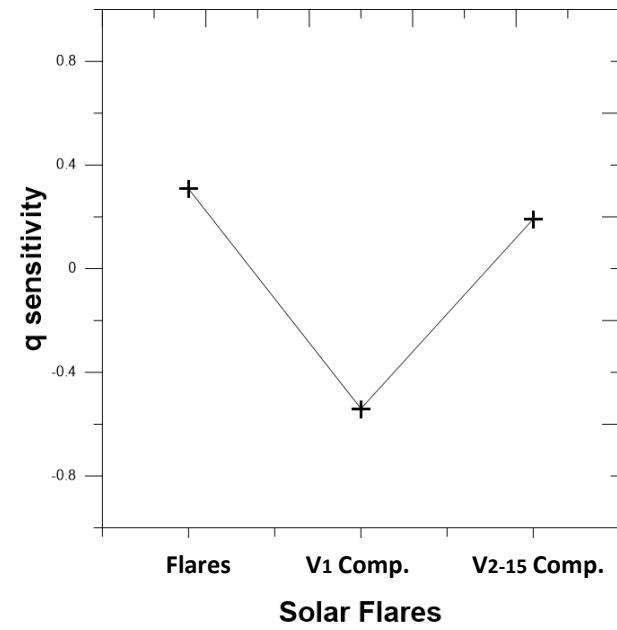
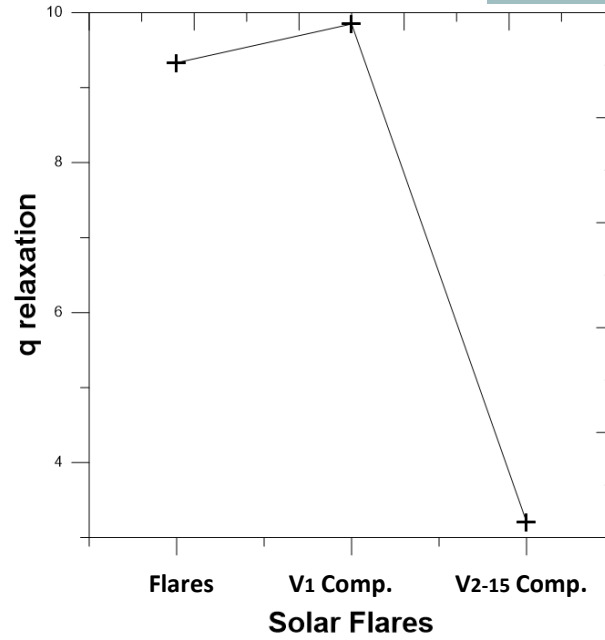
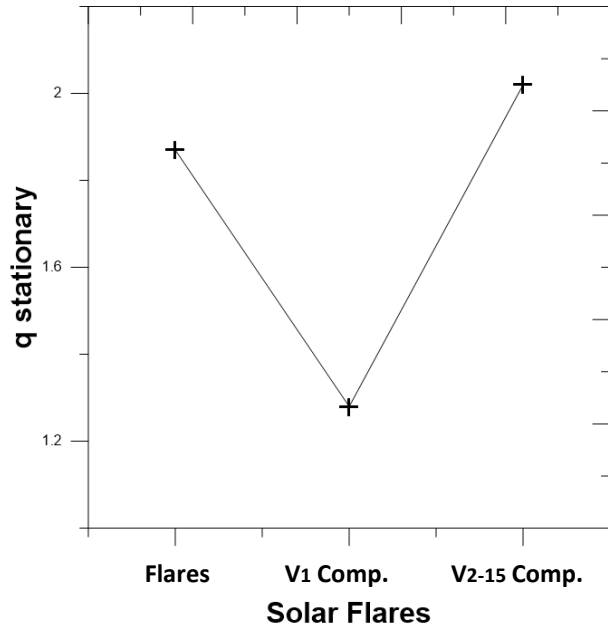
**V2-15 SVD Component - $q_{sen}=0.192$
 $\Delta\alpha= 1.37$**



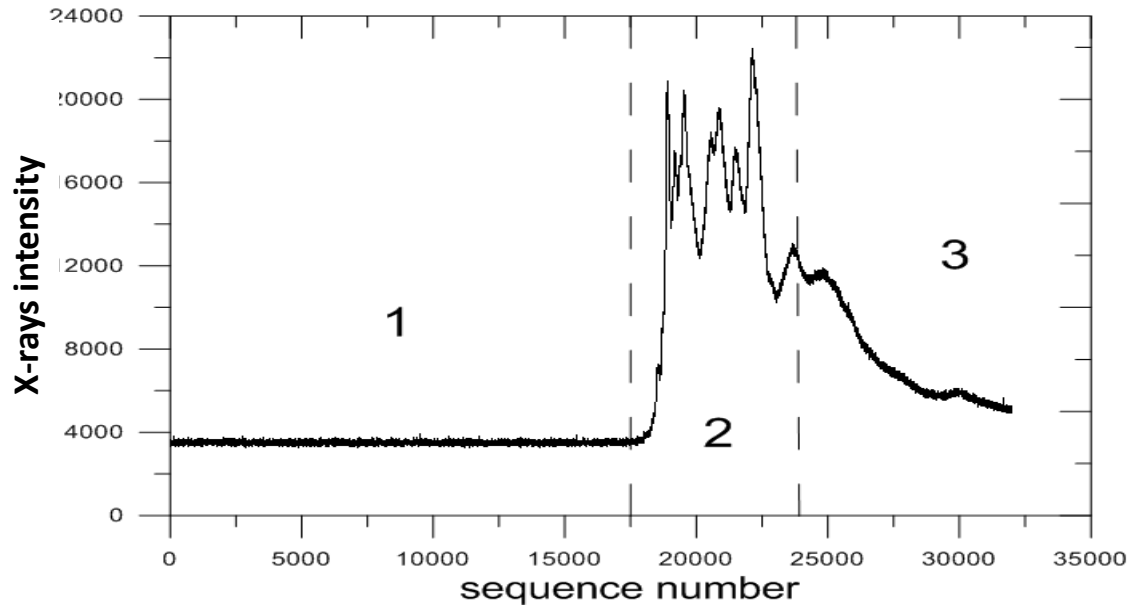
Solar Flares Index (Turbulence Analysis)



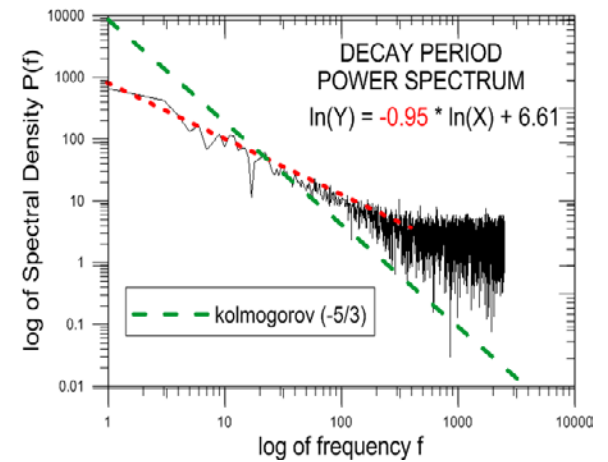
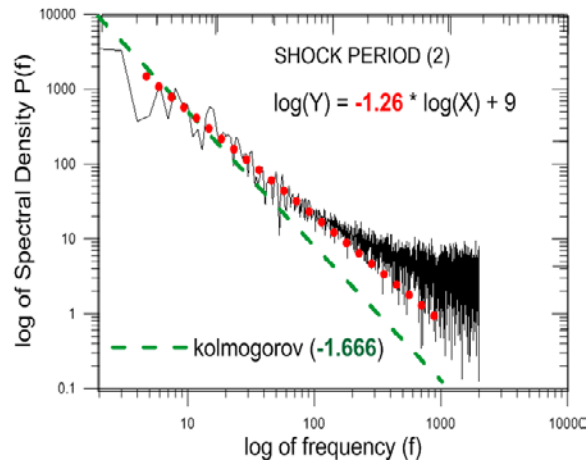
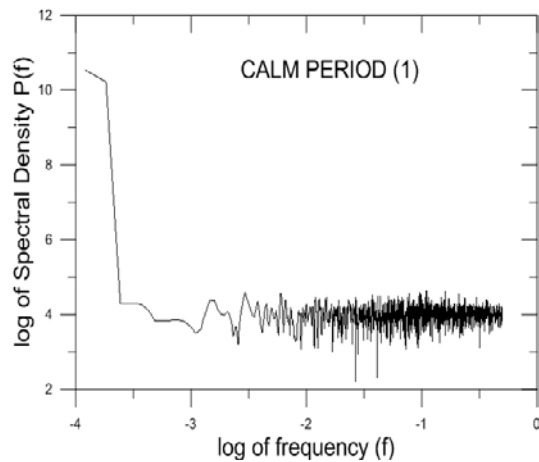
Solar Flares Index (Parameters, Generalized Hurst Exponent & Entropy Production)



Solar Flares Event 1

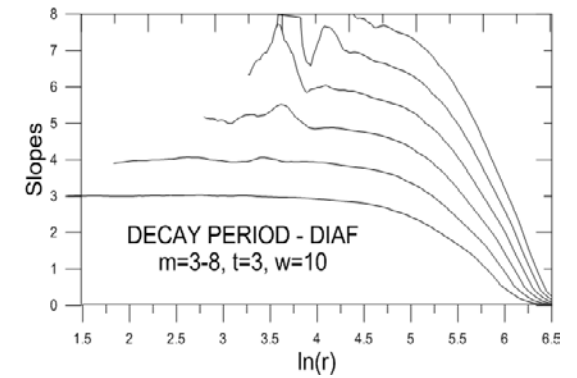
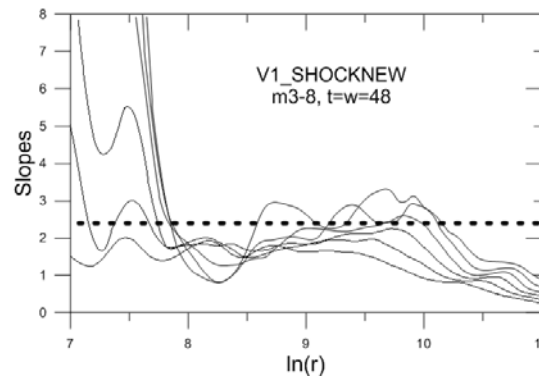
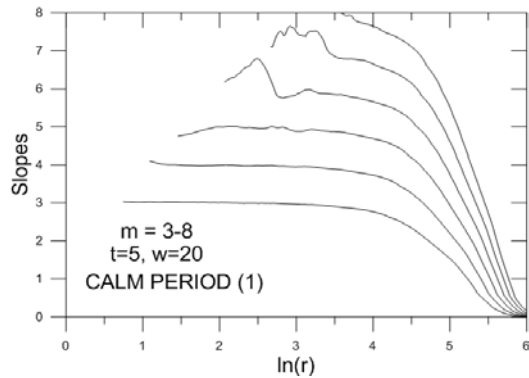


Power Spectrum



Solar Flares Event 1 (Chaotic Analysis)

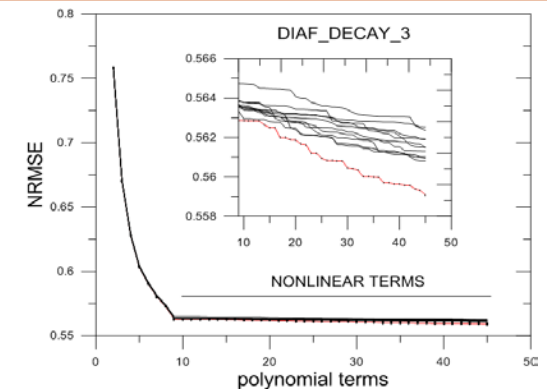
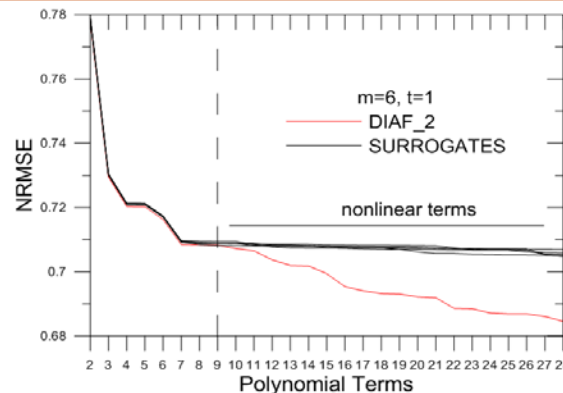
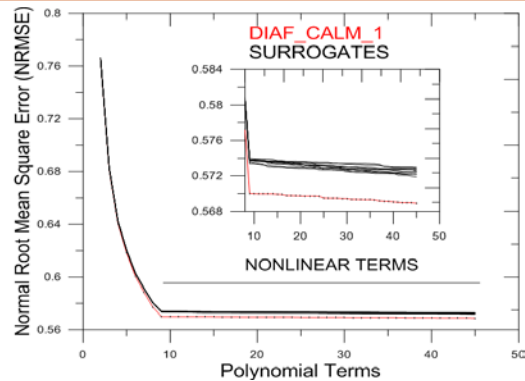
Correlation Dimension



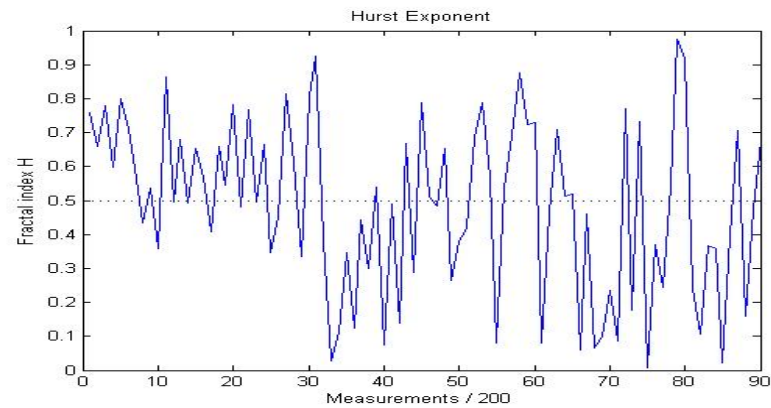
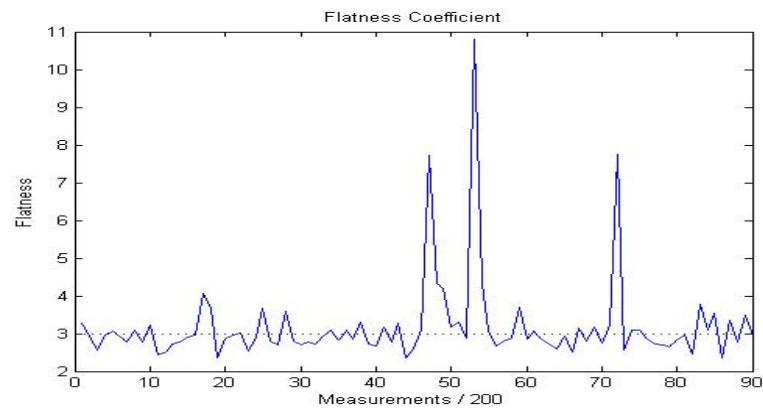
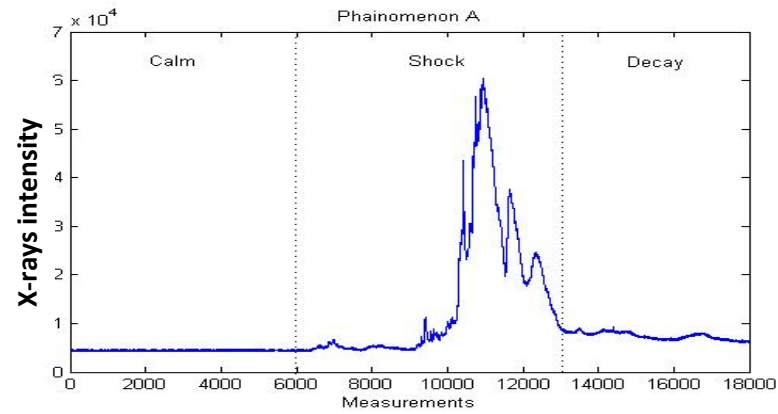
Lyapunov Spectrum Exponents

	CALM (m=10)	SHOCK (m=7)	DECAY(m=10)
L1 maximum	-0.3	0.02 > 0	-0.01
L2	-0.32	-0.04	-1.19
L3	-0.34	-0.04	-2.42
L4	-0.37	-0.05	-2.48
L5	-0.4	-0.35	-2.57

Volterra – Weiner Model

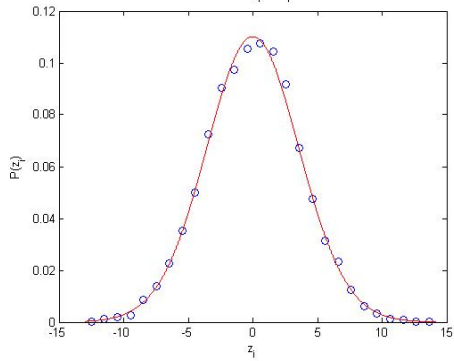


Solar Flares Event 2 (Statistical Analysis)

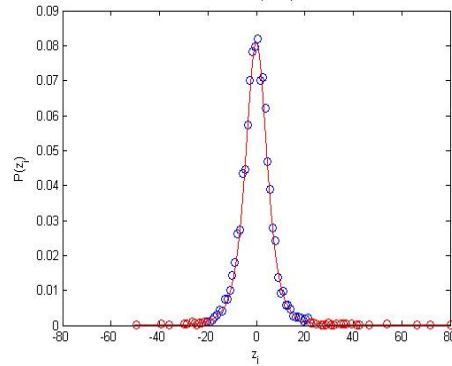


Solar Flares Event 2 (Tsallis Statistics)

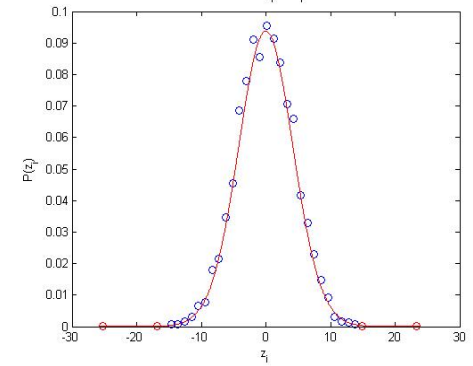
Calm- $q_{stat}=1.04\pm 0.02$



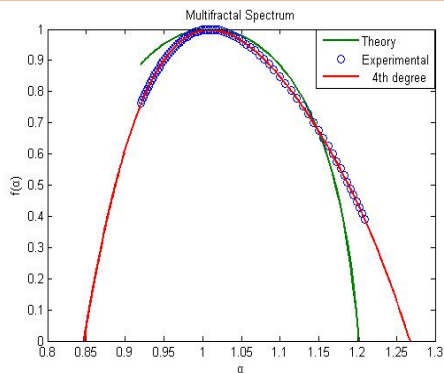
Shock- $q_{stat}=1.43\pm 0.05$



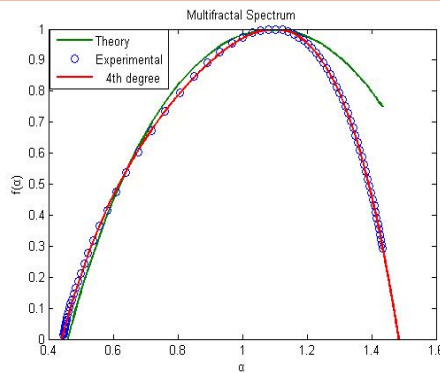
Decay- $q_{stat}=1.08\pm 0.02$



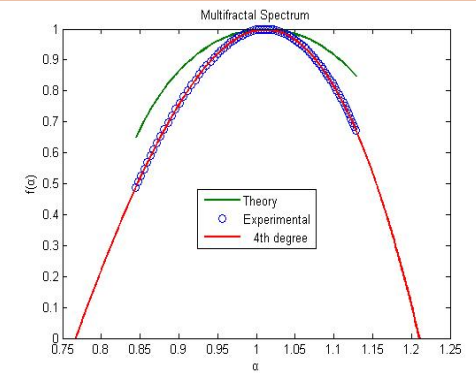
**Calm - $q_{sen}=-1.547$
 $\Delta(\alpha) = 0.421$**



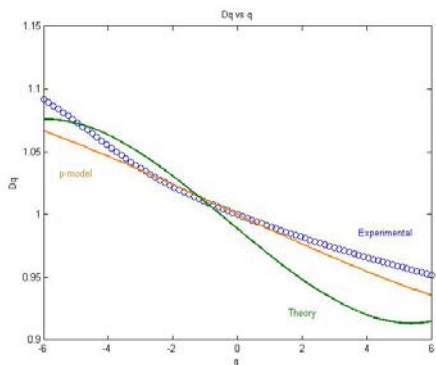
**Shock - $q_{sen}=0.376$
 $\Delta(\alpha) = 1.045$**



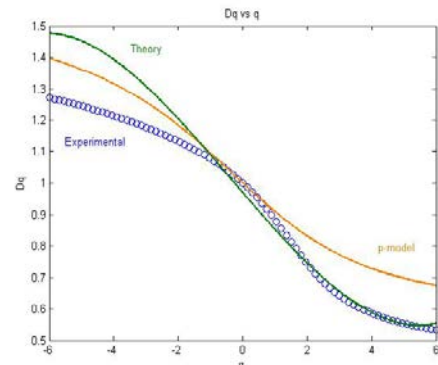
**Decay - $q_{sen}=-1.089$
 $\Delta(\alpha) = 0.441$**



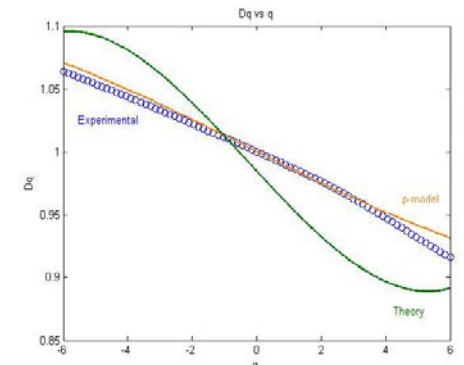
Calm - $\Delta(Dq)=0.141$



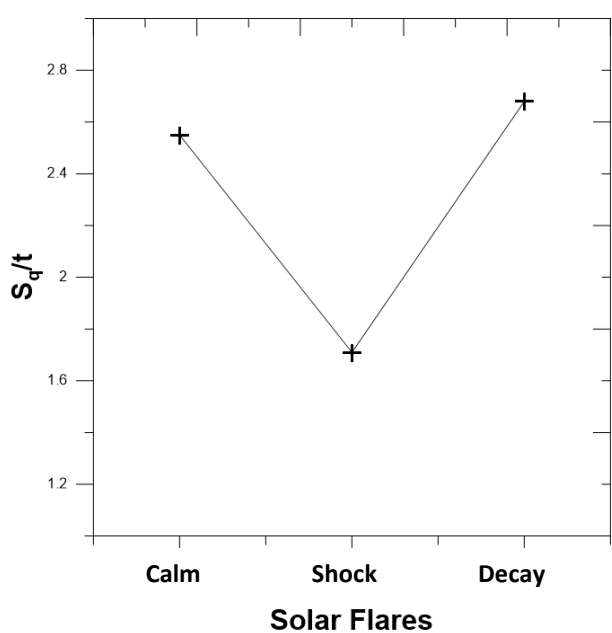
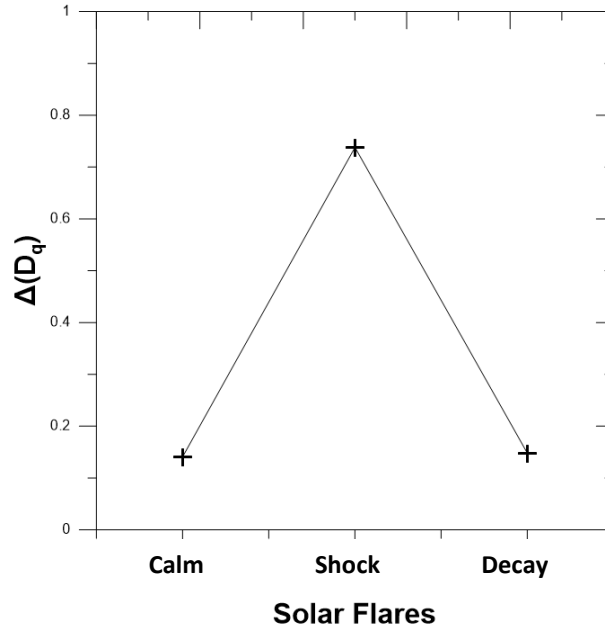
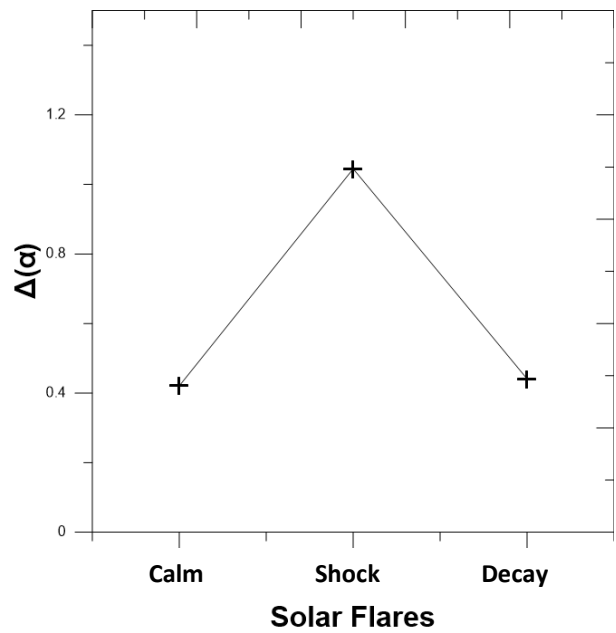
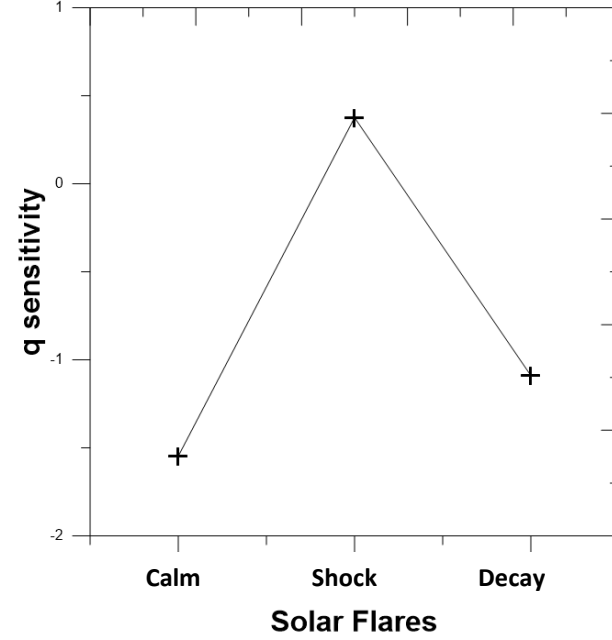
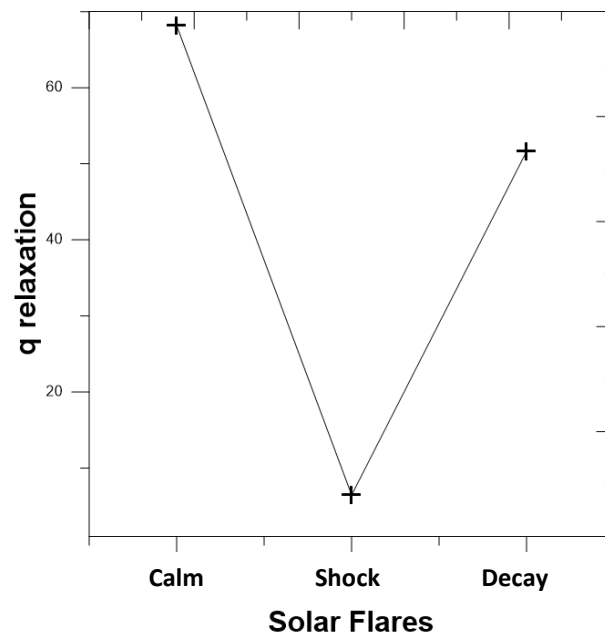
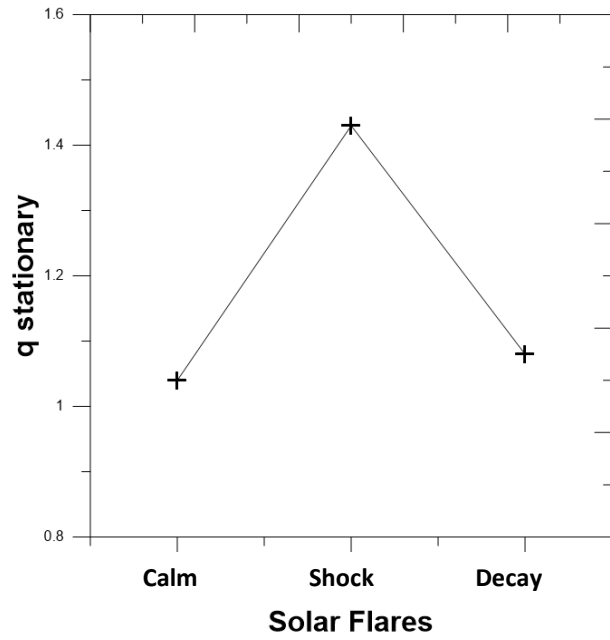
Shock - $\Delta(Dq)=0.738$



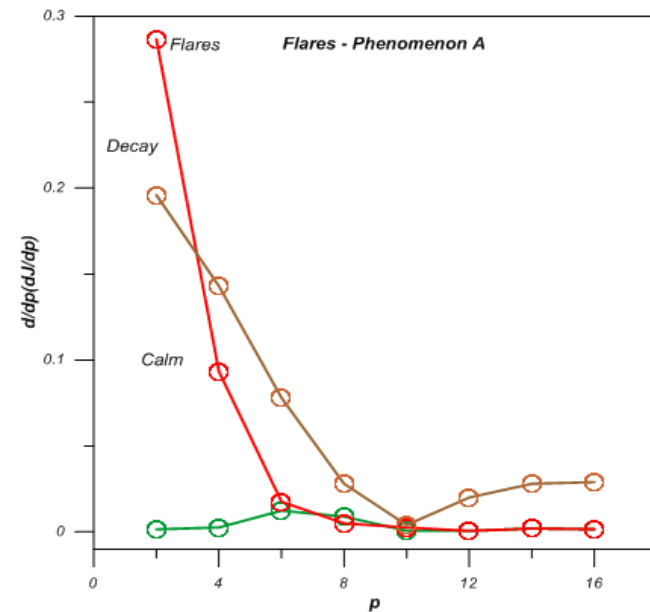
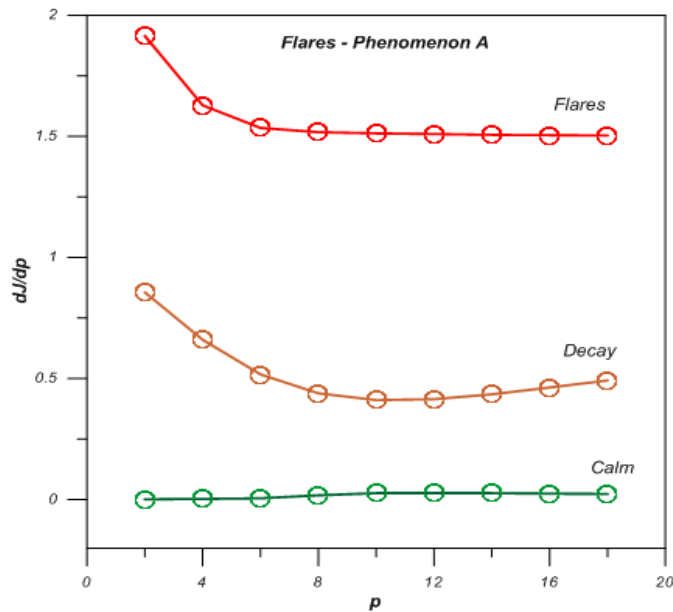
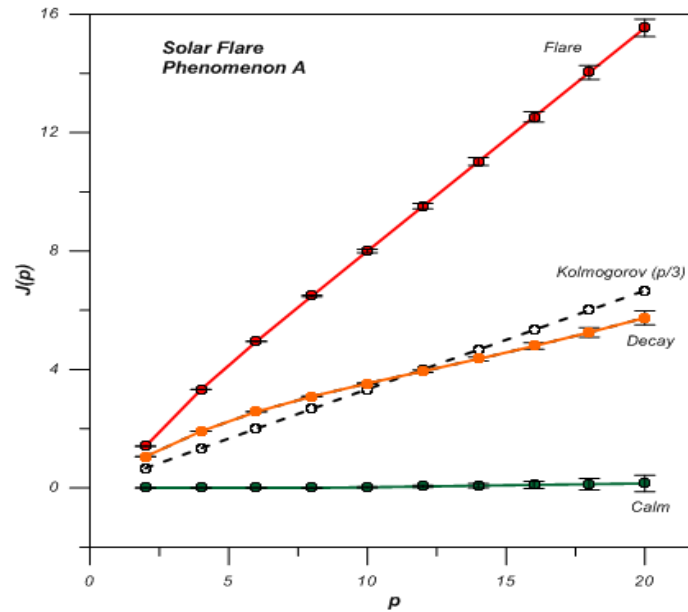
Decay - $\Delta(Dq)=0.148$



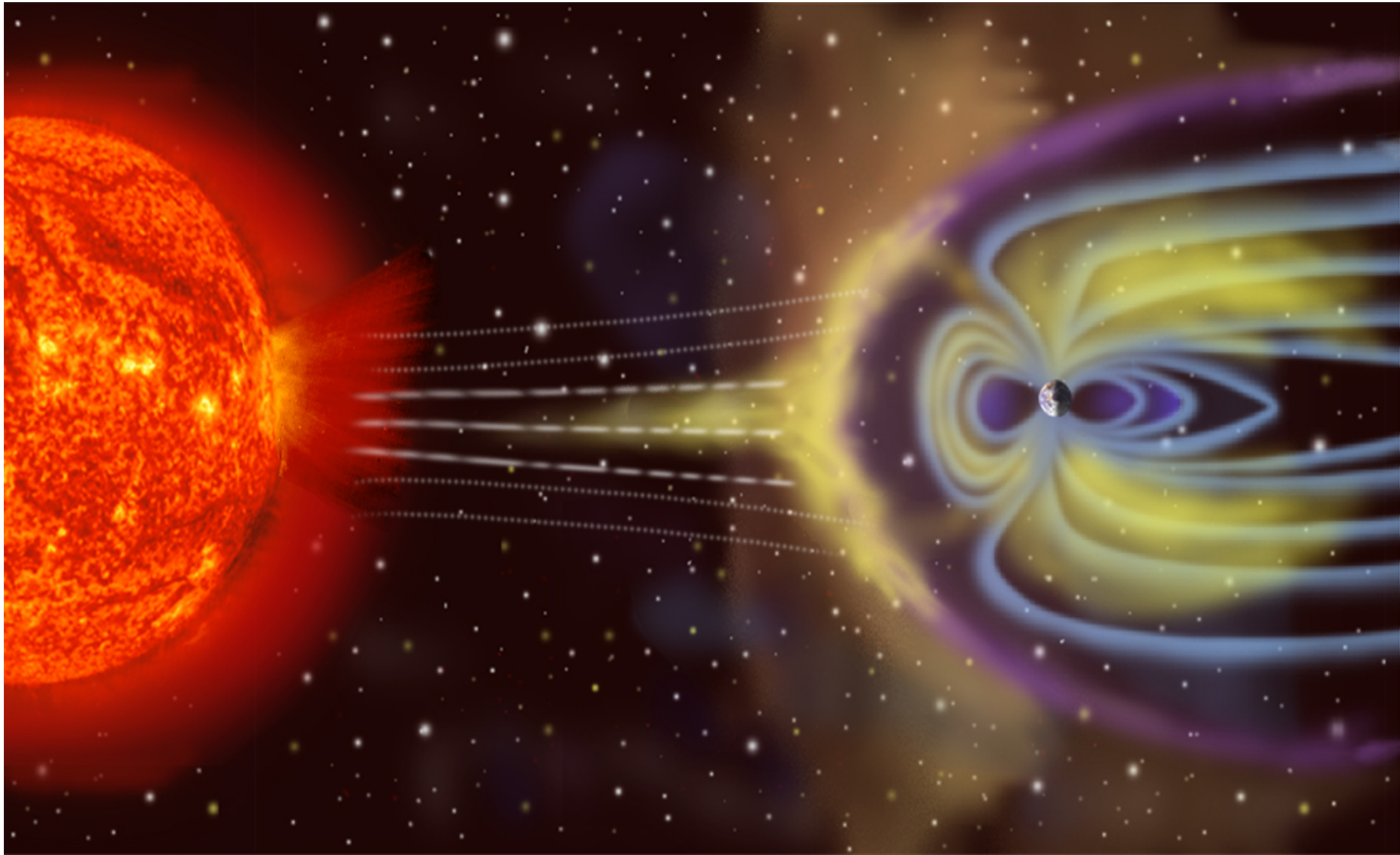
Solar Flares Event 2 (Parameters)



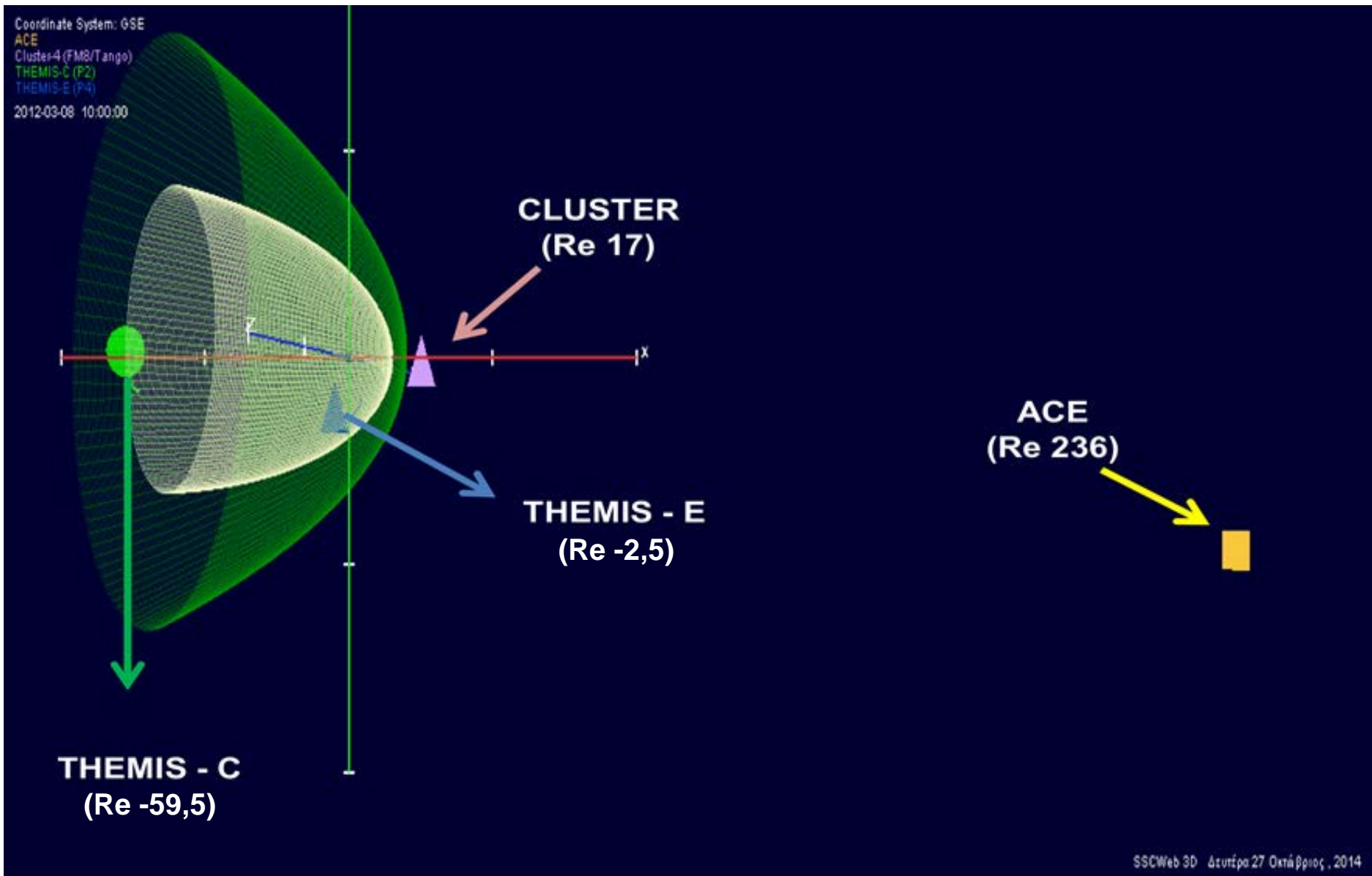
Solar Flares Event 2 (Turbulence Analysis)



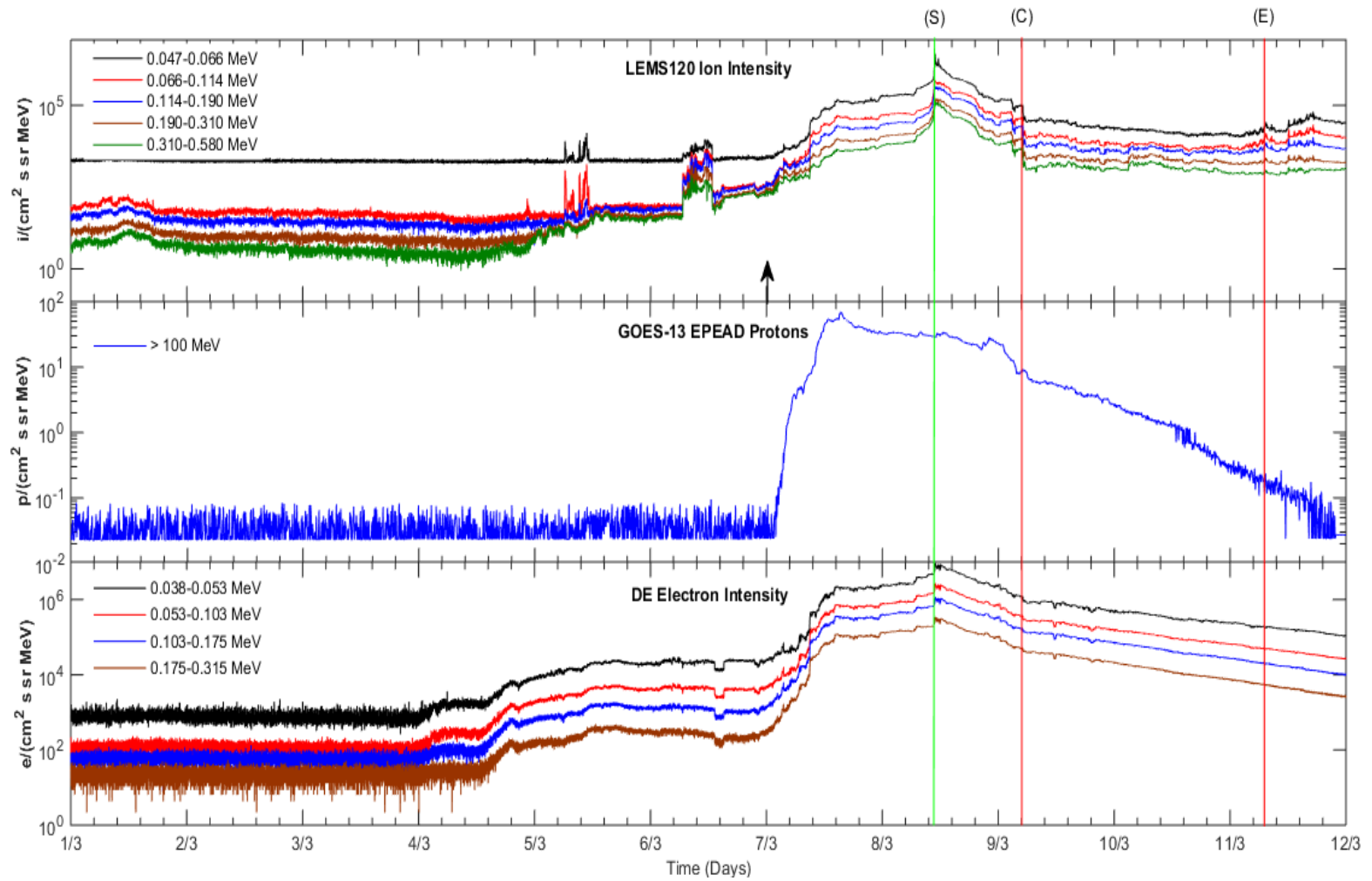
Solar Energetic Particle Interplanetary Space



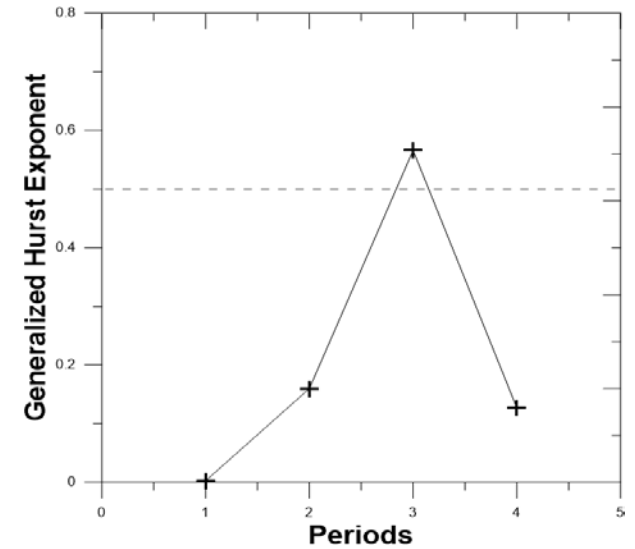
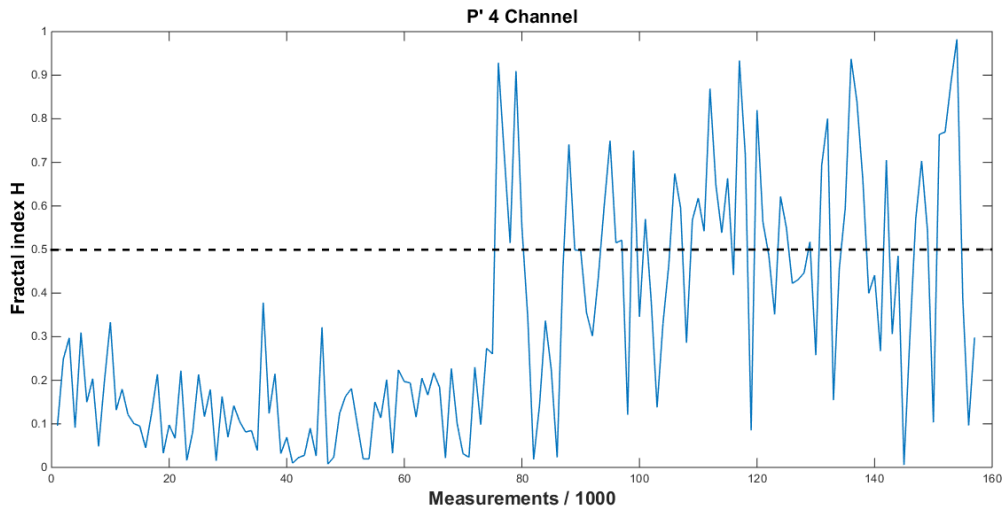
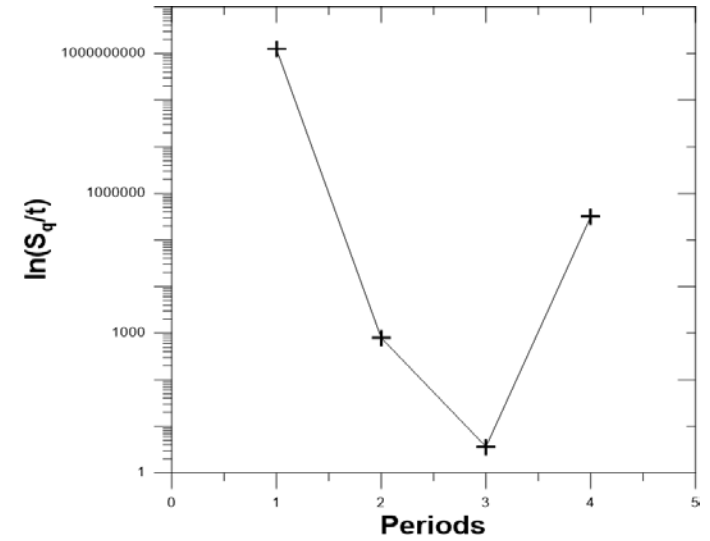
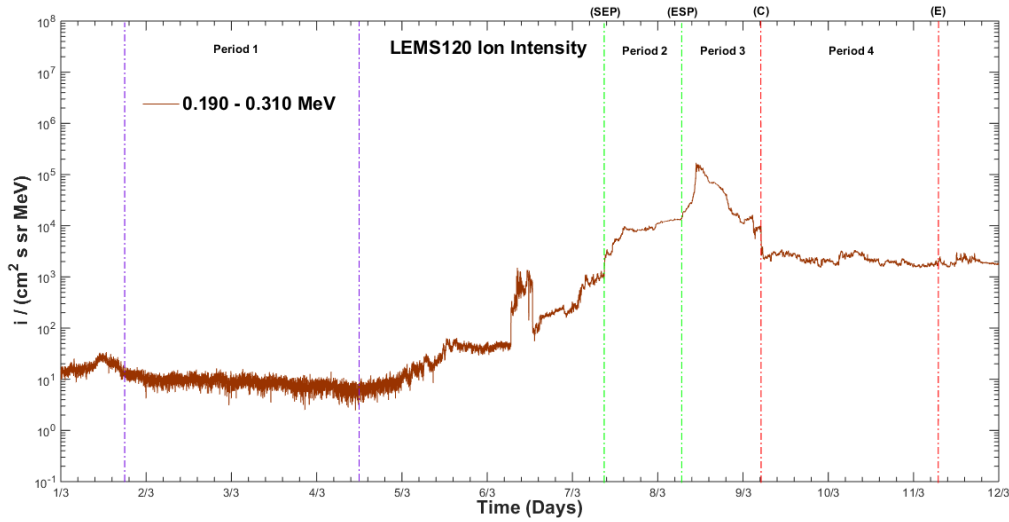
Interplanetary Observations (CME 7 March 2012)



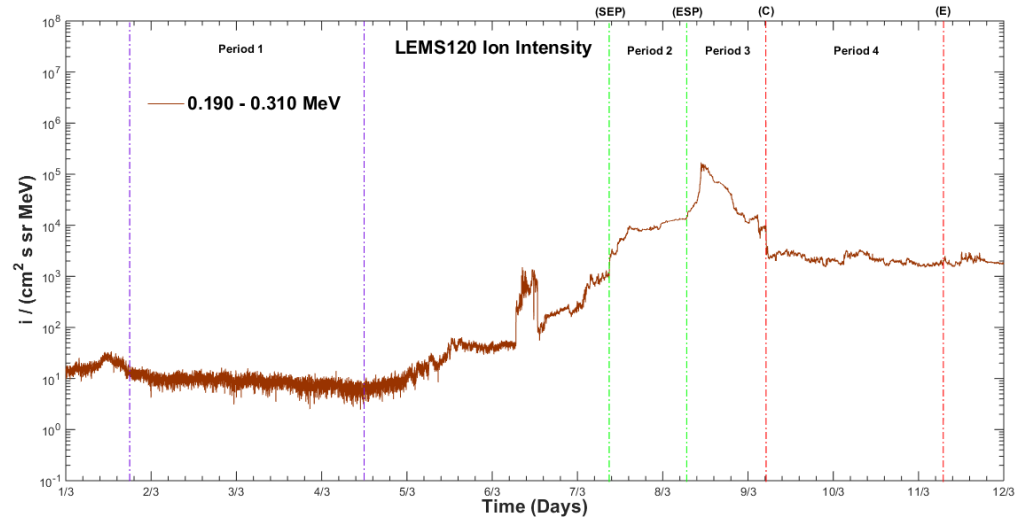
Solar Energetic Particles



Solar Energetic Particle (Generalized Hurst Exponent & Entropy Production)



Solar Energetic Particle (Tsallis Statistics)

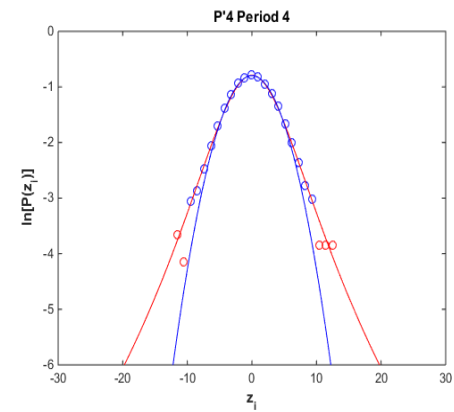
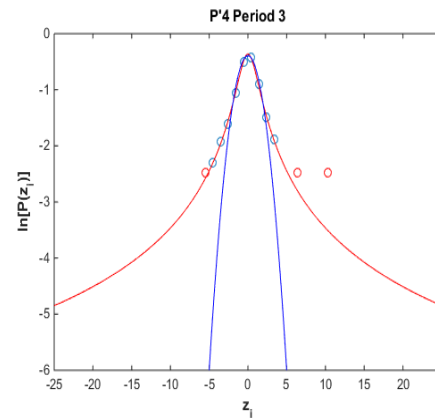
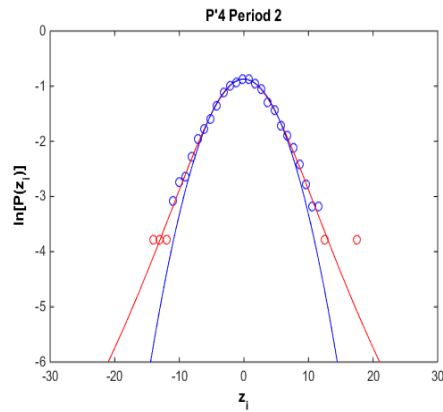
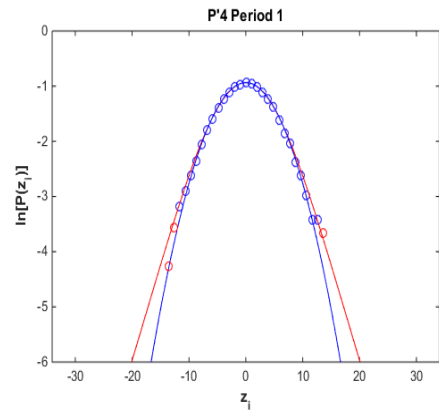


Period 1 - $q_{\text{stat}}=1.08 \pm 0.03$

Period 2 - $q_{\text{stat}}=1.18 \pm 0.03$

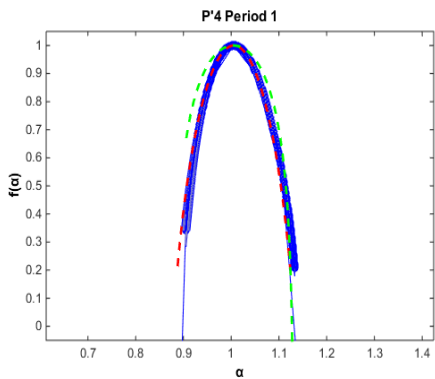
Period 3 - $q_{\text{stat}}=1.57 \pm 0.08$

Period 4 - $q_{\text{stat}}=1.11 \pm 0.05$

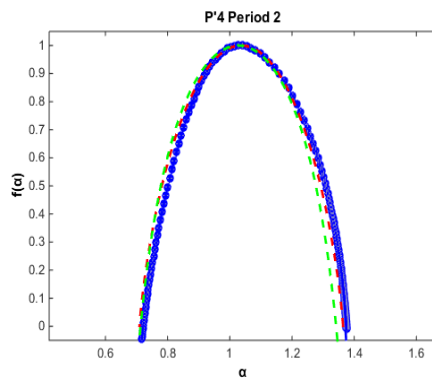


Solar Energetic Particle (Multifractal Spectrum)

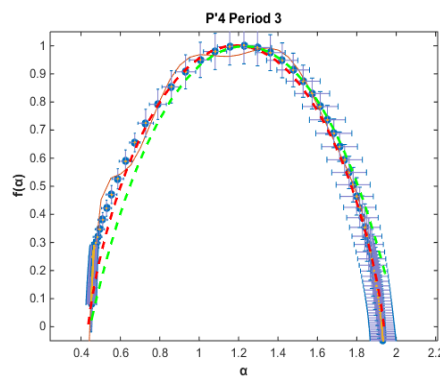
**Period 1 - $q_{sen}=-3.049$
 $\Delta\alpha=0.251$**



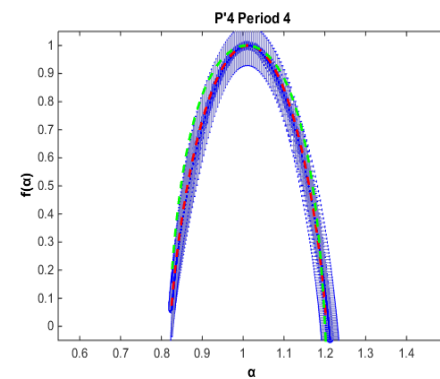
**Period 2 - $q_{sen}=-0.521$
 $\Delta\alpha=0.653$**



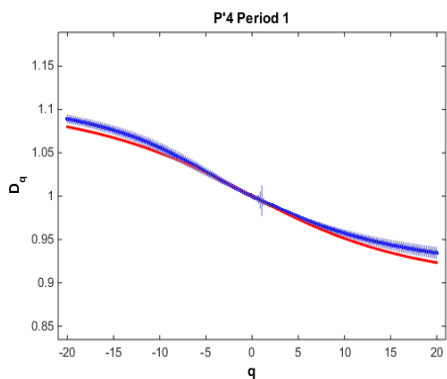
**Period 3 - $q_{sen}=0.427$
 $\Delta\alpha=1.483$**



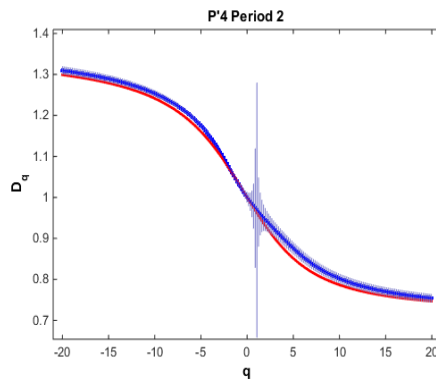
**Period 4 - $q_{sen}=-1.563$
 $\Delta\alpha=0.388$**



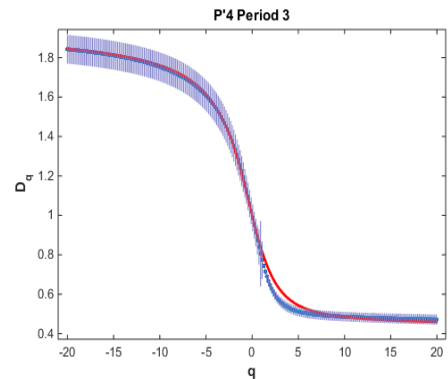
**Period 1 - $\Delta D_q=0.155$
p-model=0.546**



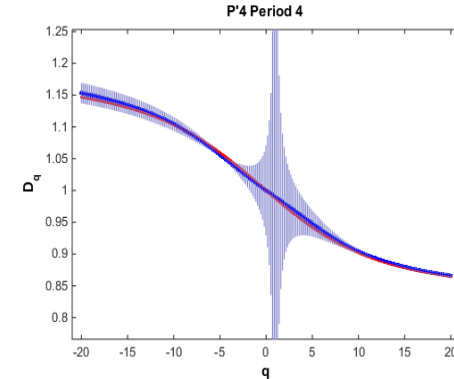
**Period 2 - $\Delta D_q=0.556$
p-model=0.612**



**Period 3 - $\Delta D_q=1.371$
p-model=0.753**

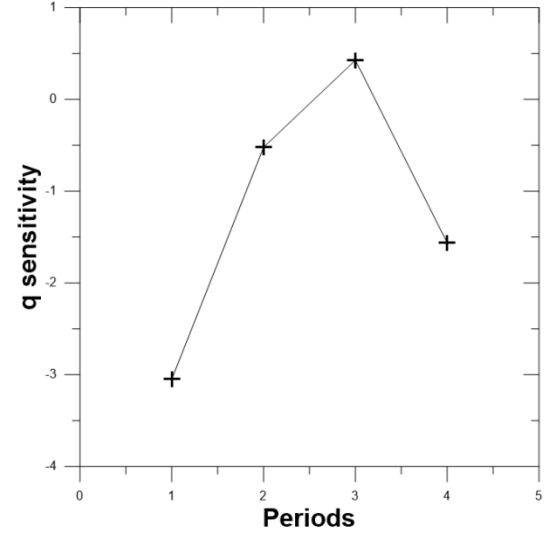
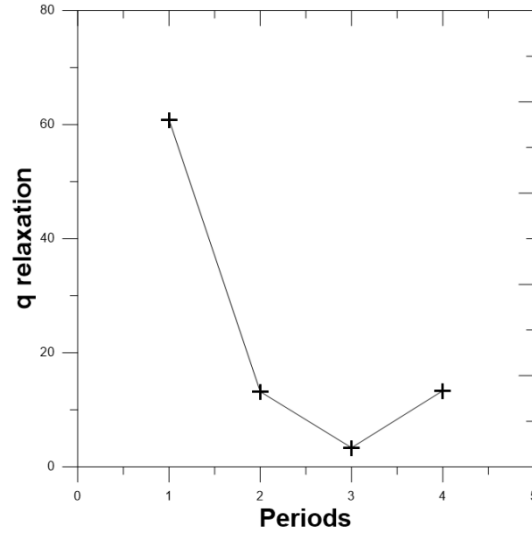
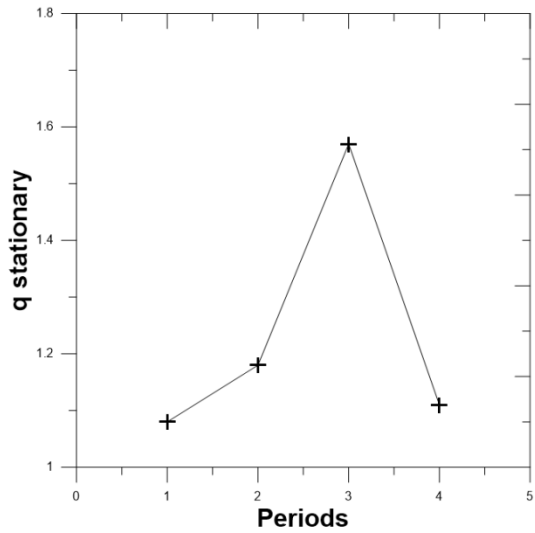


**Period 4 - $\Delta D_q=0.287$
p-model=0.566**

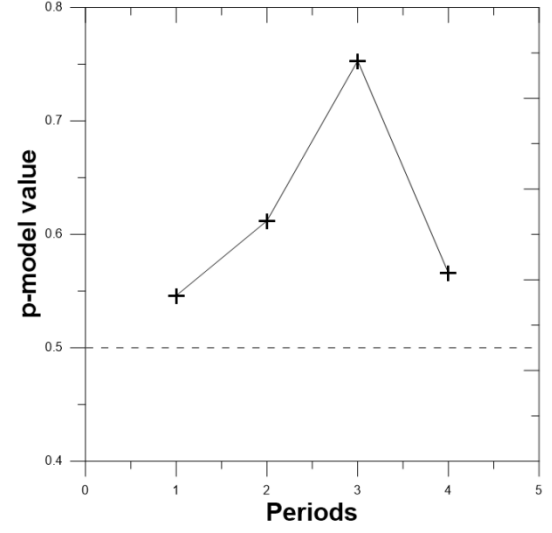
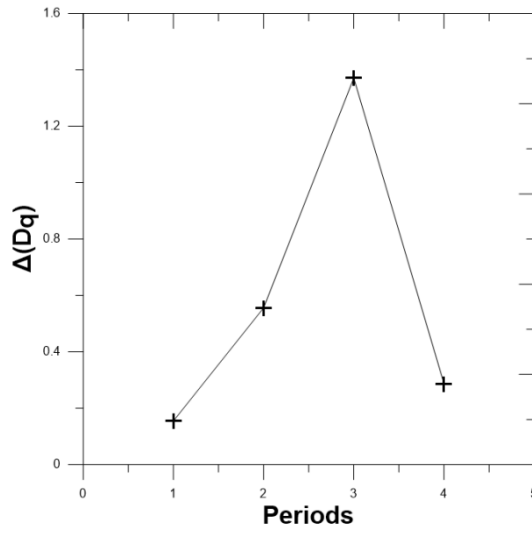
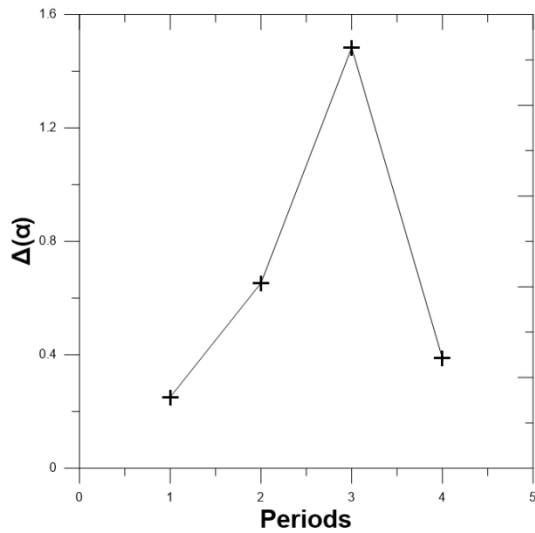


Solar Energetic Particle (Parameters)

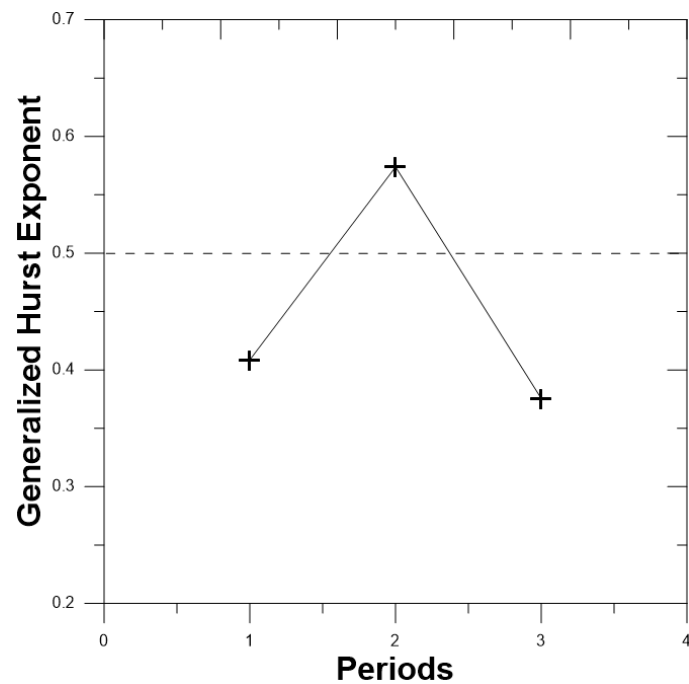
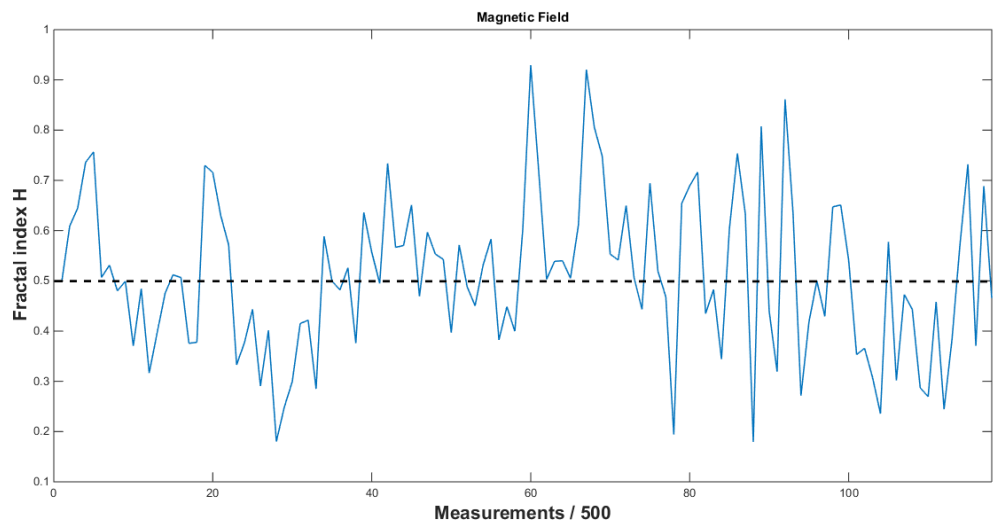
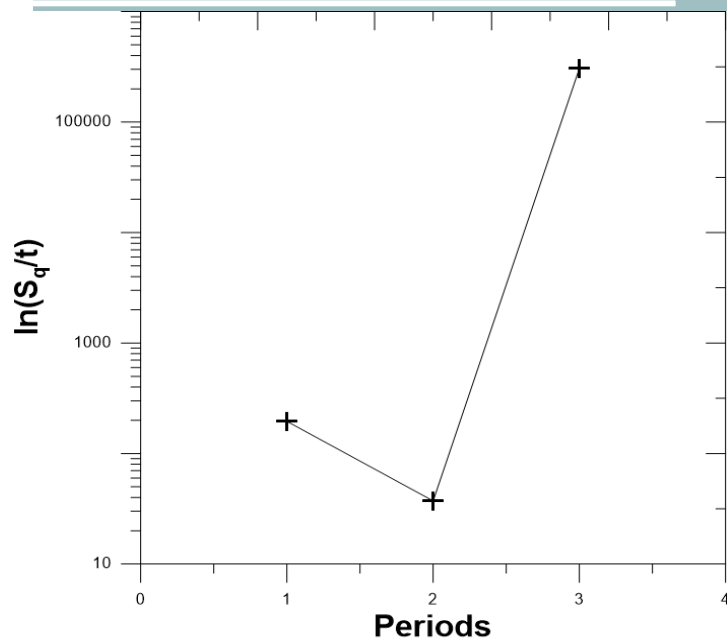
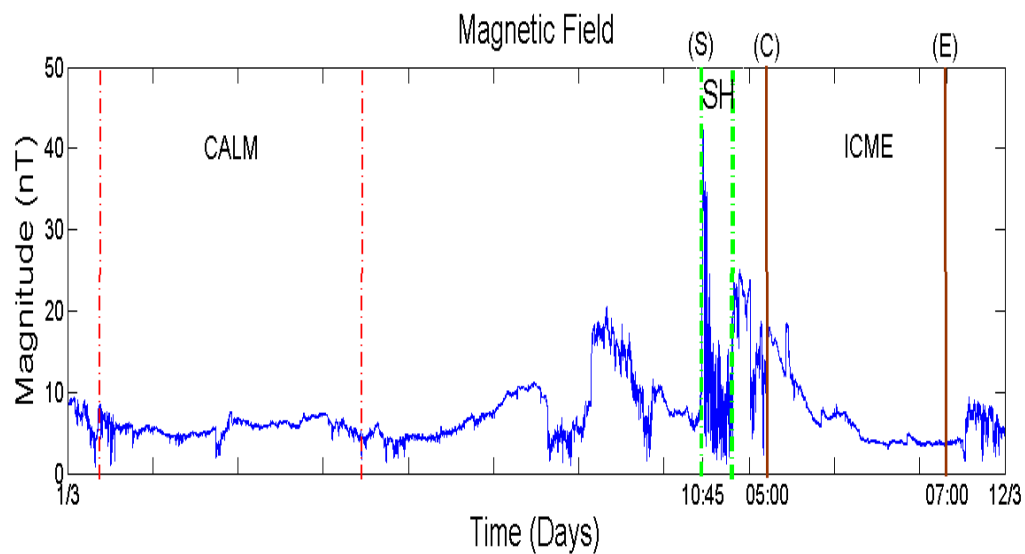
Phase Transition through Tsallis q -triplet



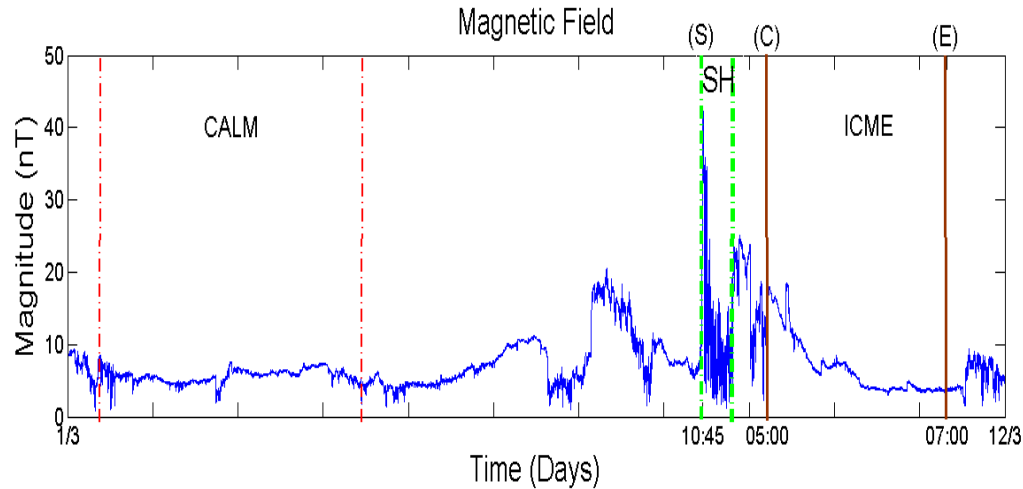
Phase Transition through Multifractal Structure and p -model values



ACE/MAG Project - Solar Wind (Magnetic Field)



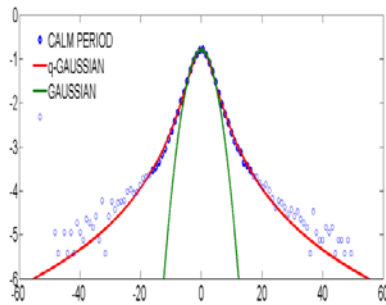
Solar Wind Magnetic Field (q-stationary & q-relaxation)



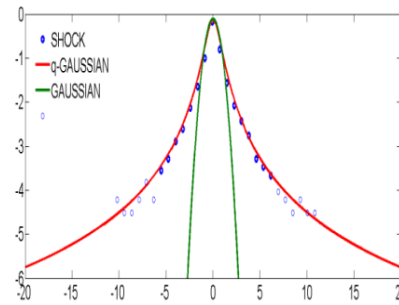
Calm - qstat = 1.38 ± 0.04

Shock - qstat = 1.74 ± 0.05

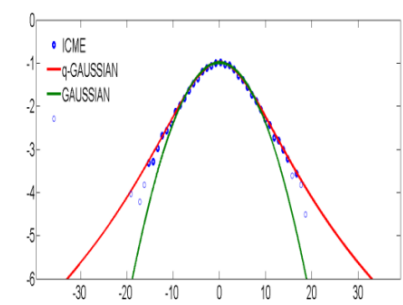
ICME - qstat = 1.17 ± 0.02



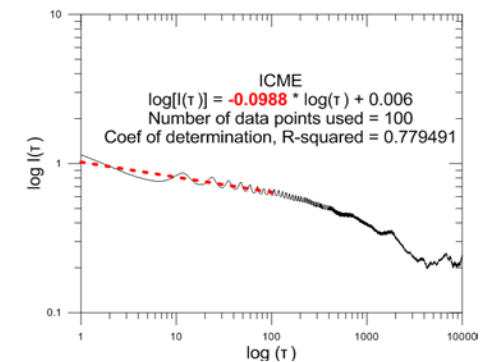
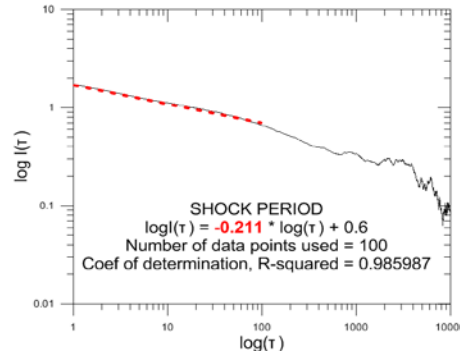
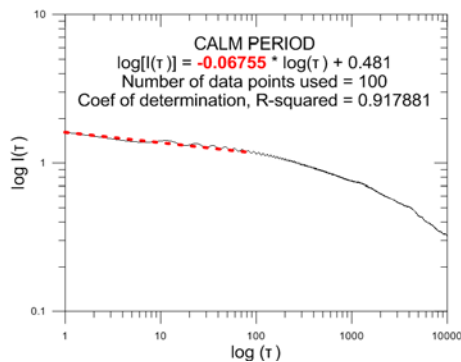
Calm - qrel = 15.804



Shock - qrel = 5.74

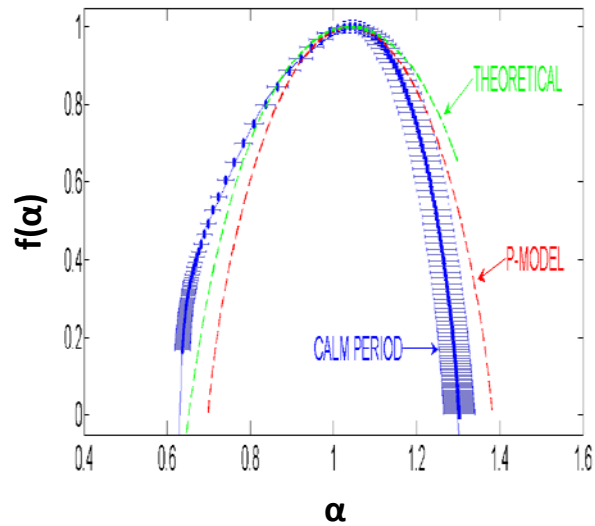


ICME - qrel = 11.12

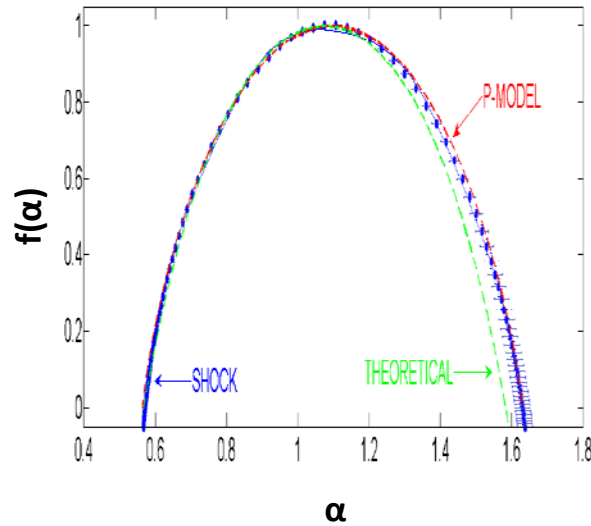


Solar Wind Magnetic Field (Multifractal Spectrum)

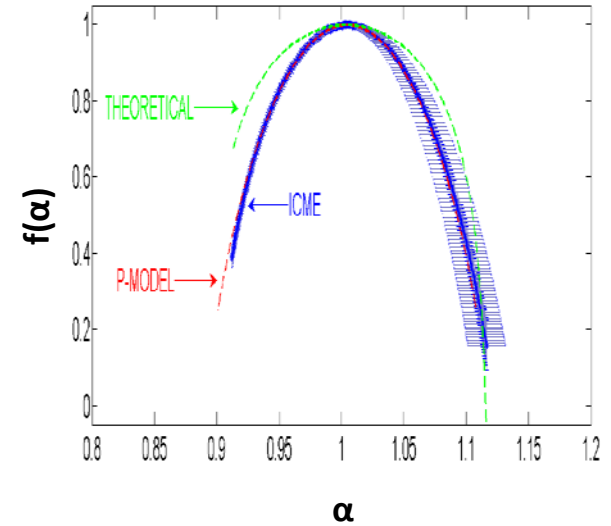
Calm - $q_{sen} = -0.198$
 $\Delta\alpha = 0.679$



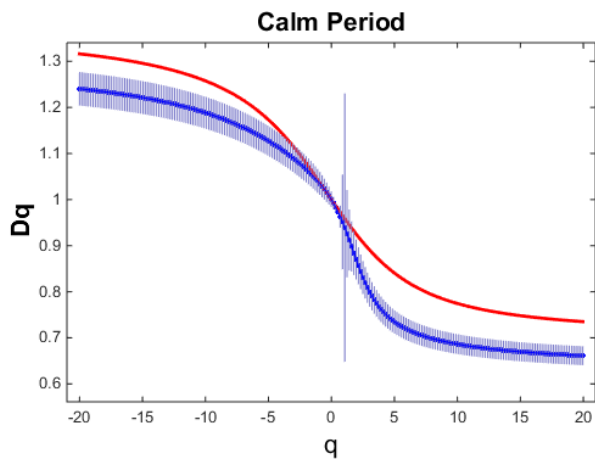
Shock - $q_{sen} = 0.117$
 $\Delta\alpha = 1.057$



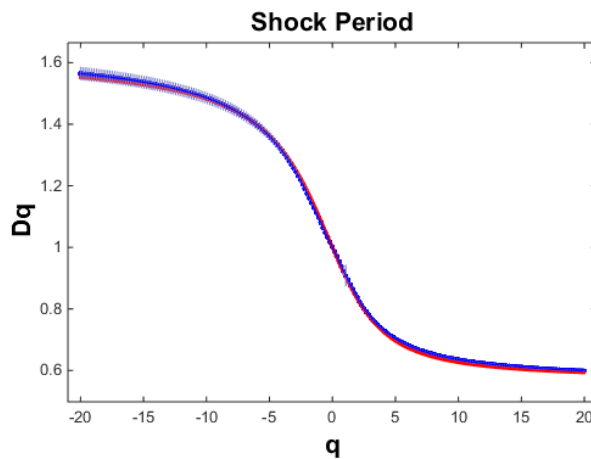
ICME - $q_{sen} = -3.456$
 $\Delta\alpha = 0.227$



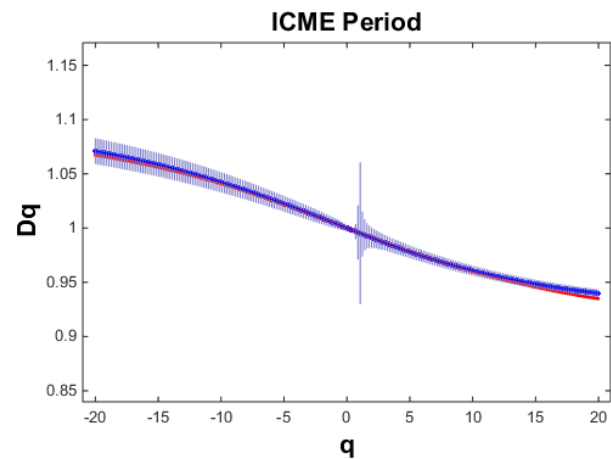
Calm - $\Delta Dq = 0.579$
p-model = 0.616



Shock - $\Delta Dq = 0.967$
p-model = 0.677

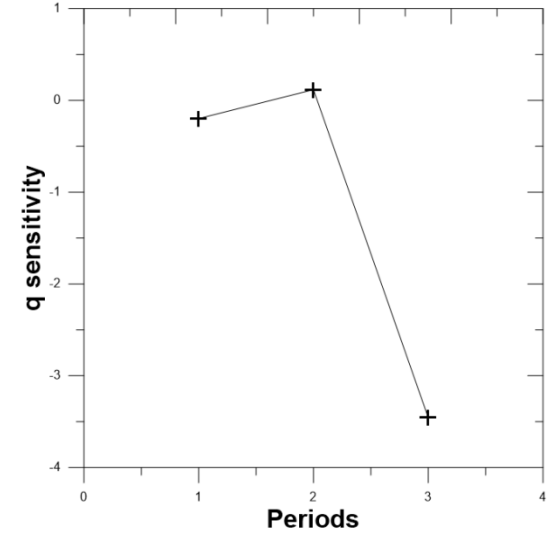
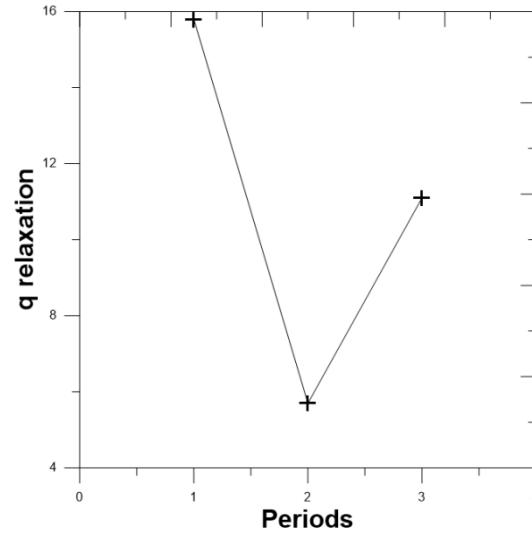
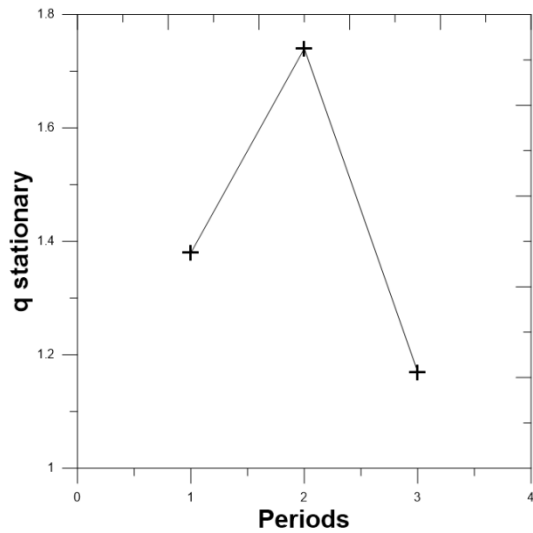


ICME - $\Delta Dq = 0.131$
p-model = 0.539

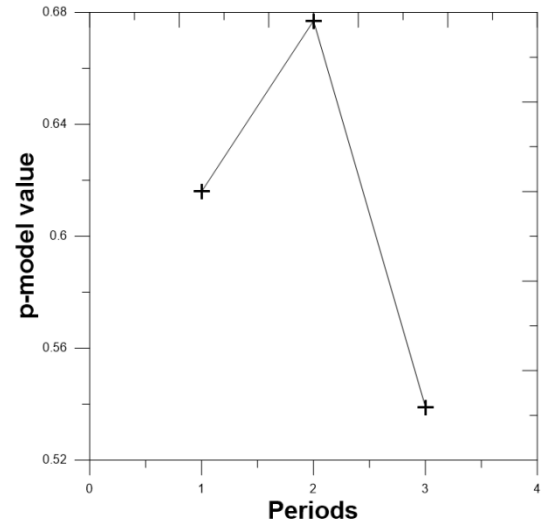
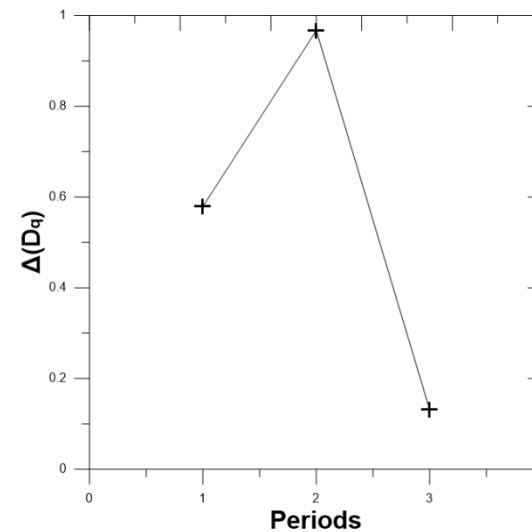
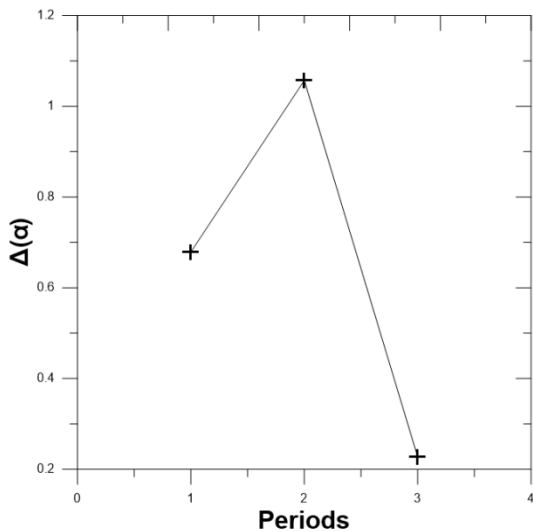


Sola Wind Magnetic Field (Parameters)

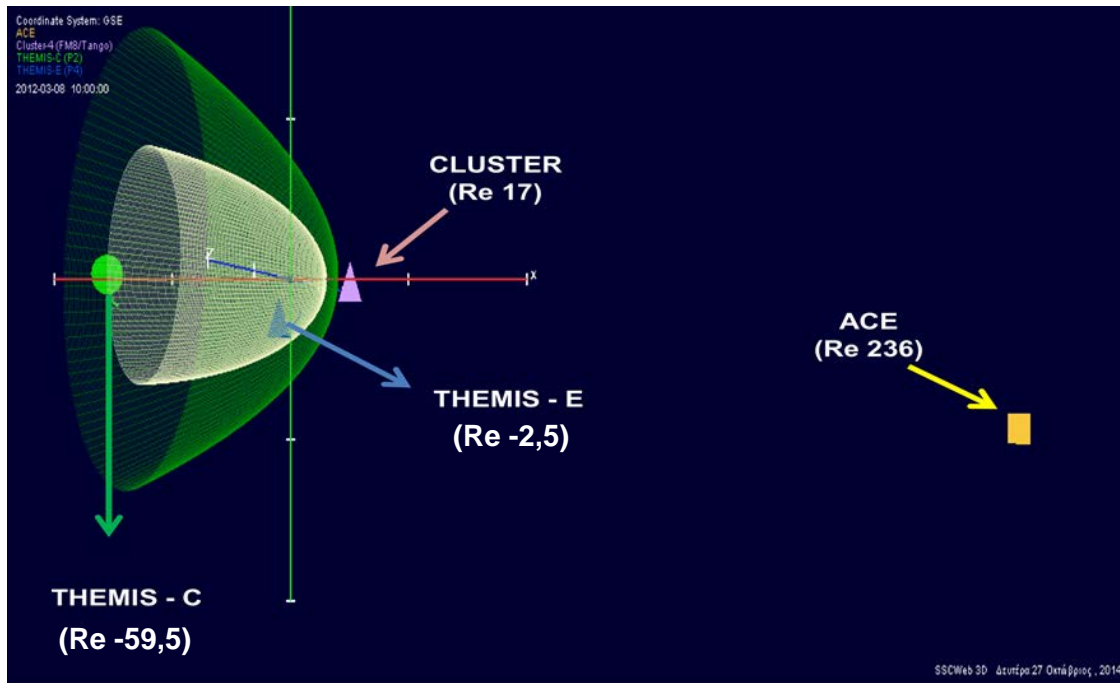
Phase Transition through Tsallis q-triplet



Phase Transition through Multifractal Structure and p-model values



Interplanetary Solar Magnetic Field



	ACE (B_{total})	CLUSTER (B_{total})	THEMIS-E (B_{total})	THEMIS-C (B_{total})
$q_{stationary}$ (CALM)	1.37 ± 0.01	1.26 ± 0.05	1.26 ± 0.05	1.27 ± 0.06
$q_{stationary}$ (SHOCK)	1.74 ± 0.03	1.53 ± 0.07	1.53 ± 0.07	1.73 ± 0.06

Solar Wind (High Time Resolution Ion Fluxes)

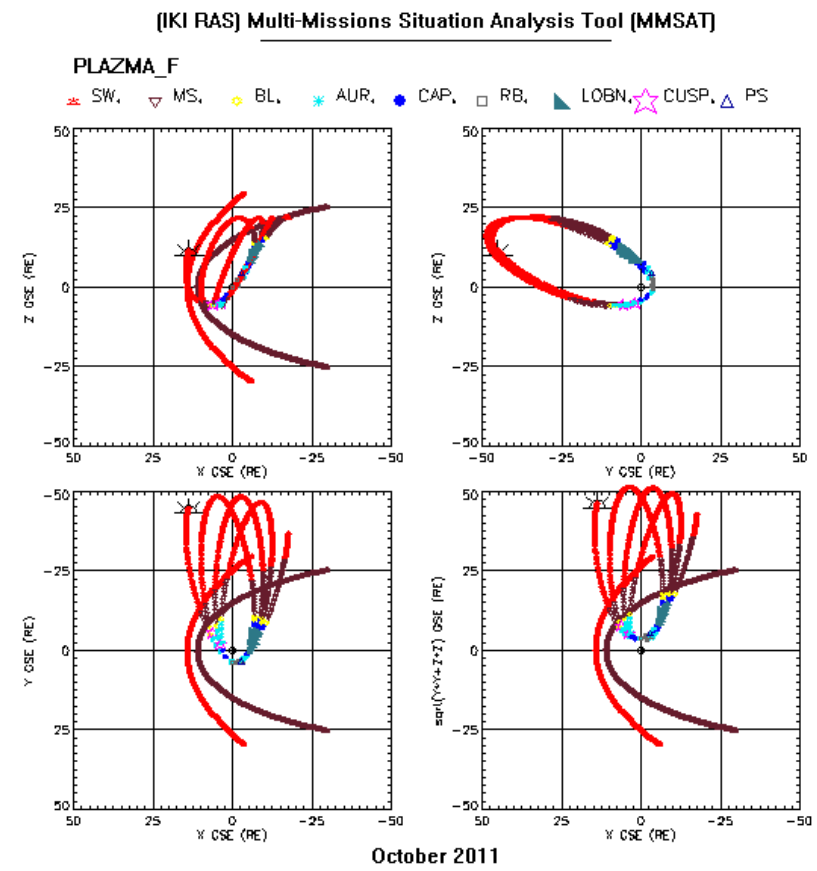
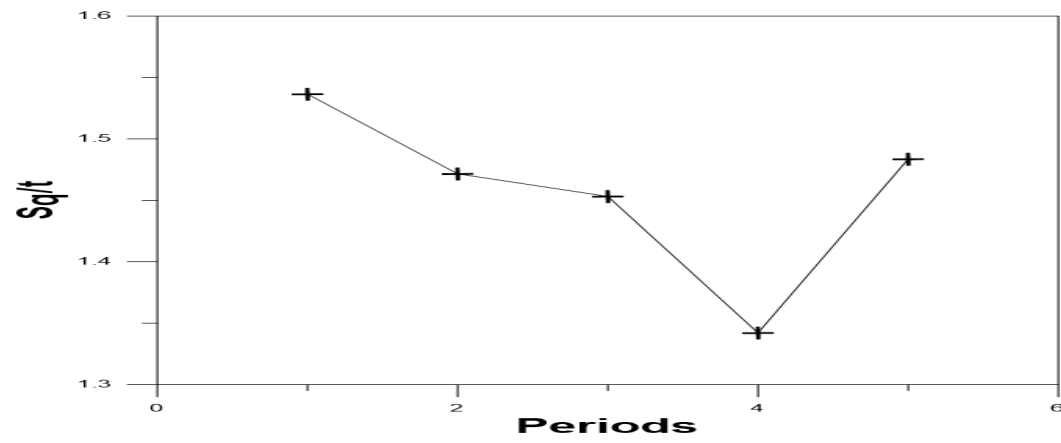
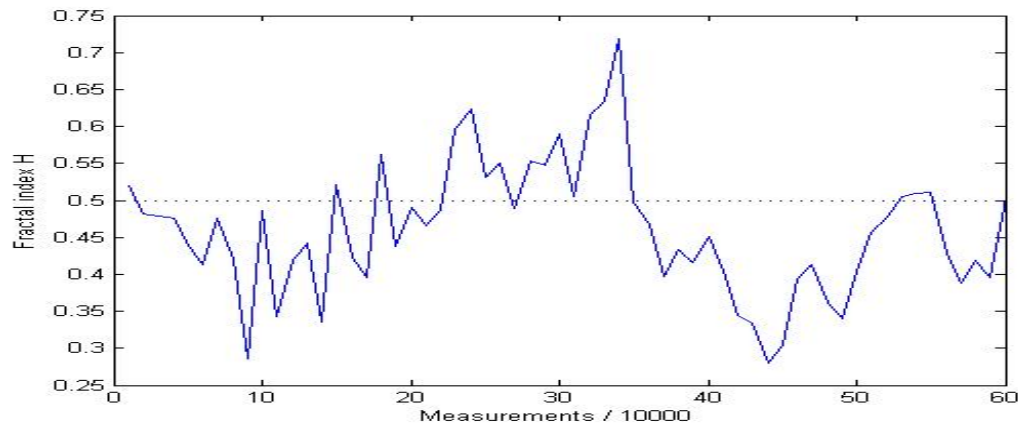
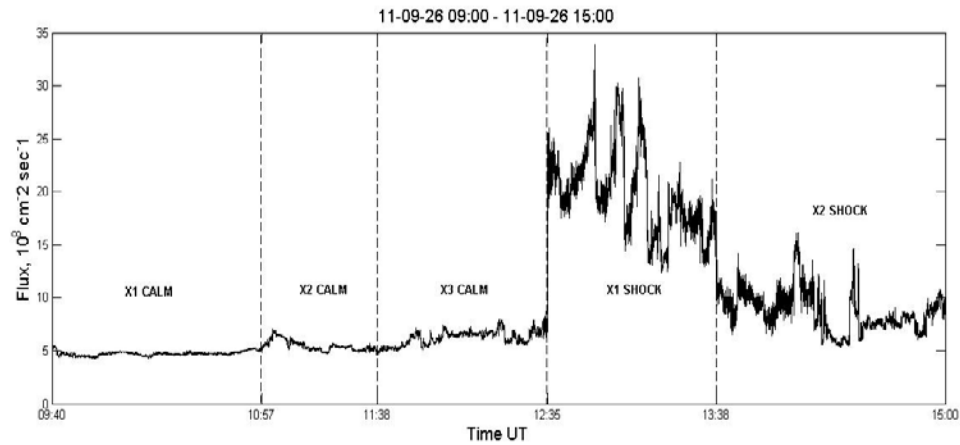


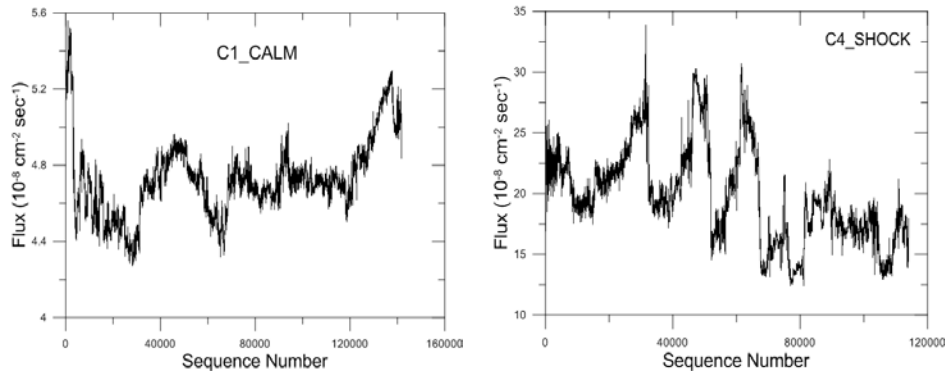
Photo of BMSW instrument (Fast Monitor of the Solar Wind) and Typical SPECTR-R orbits evolution relative the Earth's magnetosphere in October 2011.

Solar Wind Ion Fluxes

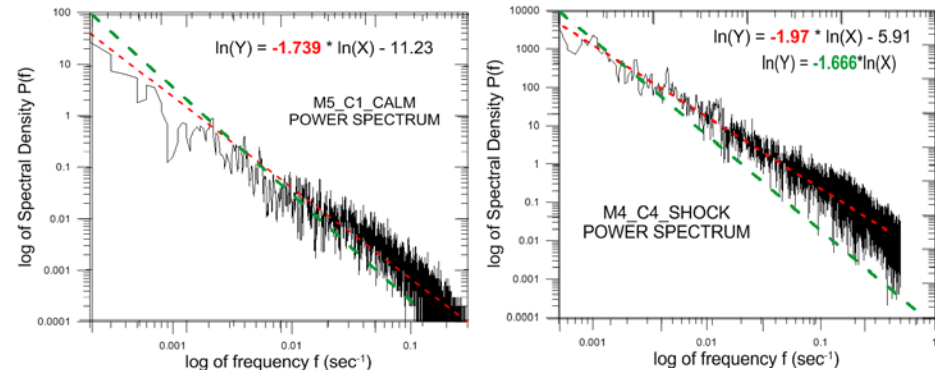


Solar Wind Ion Fluxes (Chaotic Analysis)

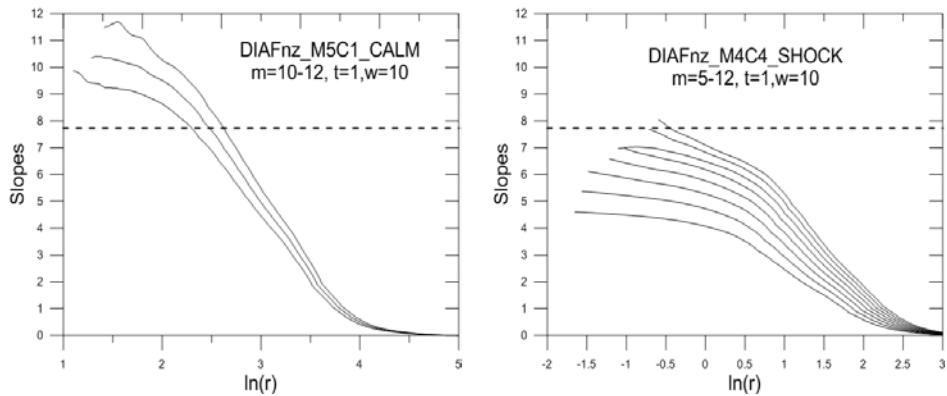
Time Series



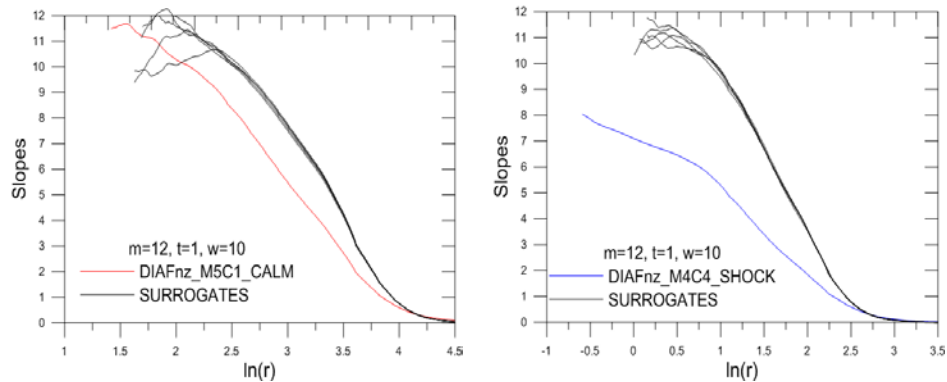
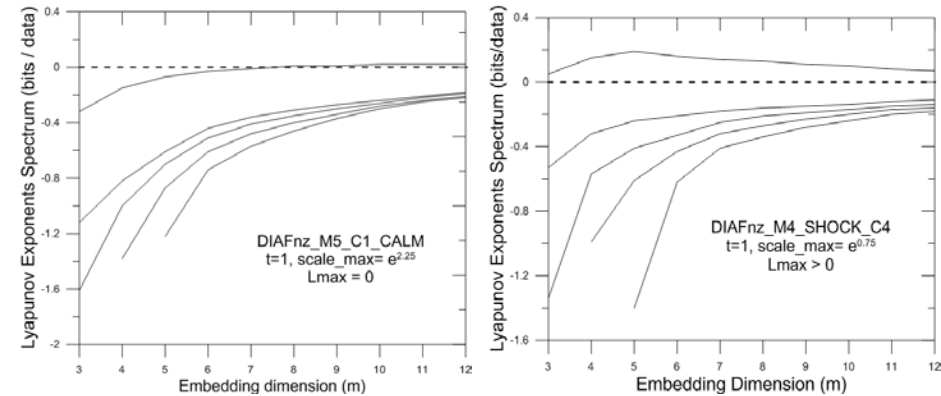
Power Spectrum



Correlation Dimension

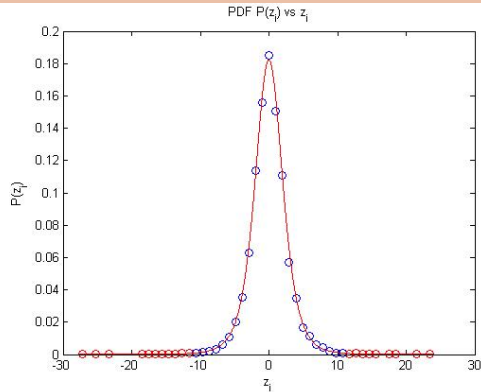


Lyapunov Spectrum Exponents

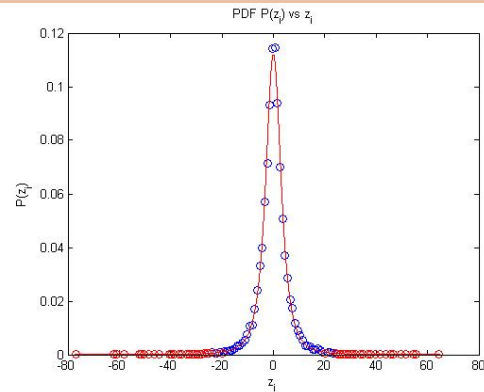


Solar Wind Ion Fluxes (Tsallis Statistics)

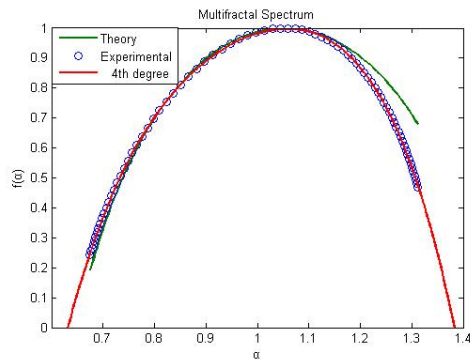
Calm - $q_{stat} = 1.37 \pm 0.05$



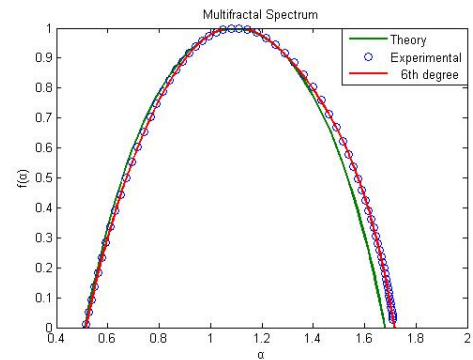
Shock - $q_{stat} = 1.61 \pm 0.03$



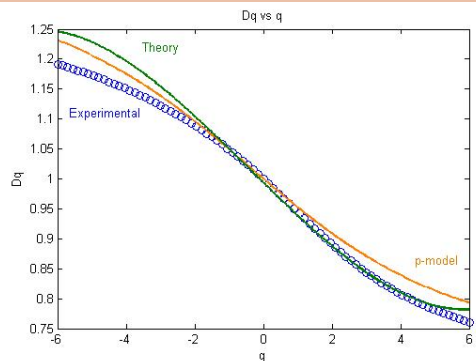
**Calm - $q_{sen} = -0.161$
 $\Delta\alpha = 0.753$**



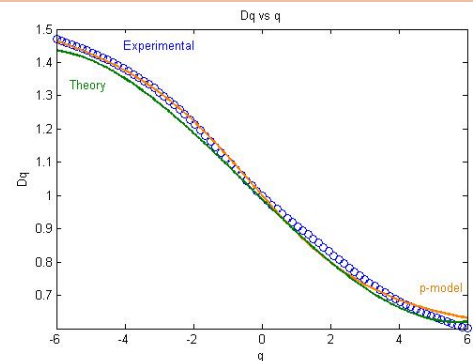
**Shock - $q_{sen} = 0.268$
 $\Delta\alpha = 1.205$**



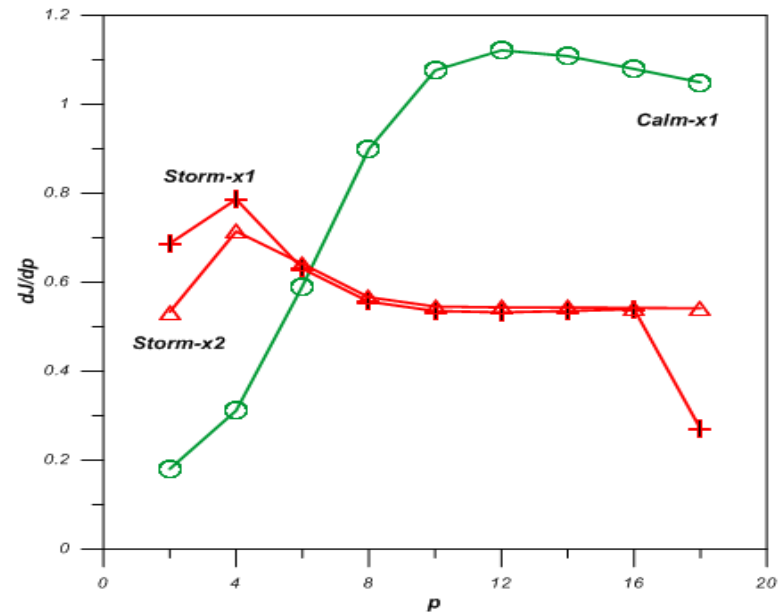
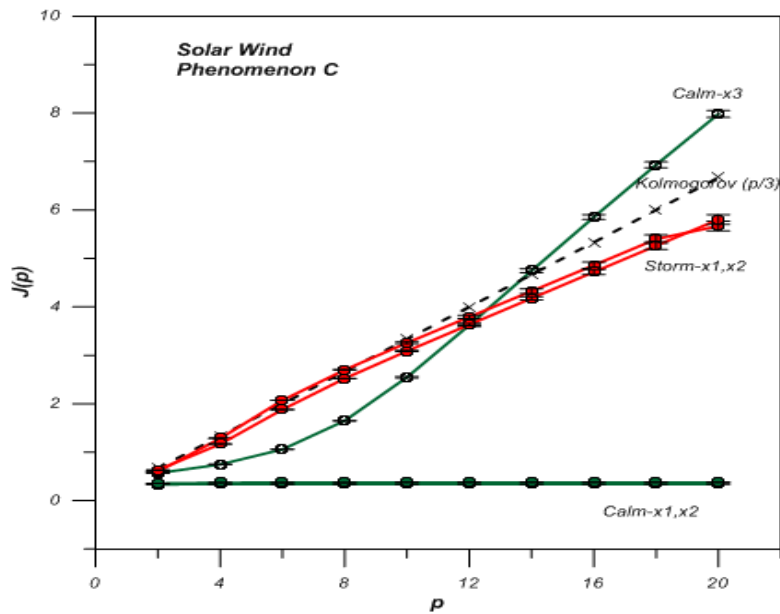
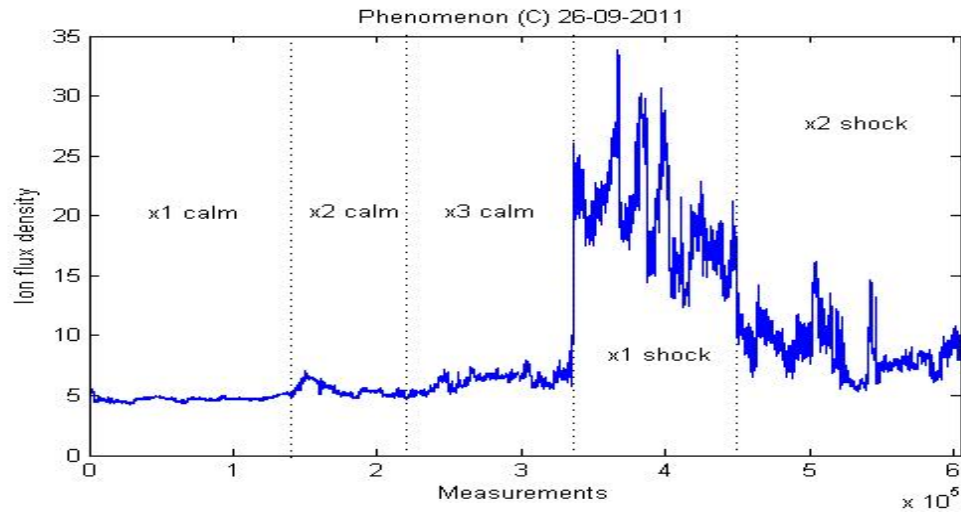
Calm - $\Delta Dq = 0.431$



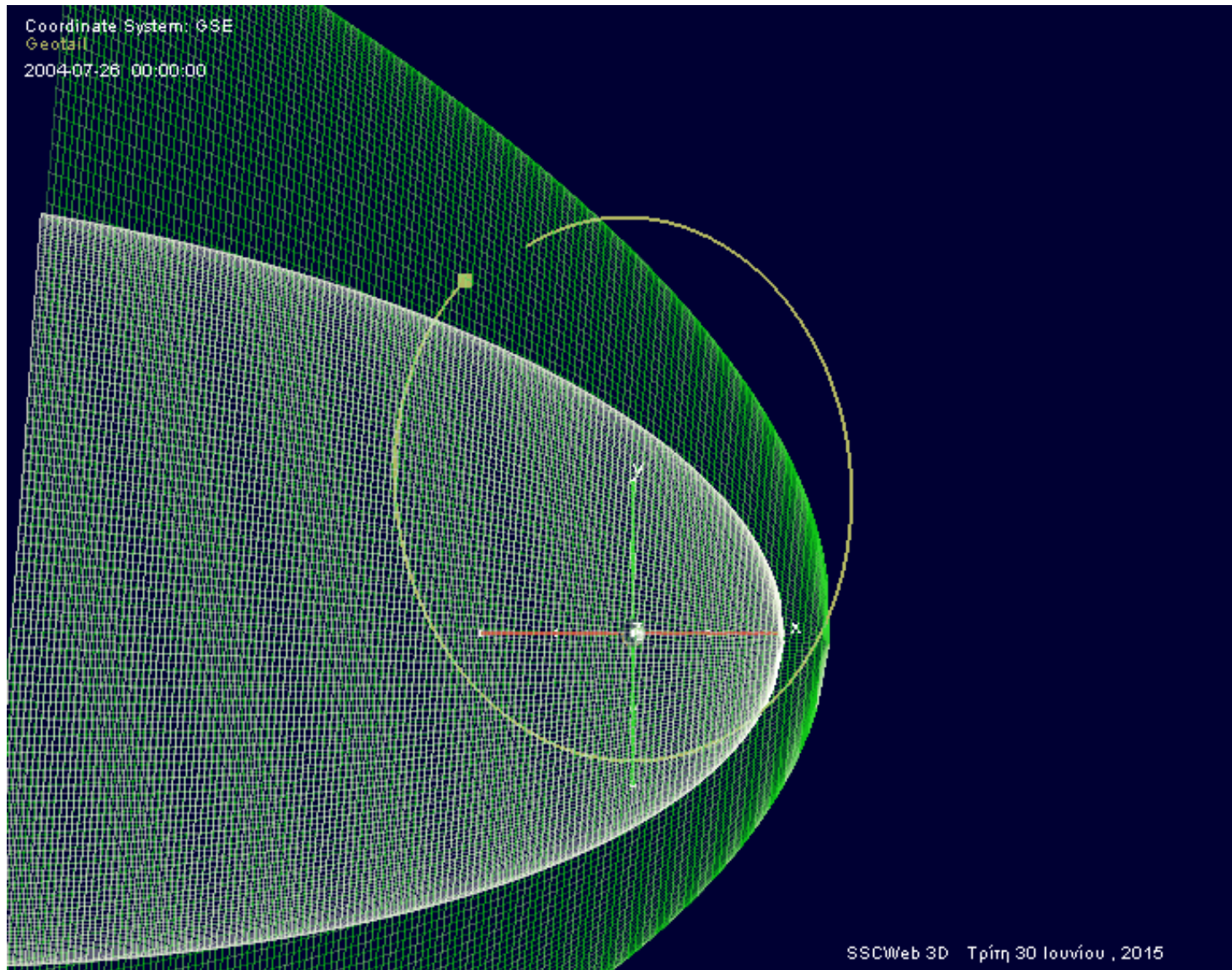
Shock - $\Delta Dq = 0.871$



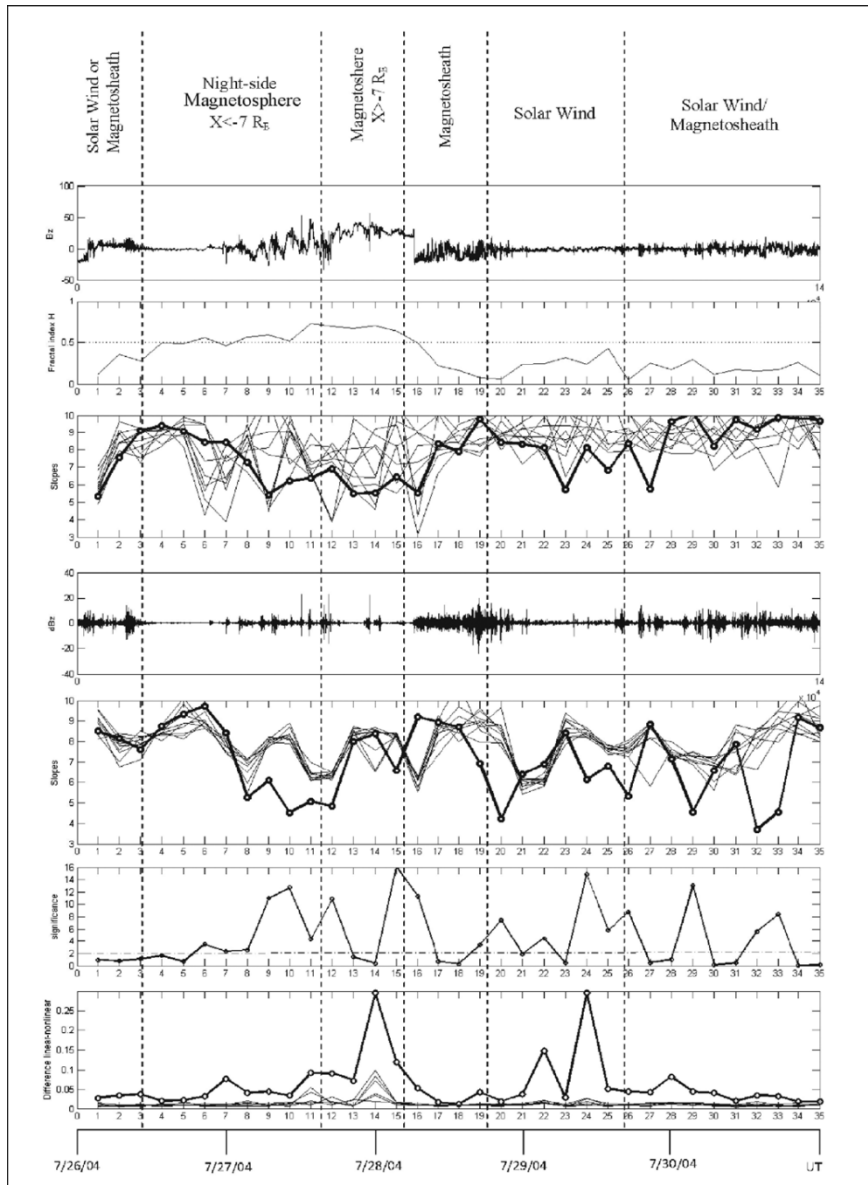
Solar Wind Ion Fluxes (Turbulence Analysis)



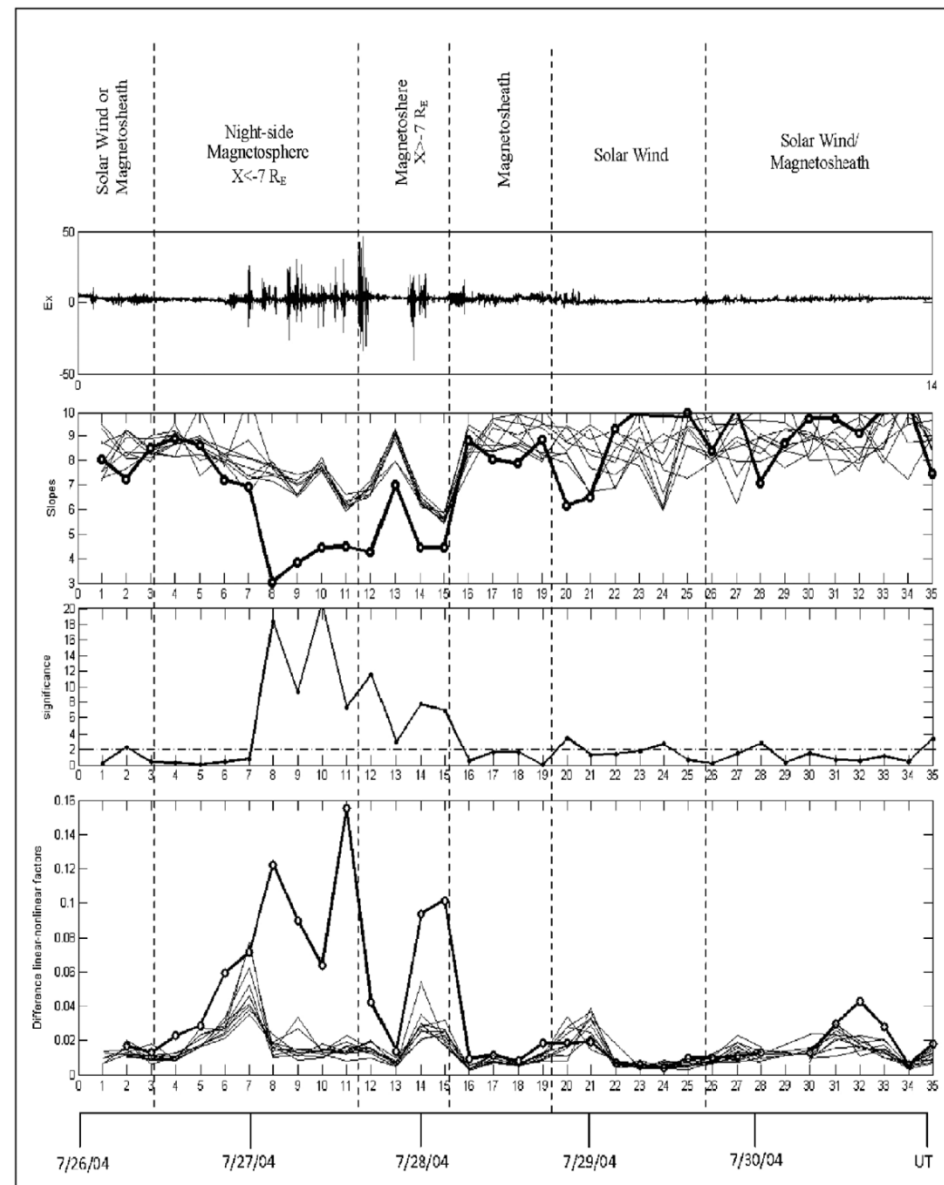
Magnetosphere 26-30 July 2004 - GEOTAIL Orbit



Bz Component



Ex Component



Summarize parameter values of non-extensive statistics in Solar Plasmas
 The q-triplet (q_{sen} , q_{stat} , q_{rel}) of Tsallis.

System	q_{sen}	q_{stat}	q_{rel} (C(τ))
Solar Wind (Bz cloud)	0.484±0.009	2.02±0.04	32.25±2.31
Solar (Sunspot Index)	0.368±0.005	1.53 ± 0.04	5.67±0.13
Solar (Flares Index)	0.308±0.005	1.870±0.005	5.33±0.22
Solar (Protons)	0.817 ± 0.024	2.31±0.13	2.48±0.23
Solar (Electrons)	0.860±0.009	2.13±0.06	2.58±0.22
Cosmic Stars (Brighness)	-0.568±0.042	1.64±0.03	7.71±0.58
Cosmic Ray (C)	-0.441±0.013	1.44±0.05	193.30±18.43

Conclusions

Everywhere in Solar plasma we have confirmed the following:

- Tsallis Sq entropy and non extensivity
- Nonequilibrium stationary states (NESS)
- Nonequilibrium topological phase transition processes
- Multifractal intermittent turbulence
- Anomalous diffusion
- Multiplicative processes
- Low Dimensional Chaos and/or SOC.

These experimental results are concluded theoretically by:

- Strange Dynamics
- Fractal extension of MHD theory
- Fractional extension of magnetized plasmas Langevin and Fokker - Planck equations
- Nonequilibrium Percolation Theory



THANK YOU
for your attention