Variability of massive stars in the Virgo Cluster galaxy NGC 4535 with the Hubble Space Telescope



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## Motivation

 Constraining evolutionary models of massive stars (e.g. rotation, metallicity, binarity)

Increase observations of massive stars in different environments

Variability is a powerful tool for identifying massive stars

 Validation of the "Hubble Catalog of Variables" (HCV) for WFPC2

### NGC 4535

- HST Key Project galaxy in Virgo Cluster
- Variability search revealed 50 Cepheids (Macri et al. 1999)
- Distance from P-L relation: 16 ± 1.9 Mpc (Macri et al. 1999)
- More variable stars expected



# Photometry

- HST/WFPC2
  4 chips (PC, WF2, WF3, WF4)
- F555W: 12 epochs
- F814W: 8 epochs
- May-August 1996
- PSF Photometry: DOLPHOT (Dolphin 2000)
- Distortion and CTE corrections were taken into account
- 24353 stars detected in both filters



# Variability Indexes

- Mean Absolute Deviation
  (MAD)
- Interquartile Range (IQR)
- von Neumann ratio (1/η)
- $\odot$  3 $\sigma$  cutoff
- Selection criteria: Variable in any of the three indexes in either filter



#### New variable stars in NGC 4535

WFPC2	# Variables
PC	17
WF2	50
WF3	23
WF4	30
Total	120

# Color Magnitude Diagram





MESA models (Choi et al. 2016)





### **Candidate RSGs**





V14



### Candidate YSGs/YHGs







#### Candidate LBV???



# LBVs

- Evolved stars: Mbol < -9.5 mag</p>
- Giant eruptions: >2 mag (η Carinae)
- Eruptions: ~1-2 mag within 10-40 yrs
- Smaller oscillations: ~ 0.5 mag
- Microvariations: <0.1 mag</p>





HST images of two bipolar LBV nebulae (Weis 1999, 2011)

## Candidate LBV

- Mv ~ -11 mag
- ♂ T<sub>eff</sub> ~ 9000 K
- Microvariations: ~ 0.1 mag





### Conclusions

- 120 new variable stars in NGC 4535 using archival HST/ WFPC2 data
  - 8 candidate RSGs
  - a 3 candidate YSGs/ YHGs
  - I candidate LBV (M<sub>v</sub> ∼ -11 mag)
- Identified massive star candidates at 16 Mpc
- Successful selection of variable stars with variability indexes

## **Future Projects**

- Apply method to other HST Key Project galaxies to search for new variablity and massive stars
- Our Derive new P-L relation with DOLPHOT photometry
- Follow-up spectroscopy and Pan-STARRS photometry for the candidate LBV

# THANK YOU!!

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