

# A PERTURBED GALACTIC DISK

Minor mergers and inflows of primordial gas

Monday, July 25<sup>th</sup>, 2022

Hellenic Astronomical Society Summer School 2022

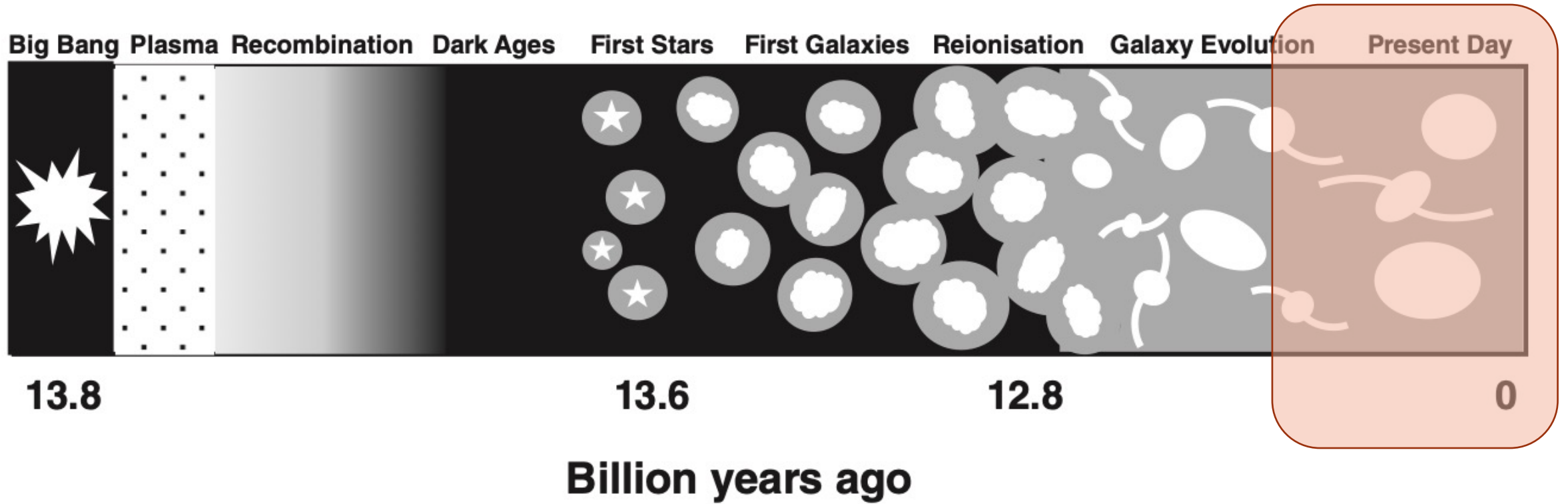


**Santi Roca-Fàbrega**

Instituto de Astronomía de la UNAM, Ensenada, México

# GALAXY FORMATION AND EVOLUTION

- The Universe in the LambdaCDM paradigm.

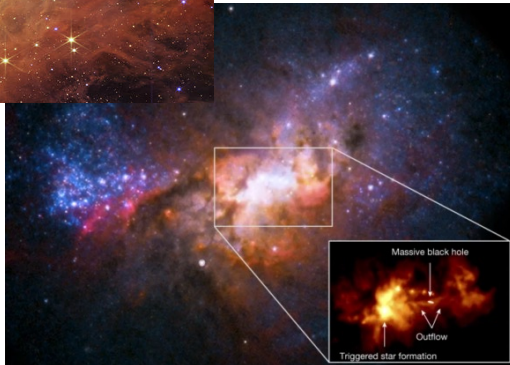


# GALAXY FORMATION AND EVOLUTION

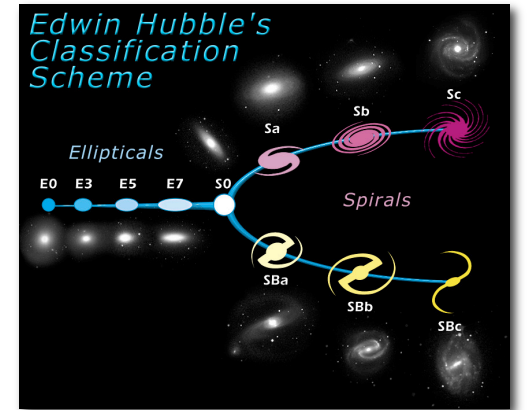
- We focus on disk galaxies (Sa-Sc).
  - Masses  $\sim$  or  $< 10^{12} M_{\text{sun}}$  ( $M_* \sim$  or  $< 10^{10.5} M_{\text{sun}}$ )
  - Star forming
  - Gas rich



JWST



JWST



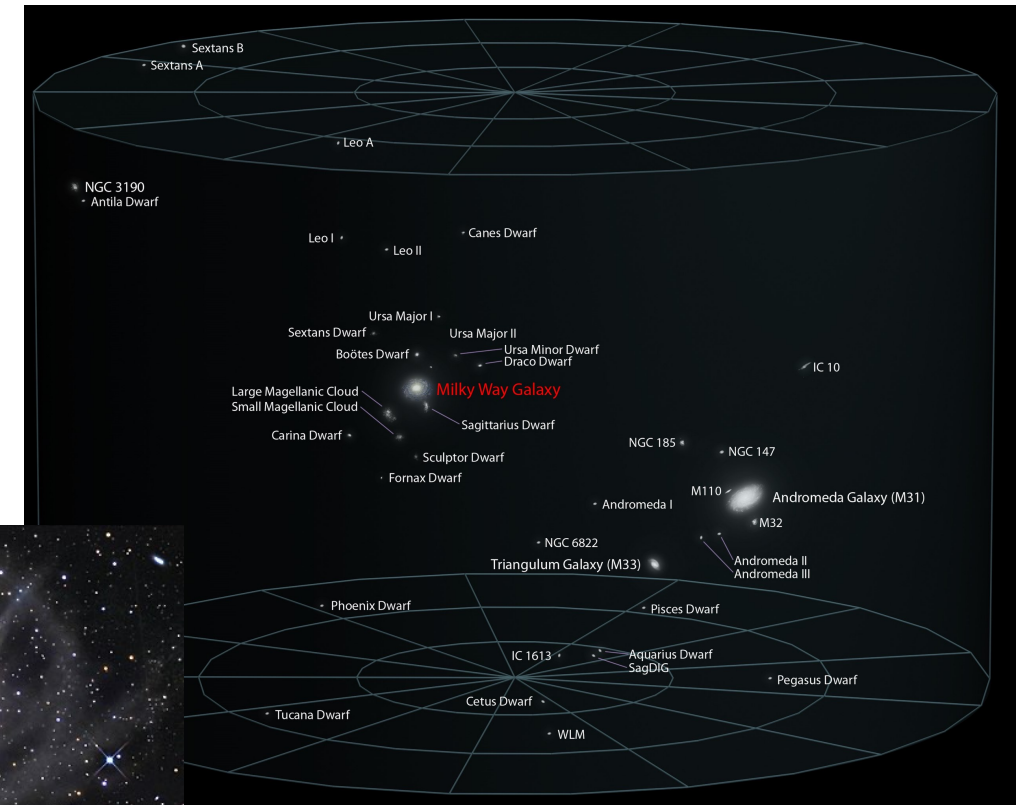
- Star formation
  - Collapse of molecular clouds:
    - Perturbation
    - Cold-dense gas
  - Gas consumption + heating by SNe
    - New gas from inflows
  - Chemistry of the disk as a proxy



# GALAXY FORMATION AND EVOLUTION

## MINOR MERGERS

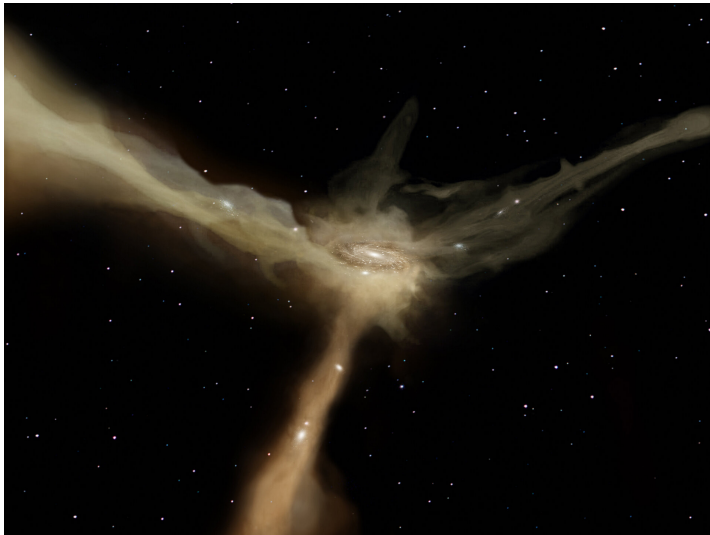
- In the current Universe the galaxies grow mainly through accretion of smaller galaxies (satellites).
- Minor mergers can bring gas to galaxies and also perturb the thin galactic disks.



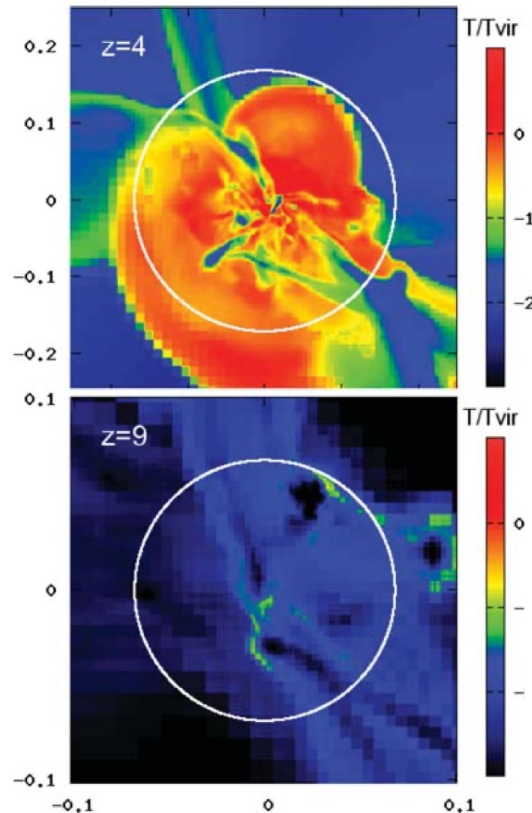


# GALAXY FORMATION AND EVOLUTION

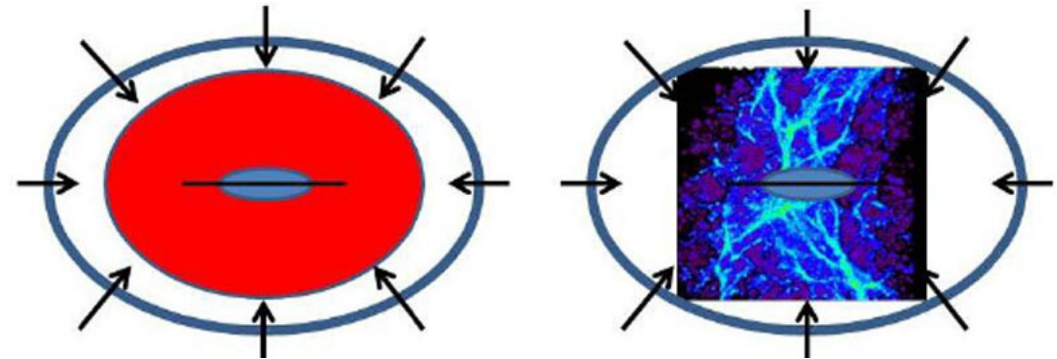
## GAS INFLOWS



Dekel & Birnboim 2006



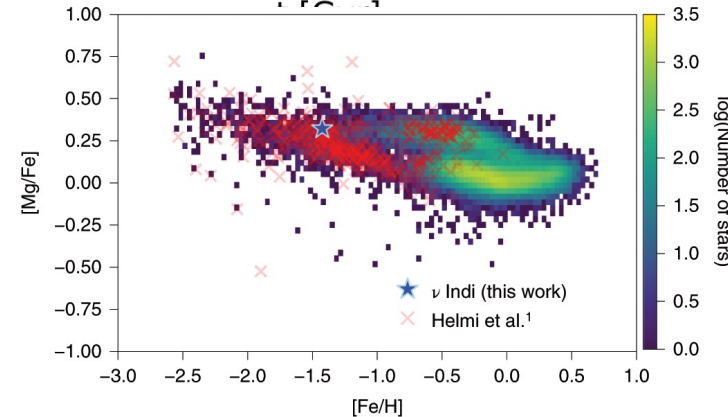
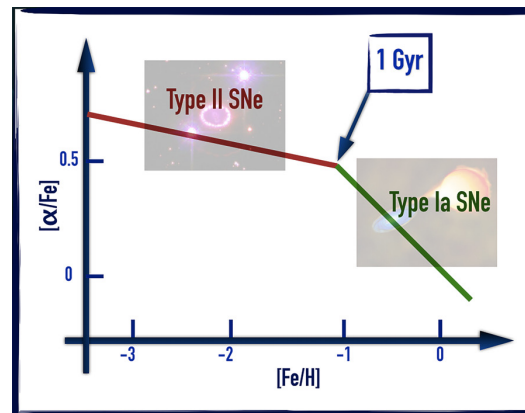
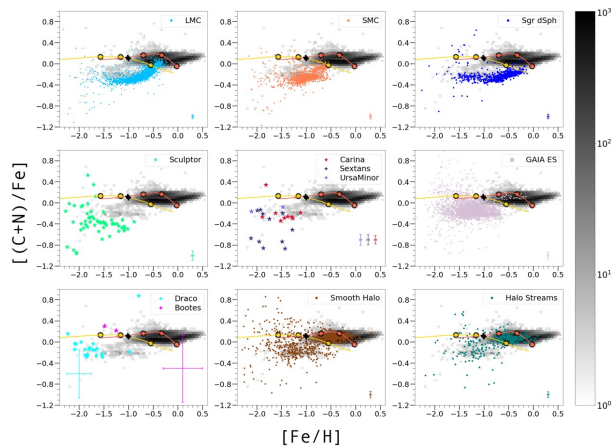
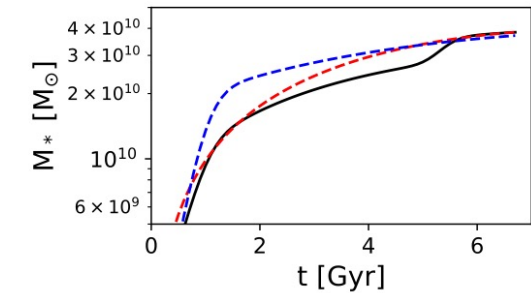
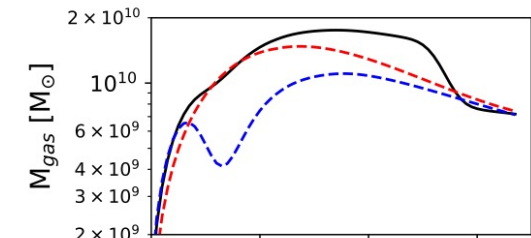
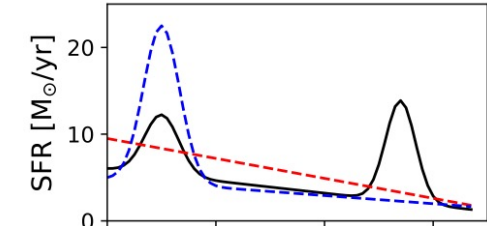
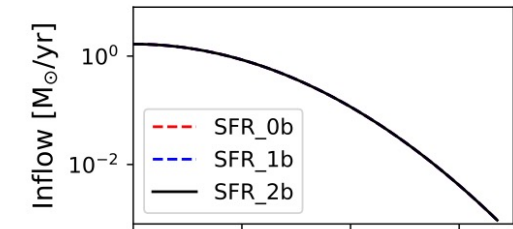
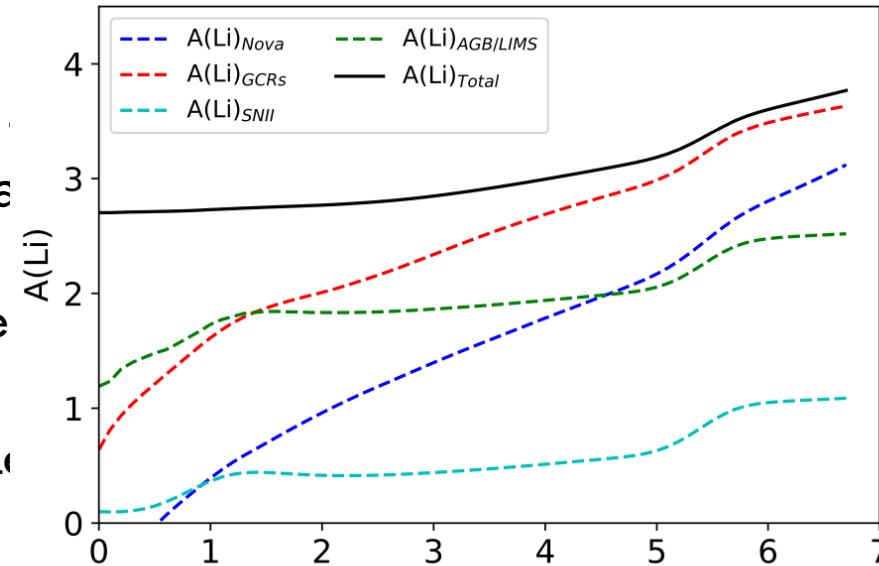
- Galaxies are connected through filaments in the cosmic net.
- Cold gas flows from the intergalactic medium to galaxies through these filaments (cold flow).
- Galaxies above the virial shock mass stop the cold flow and quench.



# RESULTS FROM SIMULATIONS

## SEMI-EMPIRICAL MODELS

- Galaxy chemistry evolution models try to reproduce the observed Galaxy chemistry from basic parameters:
  - Initial metallicity and gas mass
  - Star Formation History (implicitly include gas inflow and outflow)
  - Gas inflow History (metals dilution)
  - SNe / CR / Stellar evolution (release of metals)

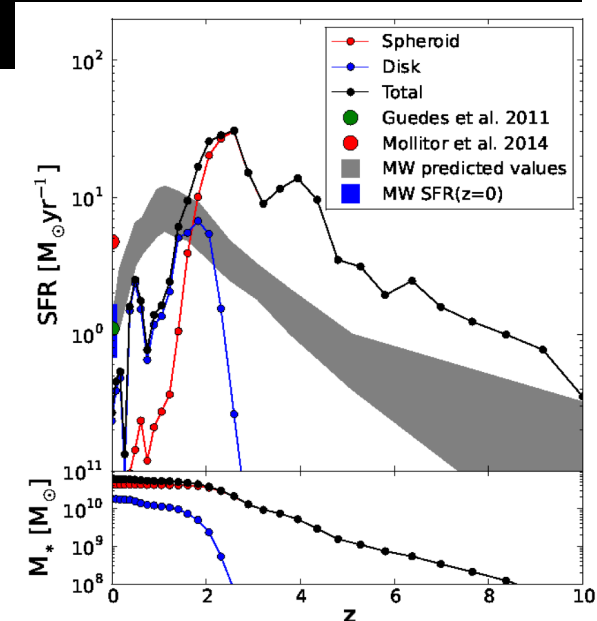
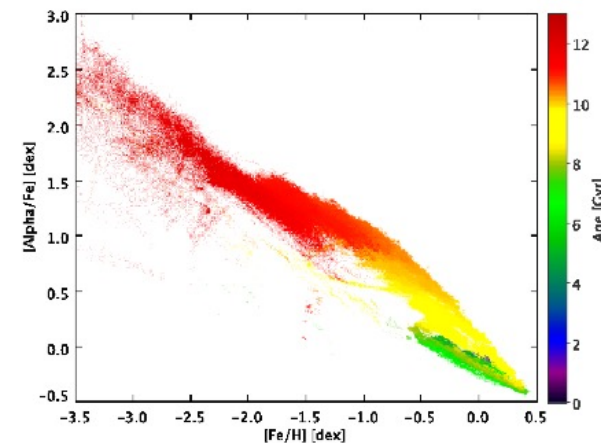
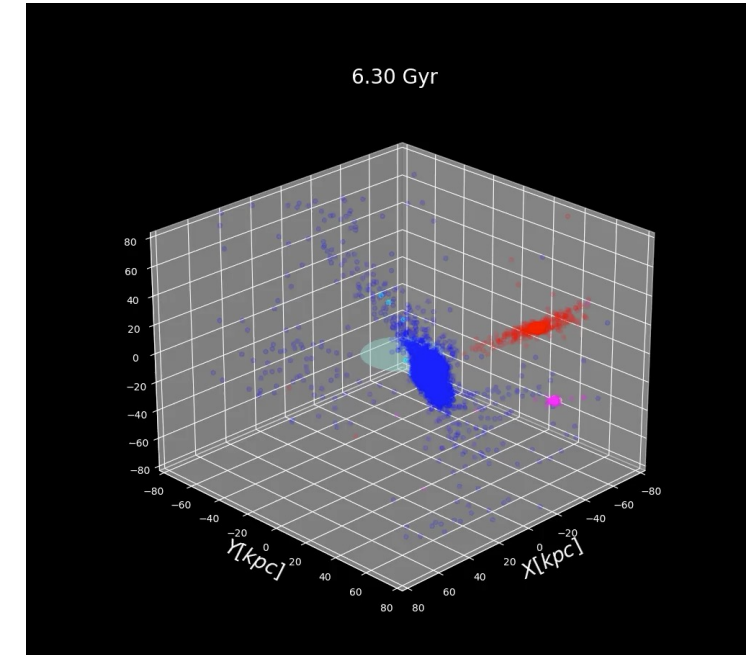


Chaplin et al. 2020

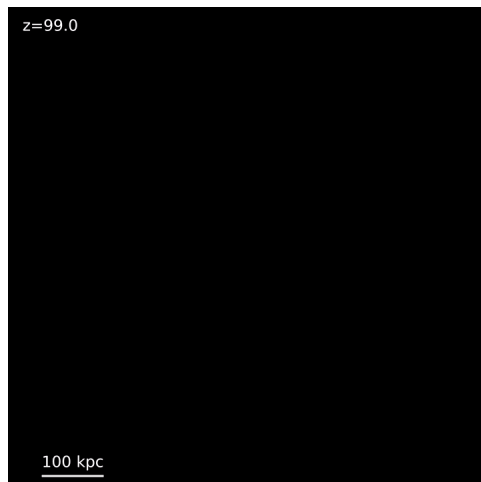
# RESULTS FROM SIMULATIONS

## SELF-CONSISTENT SIMULATIONS

- Both isolated and cosmological simulations, when including hydrodynamics should recover results from semi-empirical models.
  - Require of a large amount of CPU time
  - Subgrid physics (resolution limitations)
  - Star formation history arises naturally
  - Chemistry evolution is “subgrid-physics dependent”



Wetzel et al. 2016





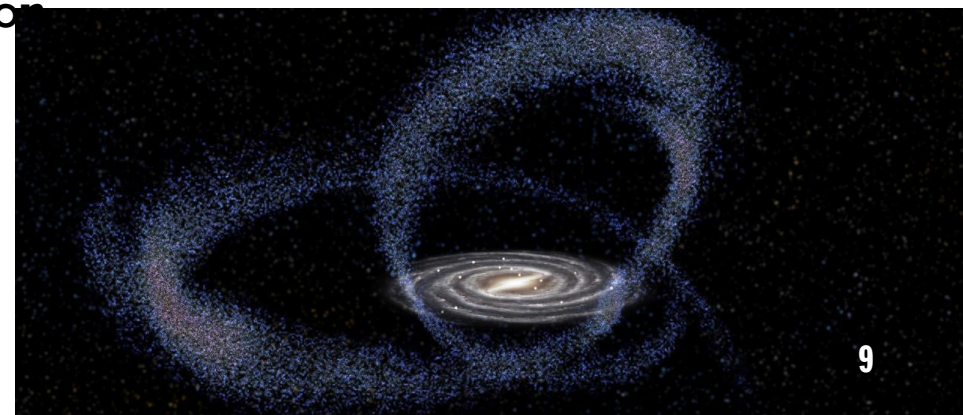
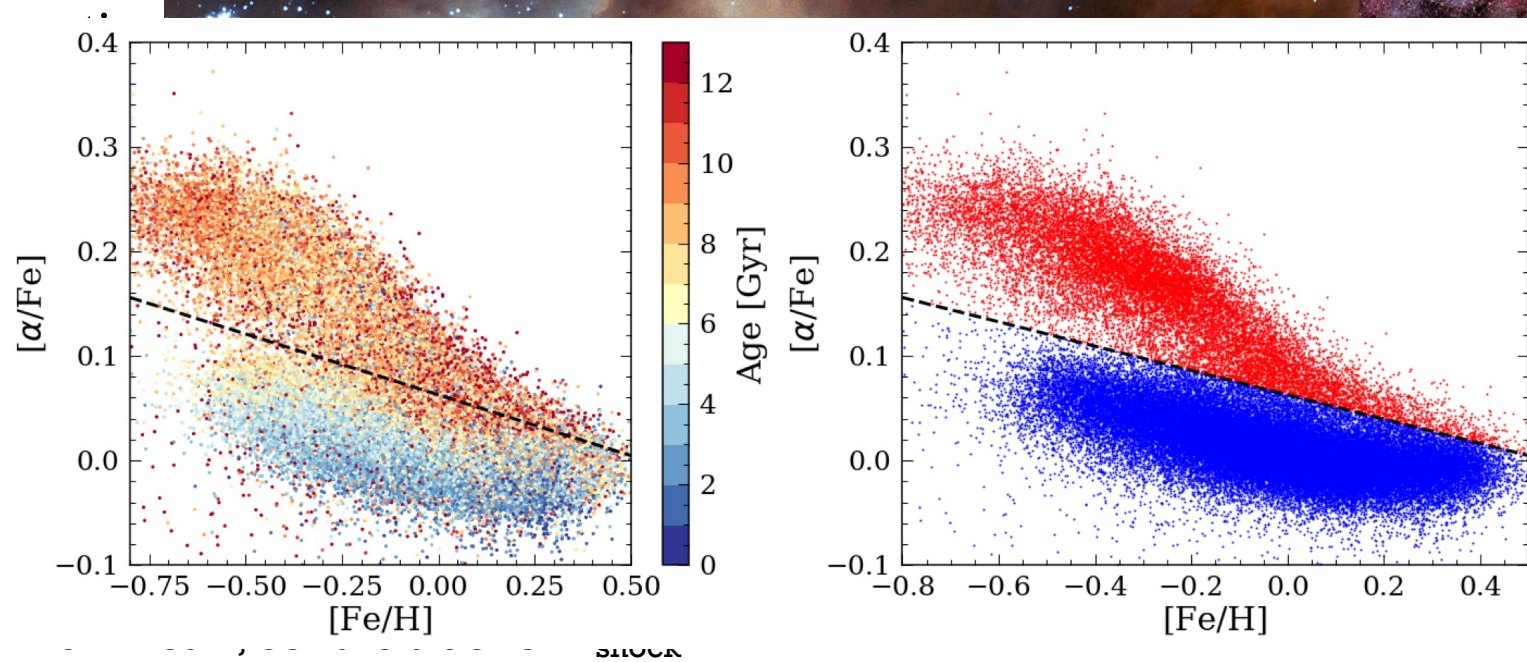
# RESULTS FROM SIMULATIONS

## SELF-CONSISTENT SIMULATIONS

- Simulations tell us that any minor merger should affect the galactic disk if enough massive and close to the disk
- The LambdaCDM predicts (confirmed by simulations) that minor mergers are usual in disk galaxies and so the disks are constantly perturbed
- Is the Milky Way an exception?



- We have the theory of galaxy formation
- We have predictions from simulations
- Which is the true story for the Milky Way?
- What do we know?
  - The MW is a disk galaxy with a central bulge
  - We know that many satellites accompanied the MW for almost the last 10 Gyr
  - We see gas and dust inside the disk and also recent star formation
  - We know the disk has two components, one thin, one thick
  - There is a clear bimodality in the disk chemistry





# THE MILKY WAY FORMATION HISTORY

- We only see an snapshot of the MW formation and evolution history: we need to do **Galactic Archaeology**
- The events that shaped the MW galaxy should be fossilized in the stellar kinematics and chemistry, and also in its morphology (spiral arms and bar).





# THE MILKY WAY FORMATION HISTORY

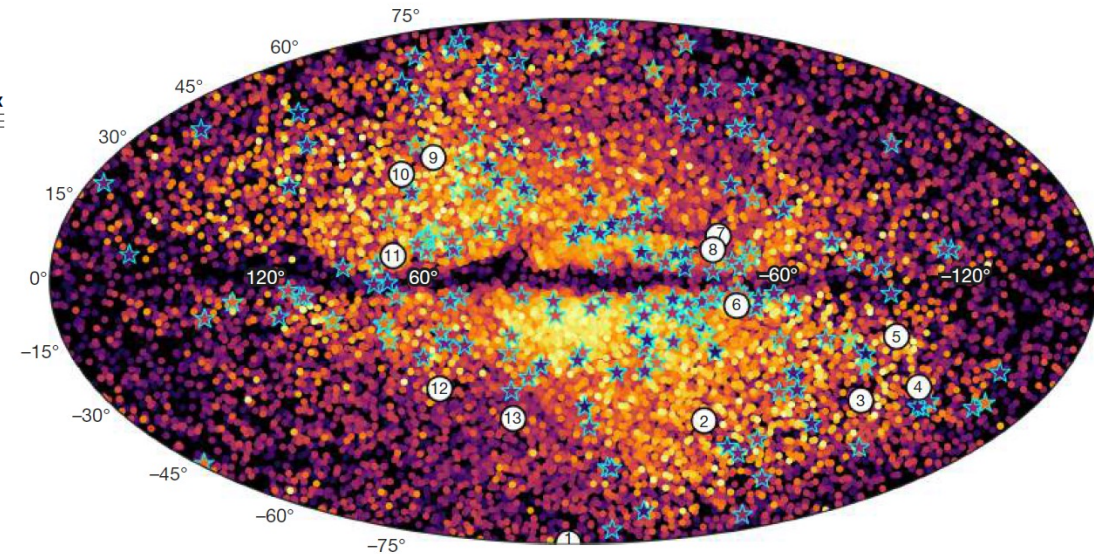
- Accretion history from stellar kinematics

LETTER 2018

<https://doi.org/10.1038/s41586-018-0625-x>

## The merger that led to the formation of the Milky Way's inner stellar halo and thick disk

Amina Helmi<sup>1\*</sup>, Carine Babusiaux<sup>2,3</sup>, Helmer H. Koppelman<sup>1</sup>, Davide Massari<sup>1</sup>, Jovan Veljanoski<sup>1</sup> & Anthony G. A. Brown<sup>4</sup>

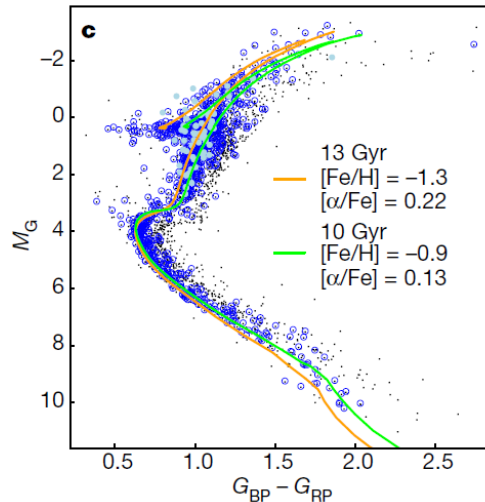


### Galactic Archaeology

*Combining Gaia data and other surveys it has been possible to study the motion, chemical composition, age, and spatial distribution of stars in the Galaxy inner halo.*

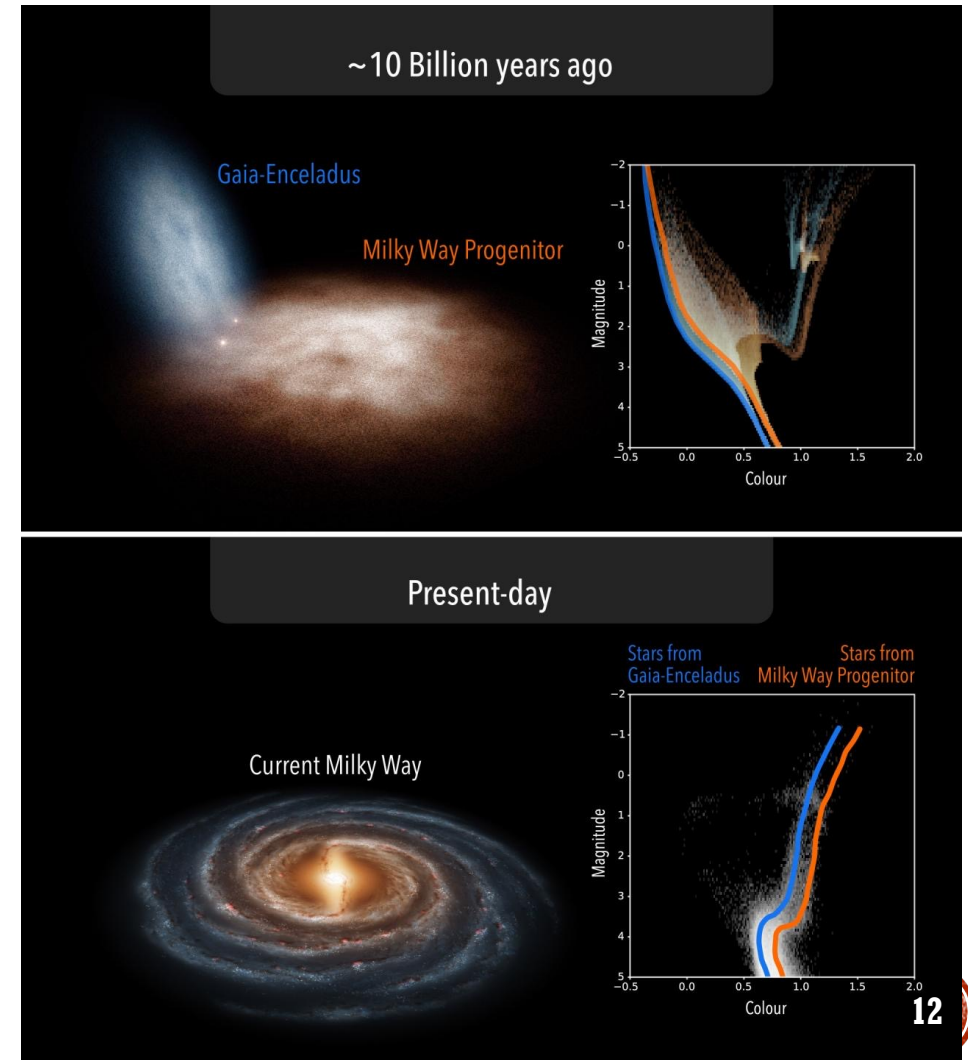
# THE MILKY WAY FORMATION HISTORY

- Accretion history from stellar kinematics



## Our galaxy merged Gaia-Enceladus

*The inner halo seems to be dominated by stars belonging to another galaxy than the Milky Way. The current hypothesis is that both galaxies merged 10 Gyr ago. The mass of Gaia-Enceladus was  $\frac{1}{4}$  of that of the MW*



# THE MILKY WAY FORMATION HISTORY

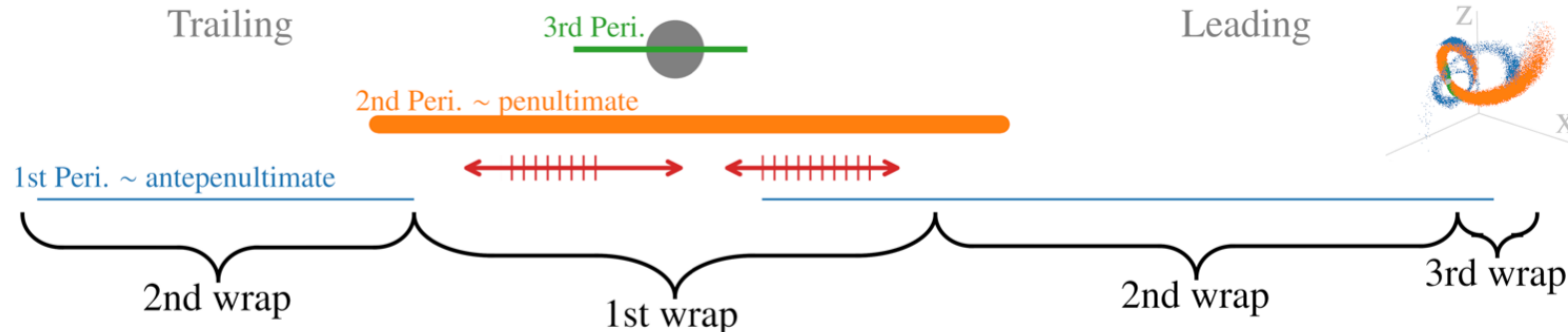
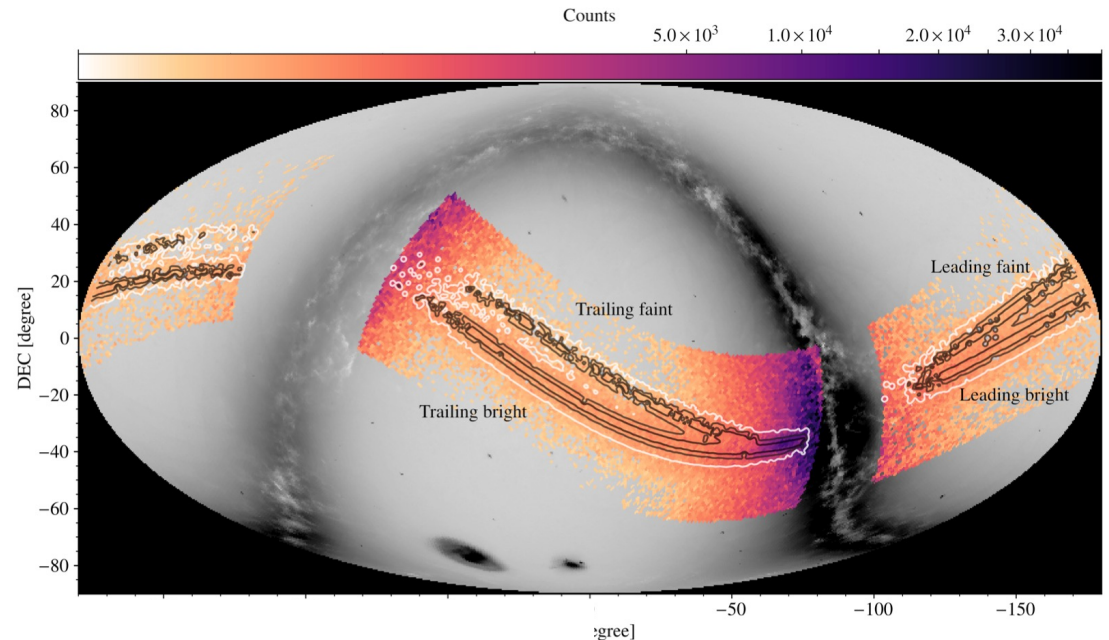
- Sagittarius stellar stream (among many others recently discovered)

## 2021 The Sagittarius stream in Gaia eDR3 and the origin of the bifurcations

P. Ramos<sup>1,\*</sup>, T. Antoja<sup>2</sup>, Z. Yuan<sup>1</sup>, A. Arentsen<sup>1</sup>, P.-A. Oria<sup>1</sup>, B. Famaey<sup>1</sup>, R. Ibata<sup>1</sup>, C. Mateu<sup>3</sup>, and J.A. Carballo-Bello<sup>4</sup>

### The Sagittarius orbit

*Using Gaia EDR3 data researchers have been able to find the stellar streams in front and behind the Sagittarius satellite's core. Modelling the formation of the stream we can now better constrain the satellite's orbit*





# THE MILKY WAY FORMATION HISTORY

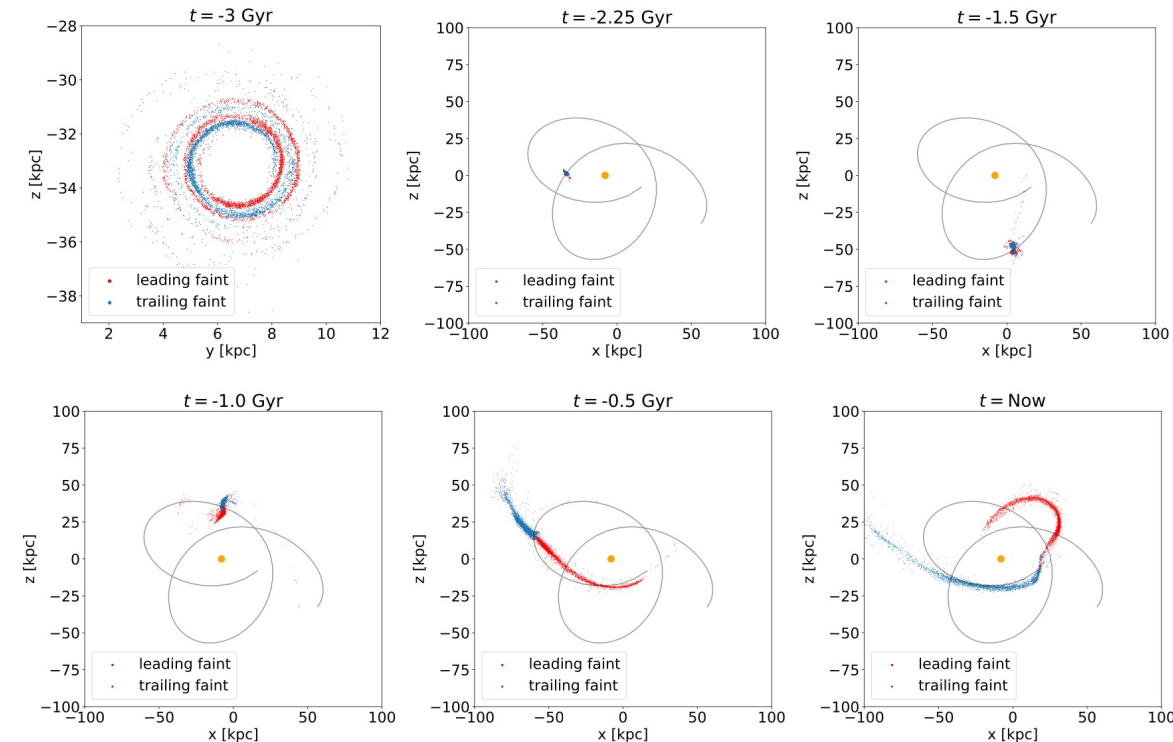
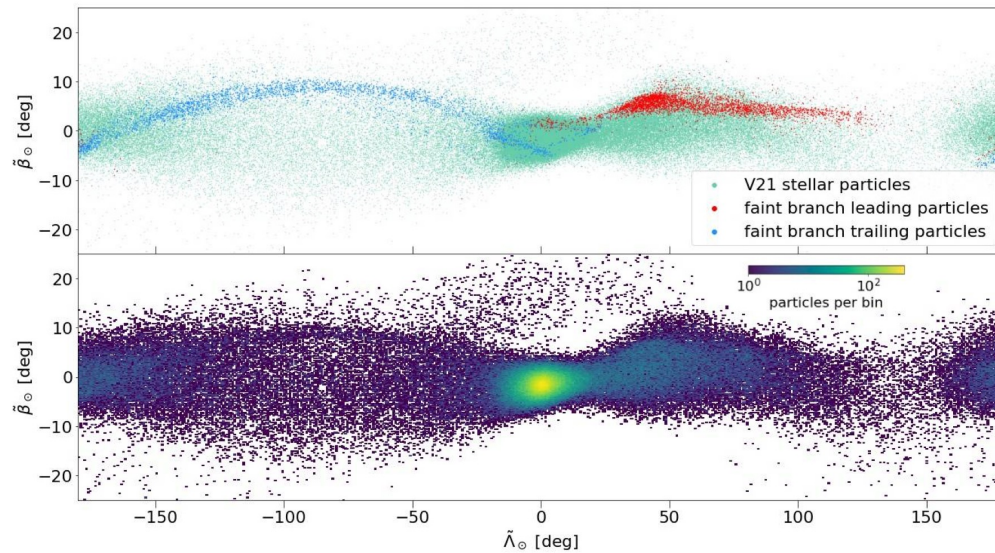
- Sagittarius stellar stream (among many others recently discovered)

2022

Revisiting a diskly origin for the faint branch of the Sagittarius stellar stream

PIERRE-ANTOINE ORIA,<sup>1</sup> RODRIGO IBATA,<sup>1</sup> PAU RAMOS,<sup>1</sup> BENOIT FAMAHEY ,<sup>1</sup> AND RAPHAËL ERRANI<sup>1</sup>

<sup>1</sup> *Université de Strasbourg, CNRS, Observatoire astronomique de Strasbourg, UMR 7550, F-67000 Strasbourg, France*

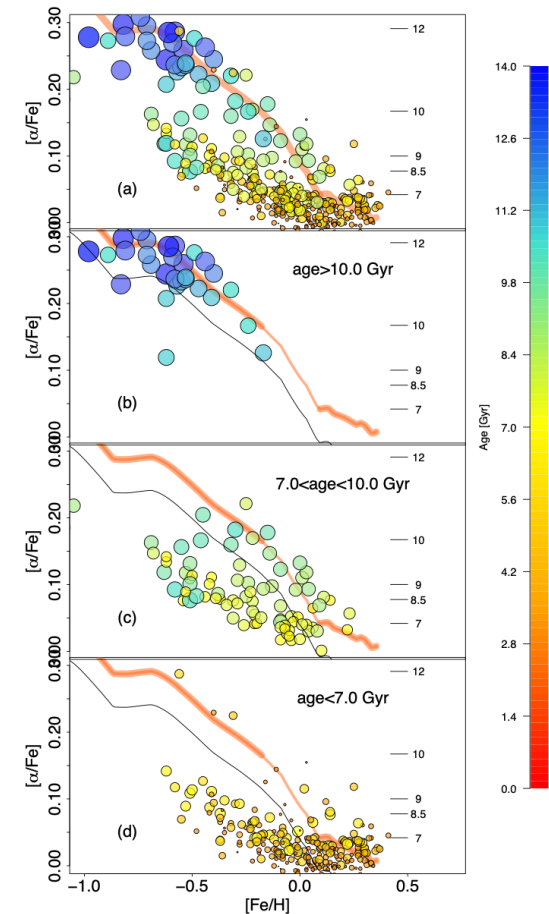
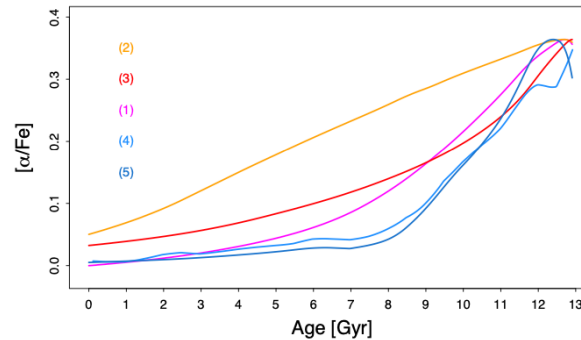
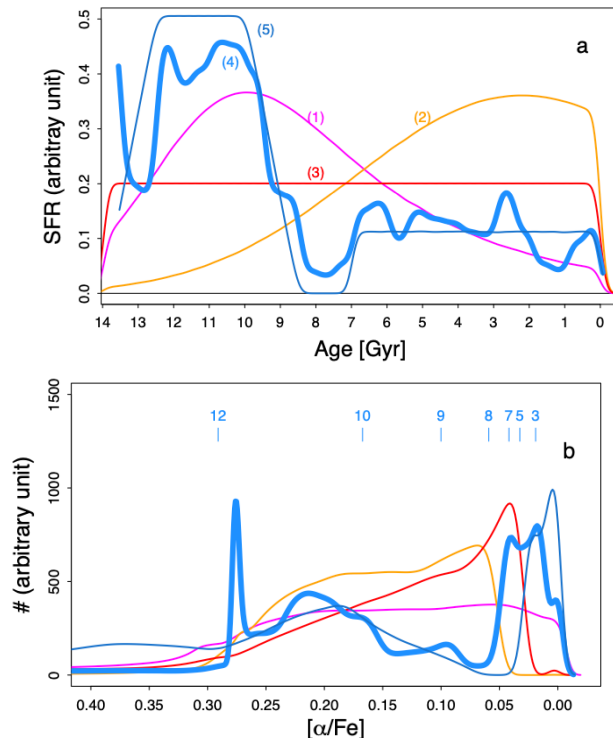


# THE MILKY WAY FORMATION HISTORY

2016 ■ Star Formation history from disk chemistry

When the Milky Way turned off the lights: APOGEE provides evidence of star formation quenching in our Galaxy

M. Haywood<sup>1</sup>, M. D. Lehnert<sup>2</sup>, P. Di Matteo<sup>1</sup>, O. Snaith<sup>3</sup>, M. Schultheis<sup>4</sup>, D. Katz<sup>1</sup>, A. Gómez<sup>1</sup>



# THE MILKY WAY FORMATION HISTORY

2019 ■ Star Formation History from CMD diagrams

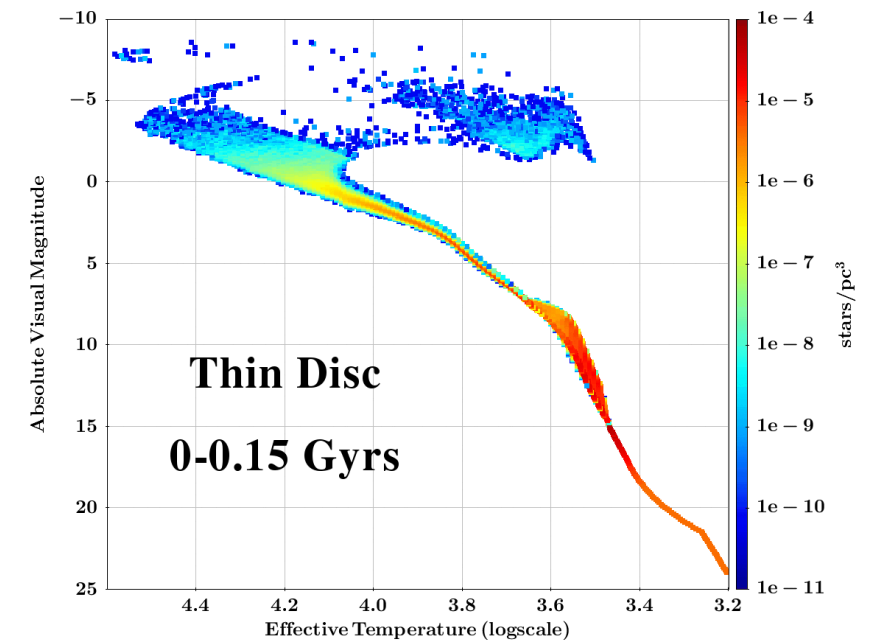
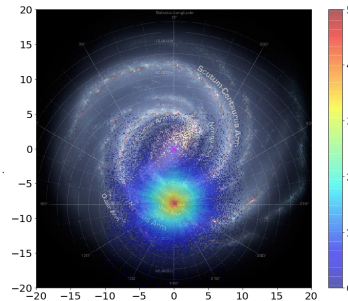
NATURE RESEARCH HIGHLIGHTS



## The cosmic drama that helped to build the Milky Way

Stellar baby boom added a slew of stars to the Galaxy's disk

R. Mor, A. Robin, F. Figueras, S. Roca-Fàbrega, X. Luri



## Star formation history of the galactic disk

*Combining distances, magnitudes, and colors of all observed stars up to a defined magnitude we have been able to trace the possible star formation history of the galactic disk's stars.*

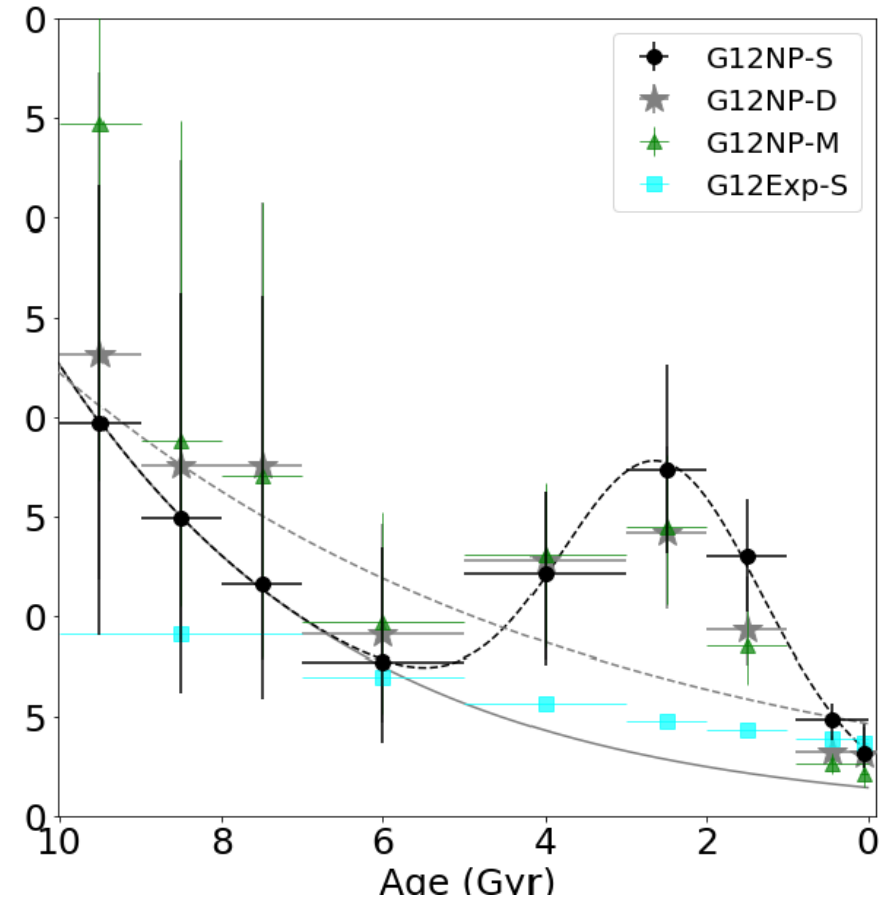


# THE MILKY WAY FORMATION HISTORY

- Star Formation History from CMD diagrams

## A strong star formation burst 2-3 Gyrs

The impact of a small satellite with the Galactic disk was the most probable origin of the star formation burst that produced more than the 50% of stars in the current MW galactic disk. It was the origin of the thin disk.



# THE MILKY WAY FORMATION HISTORY

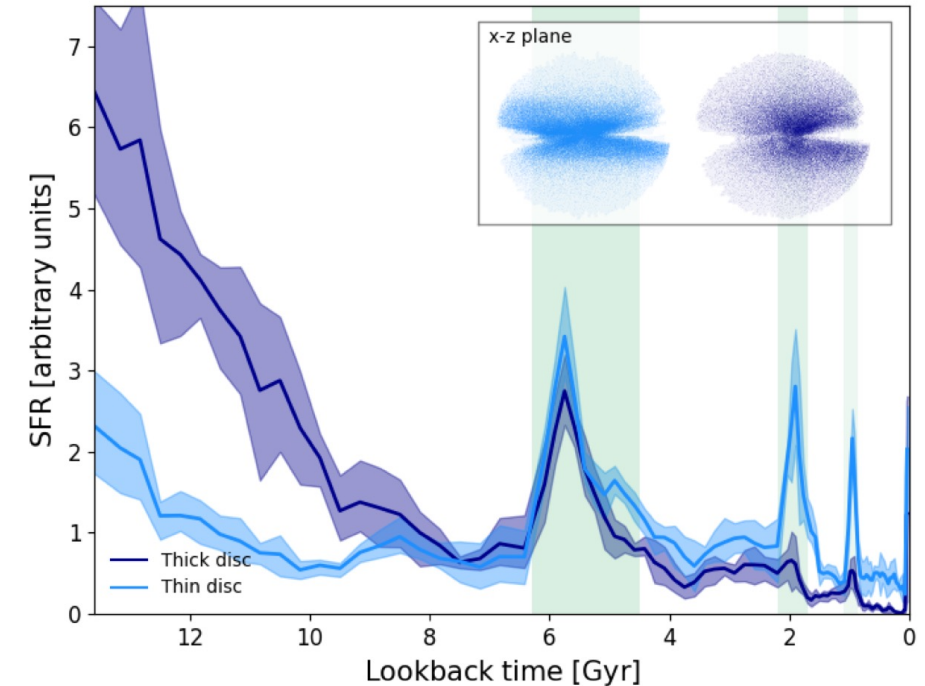
2020 ■ Star Formation History from CMD diagrams

## The recurrent impact of the Sagittarius dwarf galaxy on the star formation history of the Milky Way disc

Tomás Ruiz-Lara<sup>1,2,\*</sup>, Carme Gallart<sup>1,2</sup>, Edouard J. Bernard<sup>3</sup>, and Santi Cassisi<sup>4,5</sup>

### Three star formation bursts in the galactic thin disk well correlated with Sagittarius dwarf pericenters

The recent determination of the Sagittarius dwarf galaxy recent pericenters allowed researchers to set a correlation between those and star formation bursts derived from CMD diagrams. These works confirmed that the Sagittarius dwarf strongly perturbed the disk at least three times in the last 6 Gyrs.



# THE MILKY WAY FORMATION HISTORY

- Gas inflow history using Galaxy Chemistry Evolution models: Lithium

## The bi-modal ${}^7\text{Li}$ distribution of the Milky Way's thin-disk dwarf stars

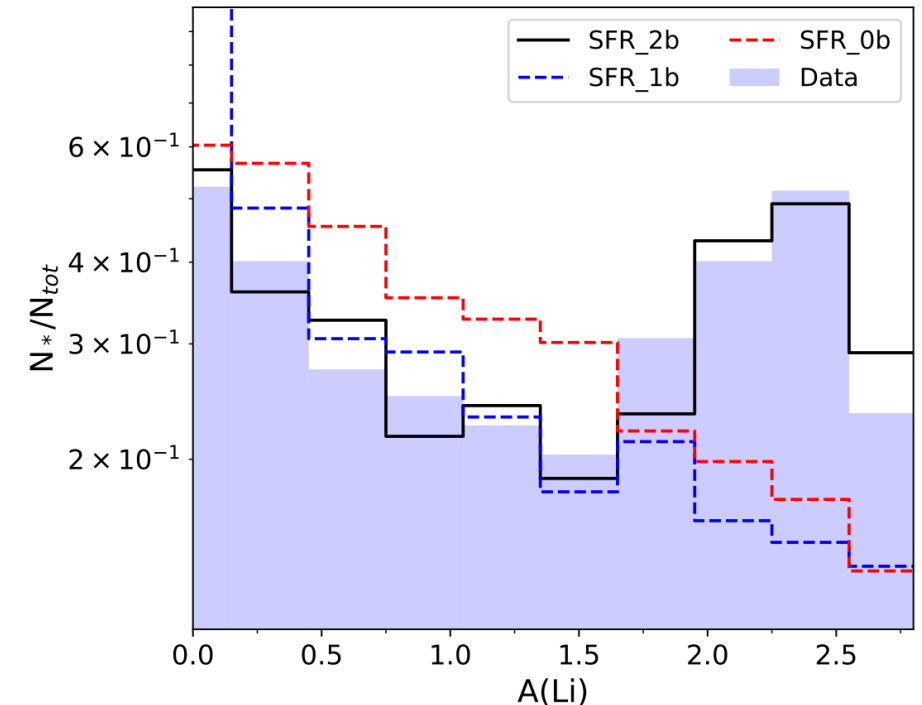
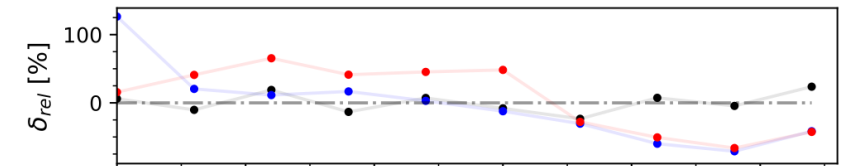
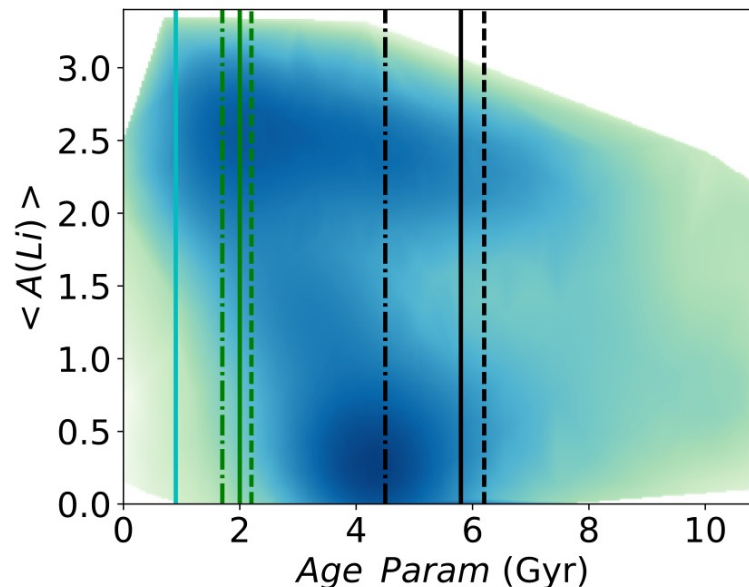
2021

### The role of Galactic-scale events and stellar evolution

S. Roca-Fàbrega<sup>1,\*</sup>, F. Llorente de Andrés<sup>2,3</sup>, C. Chavero<sup>4</sup>, C. Cifuentes<sup>3</sup> and R. de la Reza<sup>5</sup>

**Lithium is atypical, it is not produced in stellar cores**

Lithium is an element that is produced and destroyed in the external layers of stars, or in the intergalactic medium. Also, is one of the only elements that were produced in the primordial nucleosynthesis. This makes it a perfect tracer of gas inflow history.





# THE MILKY WAY FORMATION HISTORY

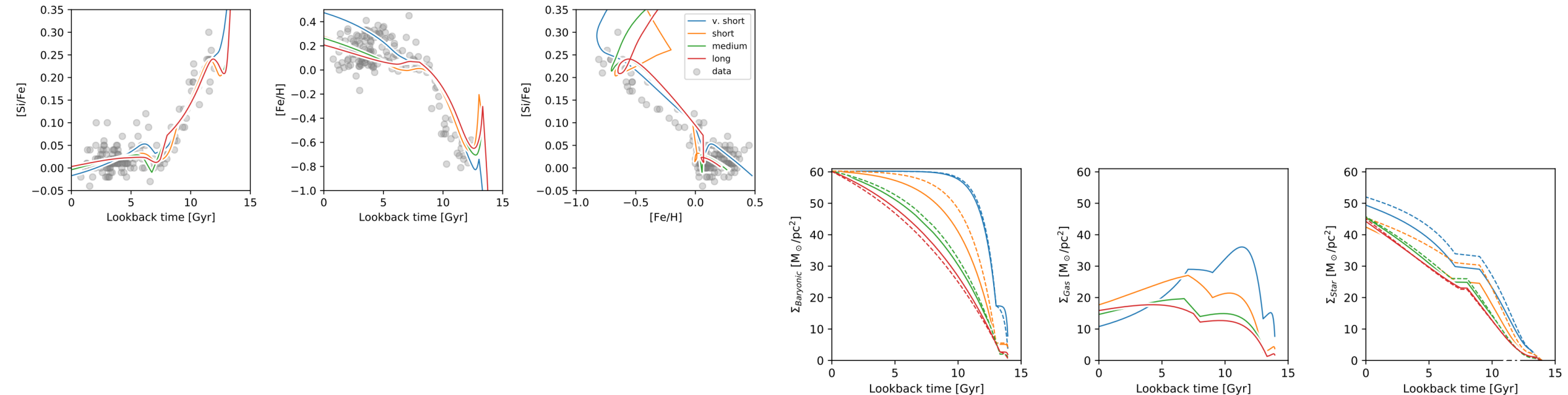
- Gas inflow history using Galaxy Chemistry Evolution models: Other tracers

## Rapid early gas accretion for the inner Galactic disc

2022

A case for a short accretion timescale

Owain Snaith<sup>1</sup>, Misha Haywood<sup>1,2</sup>, Paola Di Matteo<sup>1,2</sup>, Matthew Lehnert<sup>2\*</sup>, David Katz<sup>1</sup>, and Sergey Khoperskov<sup>3,4,1</sup>



# THE MILKY WAY FORMATION HISTORY

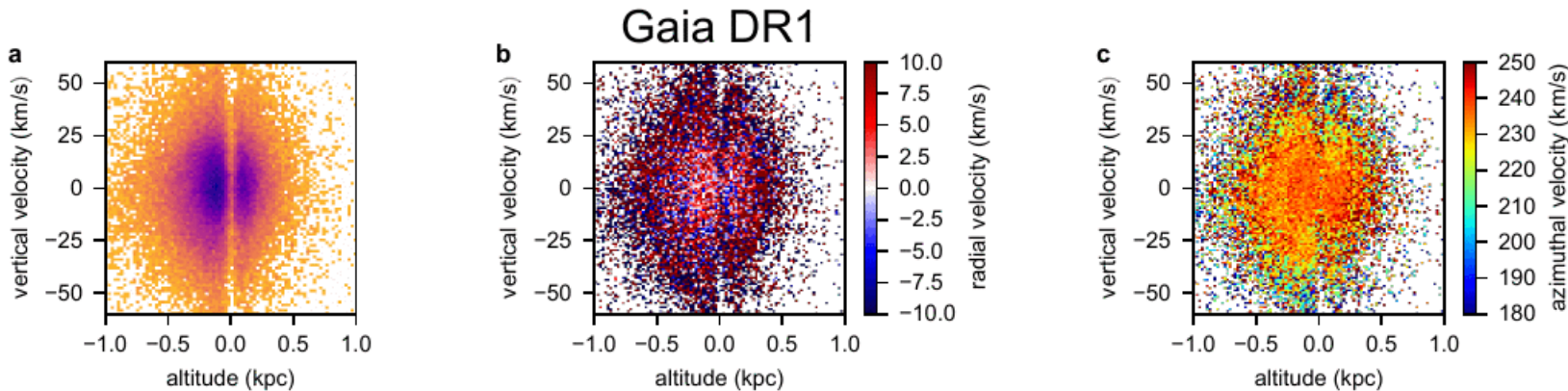
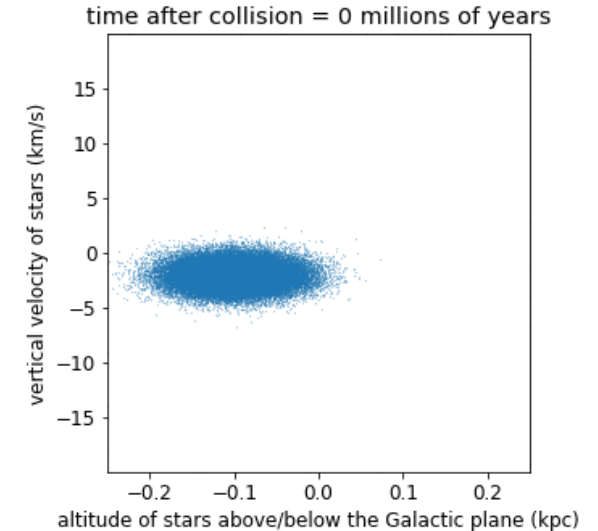
- A highly perturbed galactic disk (Gaia kinematic data)

LETTER 2018

<https://doi.org/10.1038/s41586-018-0510-7>

## A dynamically young and perturbed Milky Way disk

T. Antoja<sup>1\*</sup>, A. Helmi<sup>2</sup>, M. Romero-Gómez<sup>1</sup>, D. Katz<sup>3</sup>, C. Babusiaux<sup>3,4</sup>, R. Drimmel<sup>5</sup>, D. W. Evans<sup>6</sup>, F. Figueras<sup>1</sup>, E. Poggio<sup>5,7</sup>, C. Reylé<sup>8</sup>, A. C. Robin<sup>8</sup>, G. Seabroke<sup>9</sup> & C. Soubiran<sup>10</sup>

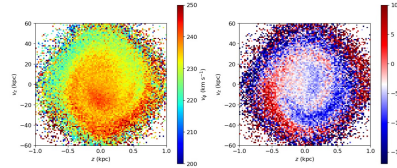


A possible explanation: a small galaxy (Sagittarius) caused 'snail shell' structure in our part of the MW. The spiral allows to predict the pericentric passage (300-900 Myr ago)

# THE MILKY WAY FORMATION HISTORY

- A highly perturbed galactic disk (Gaia kinematic data)

Confirmation  
by isolated  
galaxy  
simulations

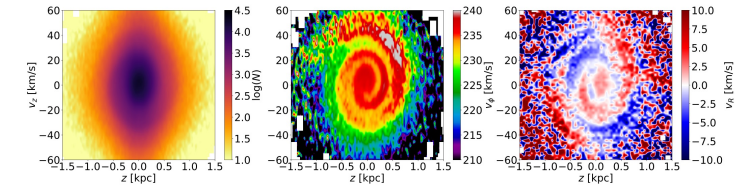


Galactic seismology: the evolving “phase spiral” after the  
Sagittarius dwarf impact 2020

Joss Bland-Hawthorn<sup>1,2\*</sup> and Thor Tepper-García<sup>1,2,3</sup>

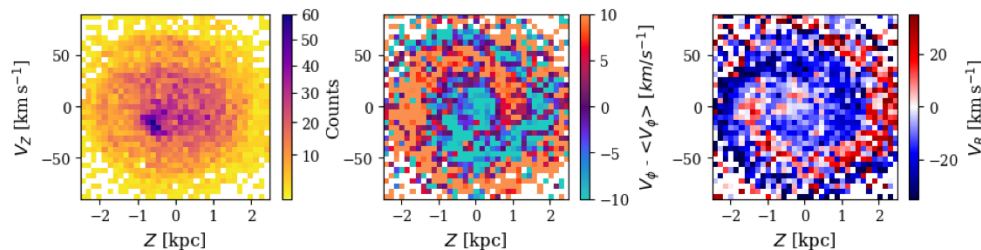
Footprints of the Sagittarius dwarf galaxy in the *Gaia* data  
set 2018

Chervin F. P. Laporte,<sup>1\*†</sup> Ivan Minchev,<sup>2</sup> Kathryn V. Johnston,<sup>3</sup> Facundo A. Gómez<sup>4,5</sup>

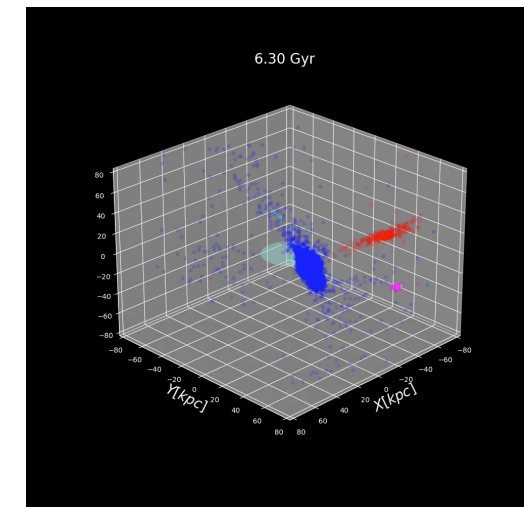


Phase spirals in cosmological simulations of Milky Way-size  
galaxies 2022

B. García-Conde,<sup>1\*</sup> S. Roca-Fàbrega,<sup>1</sup> T. Antoja<sup>2</sup>, P. Ramos<sup>3</sup>, and O. Valenzuela<sup>4</sup>



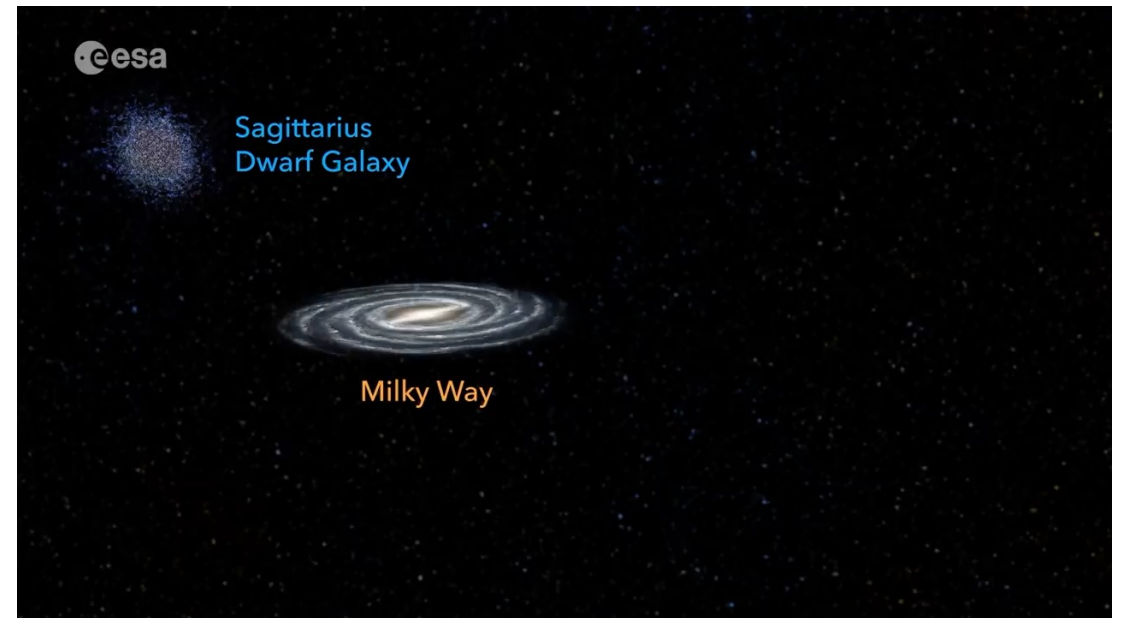
But a galaxy in a  
cosmological  
context is far more  
complex ...





- The formation and evolution of disk galaxies is complex, many physical processes need to be taken into account.
- With the recent release of the Gaia data and from the many active collaborations that studies stellar and gas physical properties we now know much better the recent past of our Galaxy.
- However, the Galaxy still keeps many secrets waiting to be unveiled by new generations of astronomers.

# Thank you!



*“The past gave us experience and made us wiser so that we can create a beautiful and brighter future.”*

— **Debasish Mridha**